





Dr. Philipp Lauber, Dr. G. Birkenmeier

- https://pwl.home.ipp.mpg.de/tum/2023\_WS.html
- please choose a topic from the list below according to your level: pro-seminar/first contact students have priority on introductory topics
- contact us (<u>philipp.lauber@ipp.mpg.de</u>, gregor.birkenmeier@ipp.mpg.de) <u>by Oct 31st</u> via email with your preferred topic and one alternative (first come first serve basis);
- you are encouraged to choose your own topic (e.g. bachelors, masters, interns) related to your thesis/work

#### aim:

introductory/deeper knowledge about some aspects of plasma physics
train presentation skills:

- preparation of scientific material (reading/understanding)
- combine and present material in your own style/words/slides
- explain and 'teach' your material in class, answer questions

## style:

- duration 25+15 mins, language English
- discussion: all attendants should try to ask questions
- prepare slides (and/or blackboard)
- slides to be discussed and iterated with Gregor/Philipp before the presentation (a couple of days before)

#### **Basic Plasma Properties**



classification of plasmas, Debye theory

derivation, consequences, application to ionosphere



#### cross section, Coulomb logarithm, friction force

#### models for describing plasmas: fluid - kinetic

$$\frac{d\rho}{dt} + \rho\,\nabla\cdot V \quad = 0,$$

$$\rho \frac{d\mathbf{V}}{dt} + \nabla p - \frac{(\nabla \times \mathbf{B}) \times \mathbf{B}}{\mu_0} = \mathbf{0},$$

$$-\frac{\partial \mathbf{B}}{\partial t} + \nabla \times (\mathbf{V} \times \mathbf{B}) = \mathbf{0},$$

$$\frac{d}{dt}\left(\frac{p}{\rho^{\Gamma}}\right) = 0,$$

# Charged particle motion in inhomogeneous magnetic fields



drifts, guiding centre description (numerical approaches: implicit, explicit, symplectic)

## Confinement concepts





#### introductory Tokamaks (IET/AUG/IT60SA/ITER/DEMO)







Stellarators

#### introductory

basic concepts confinement experiments



## Runaway electrons



#### Kinetic Alfvén waves: theory, applications in space/astro/fusion



#### Energetic lons in Tokamaks

Alfvén waves, resonant interaction, non-linear saturation



## I.theoretical framework 2. experiment

#### introductory

#### Low temperature plasmas: principles and applications



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#### Plasma Thrusters



#### The Sun



nuclear processes, solar structure, solar equilibrium and equations of state, radiation transport, convection

#### The solar Co

#### vind





#### NASA, easa/SOHO

solar structure, magnetic fields in the sun(dynamo), solar spots, the corona heating problem

history, origin, Chapman model (static), Parker Model, interaction with earth magnetic field

#### **Electron-Positron Plasmas**



#### theoretical properties, experimental setup

## **Confinement regimes**





L-mode, H-mode, I-phase, I-mode, QH-mode, EDA-H-mode, Super-H-mode turbulence, magnetic confinement, transport barriers

#### H-mode ELMs





#### **Divertor detachment**





#### Plasma filaments in the scrape-off layer



PP





+ Plasmadiagnostiken...



#### 

#### **Plasma diagnostics**









C. Lechte et al., 2020 Plasma Sci. Technol. 22 064006

M. Griener et al., Review of Scientific Instruments 89, 10D102 (2018); https://doi.org/10.1063/1.5034446