

Dr. Philipp Lauber, Dr. G. Birkenmeier

- https://pwl.home.ipp.mpg.de/tum/2022_WS.html
- please choose a topic from the list below according to your level: pro-seminar/first contact students have priority on introductory topics
- contact us (philipp.lauber@ipp.mpg.de, gregor.birkenmeier@ipp.mpg.de) by Oct 31st via email with your preferred topic and one alternative (first come first serve basis);
- you are encouraged to choose your own topic (e.g. bachelors, masters, interns) related to your thesis/work

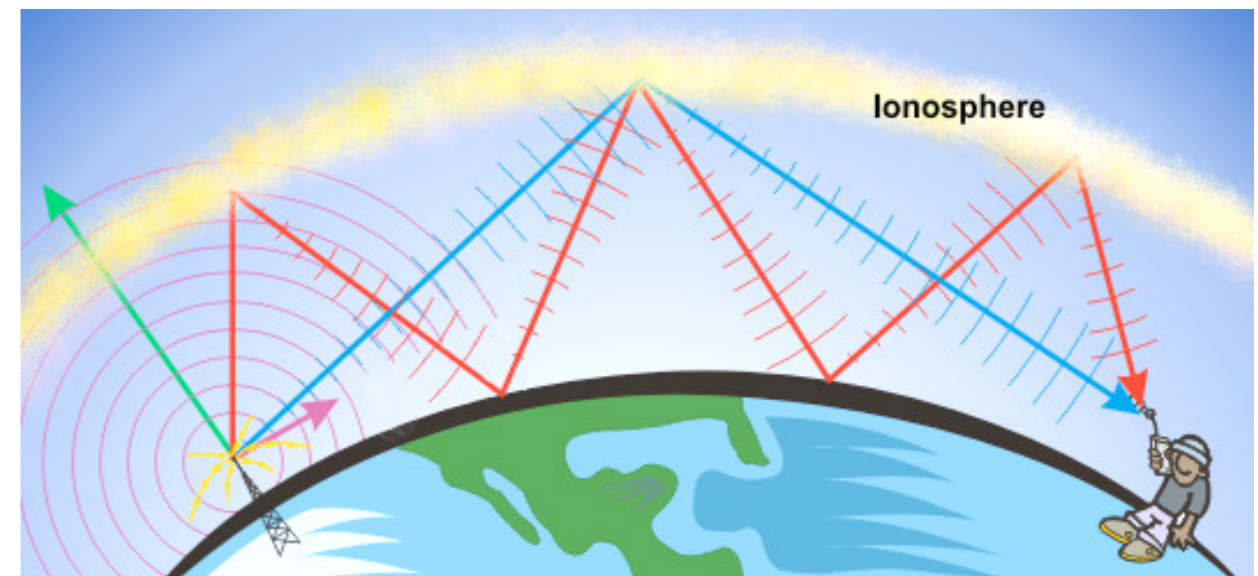
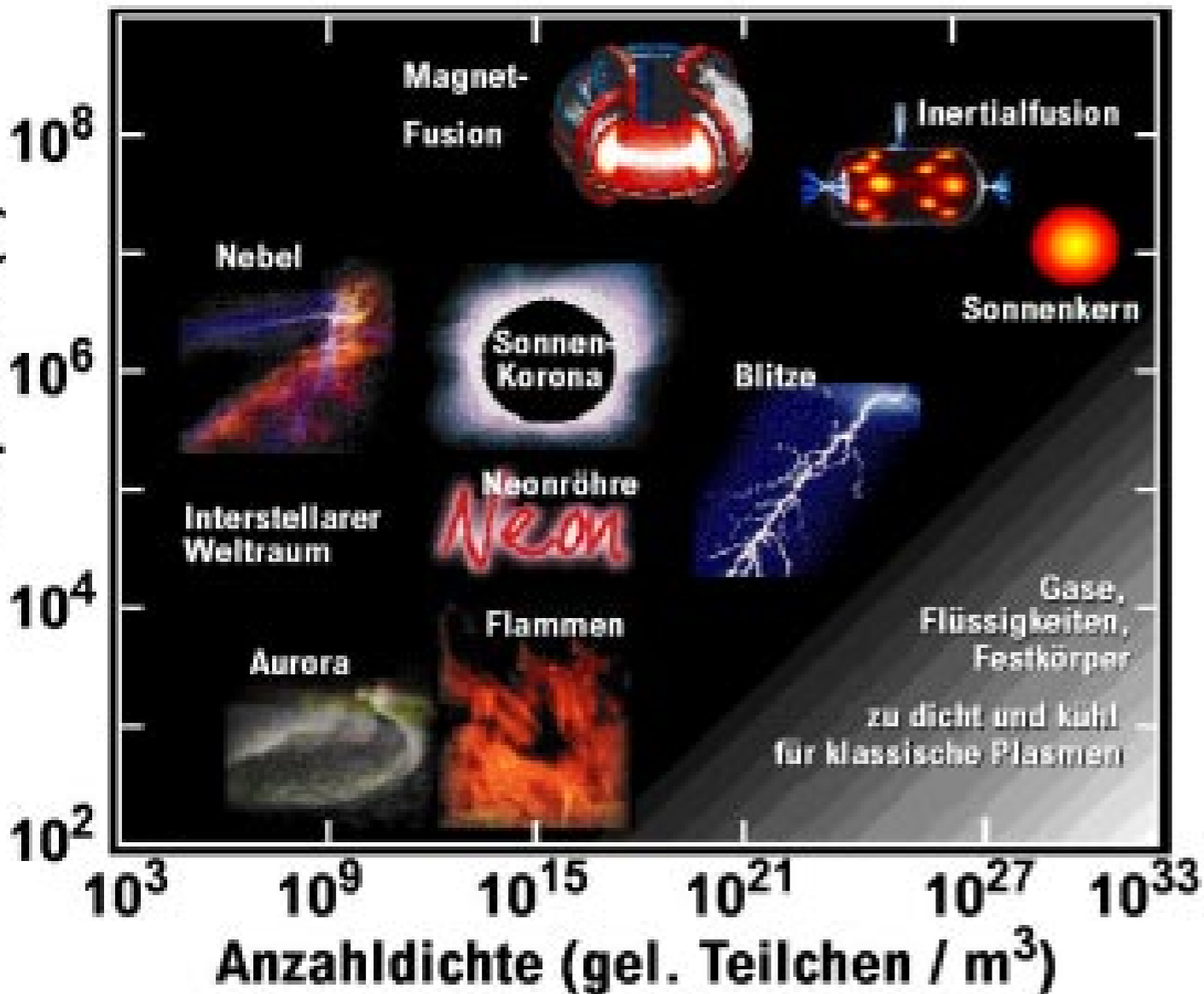
aim:

- introductory/deeper knowledge about some aspects of plasma physics
- train presentation skills:
 - preparation of scientific material (reading/understanding)
 - combine and present material in your own style/words/slides
 - explain and 'teach' your material in class, answer questions

style:

- duration 25+15 mins, language English
- discussion: all attendants should try to ask questions
- prepare slides (and/or blackboard)
- slides to be discussed and iterated with Gregor/Philipp before the presentation (a couple of days before)

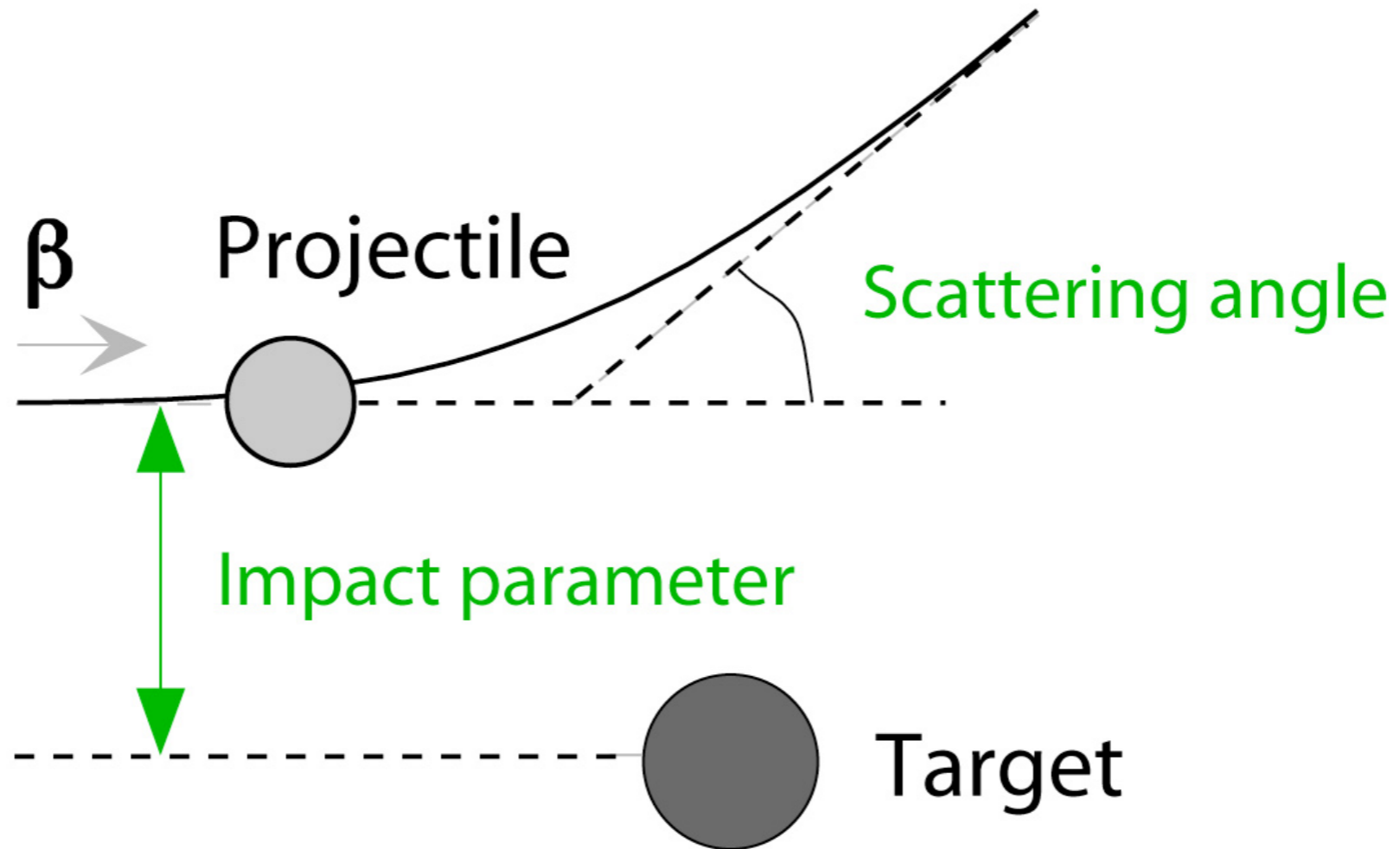
Basic Plasma Properties



classification of plasmas,
Debye theory

derivation, consequences,
application to ionosphere

Coulomb collisions



cross section, Coulomb logarithm, friction force

Magnetohydrodynamics (MHD)

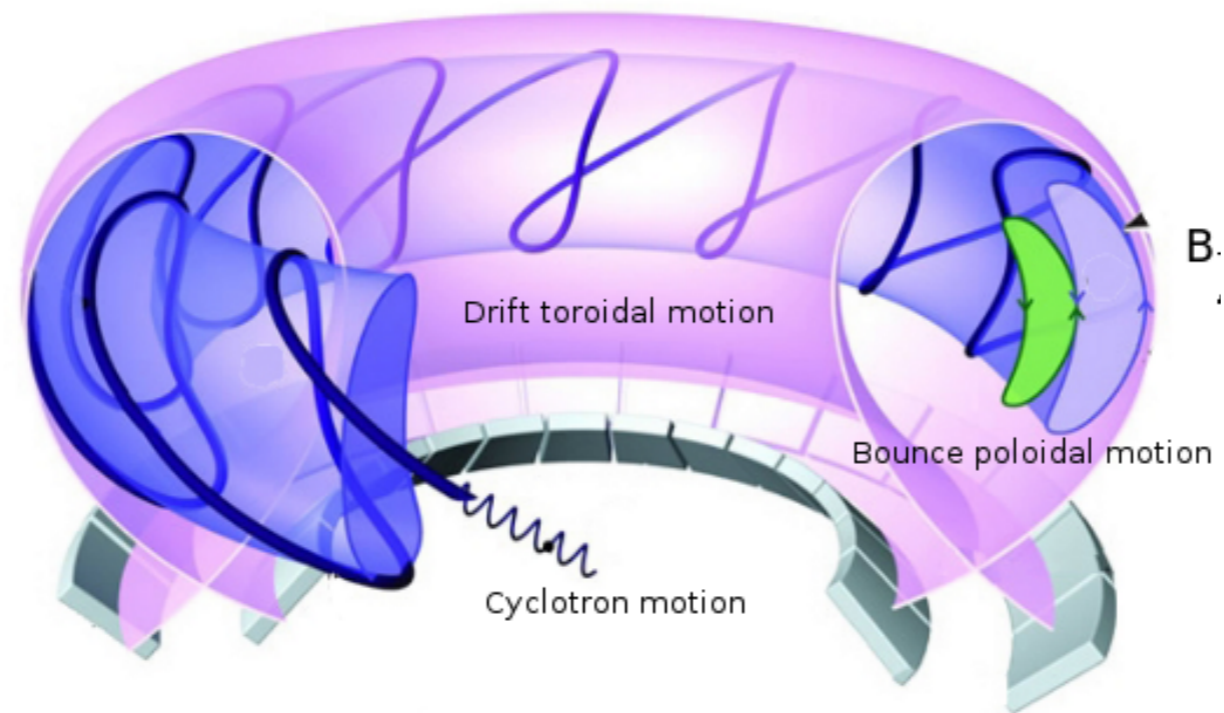
$$\frac{d\rho}{dt} + \rho \nabla \cdot \mathbf{V} = 0,$$

$$\rho \frac{d\mathbf{V}}{dt} + \nabla p - \frac{(\nabla \times \mathbf{B}) \times \mathbf{B}}{\mu_0} = 0,$$

$$-\frac{\partial \mathbf{B}}{\partial t} + \nabla \times (\mathbf{V} \times \mathbf{B}) = 0,$$

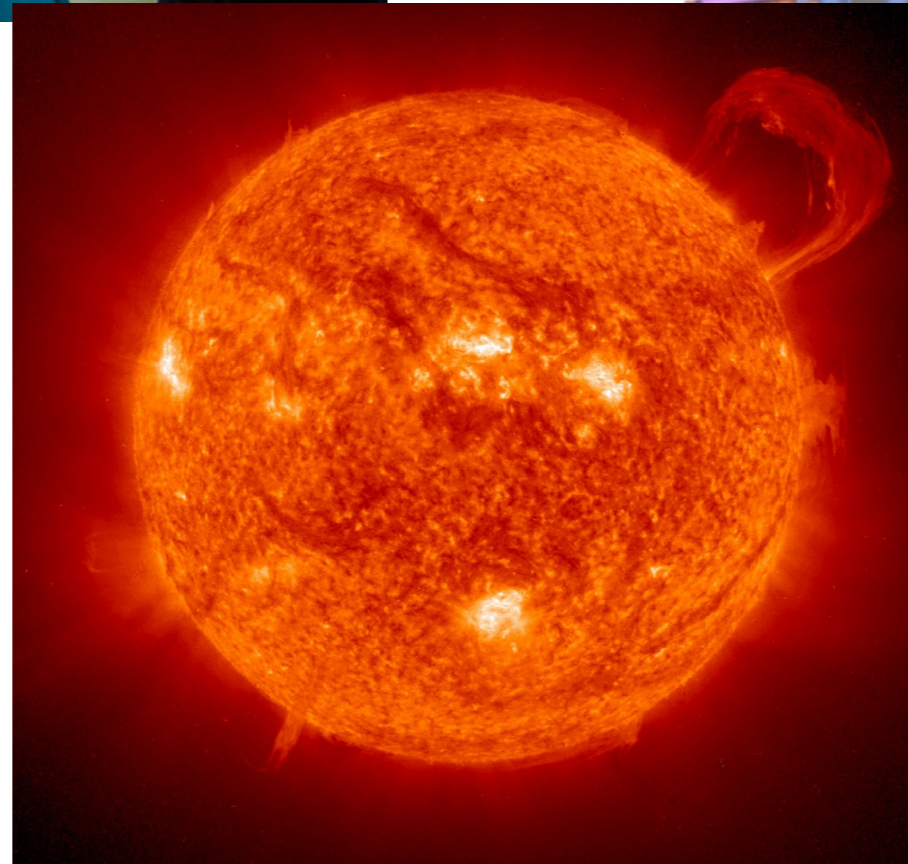
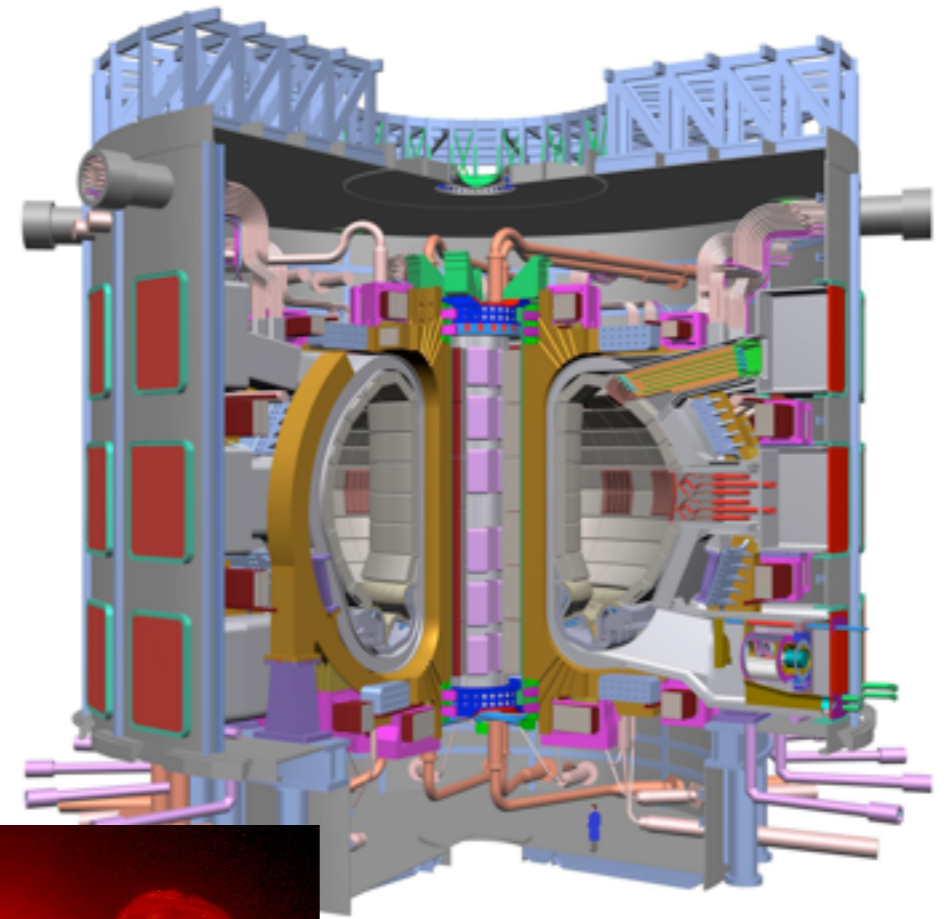
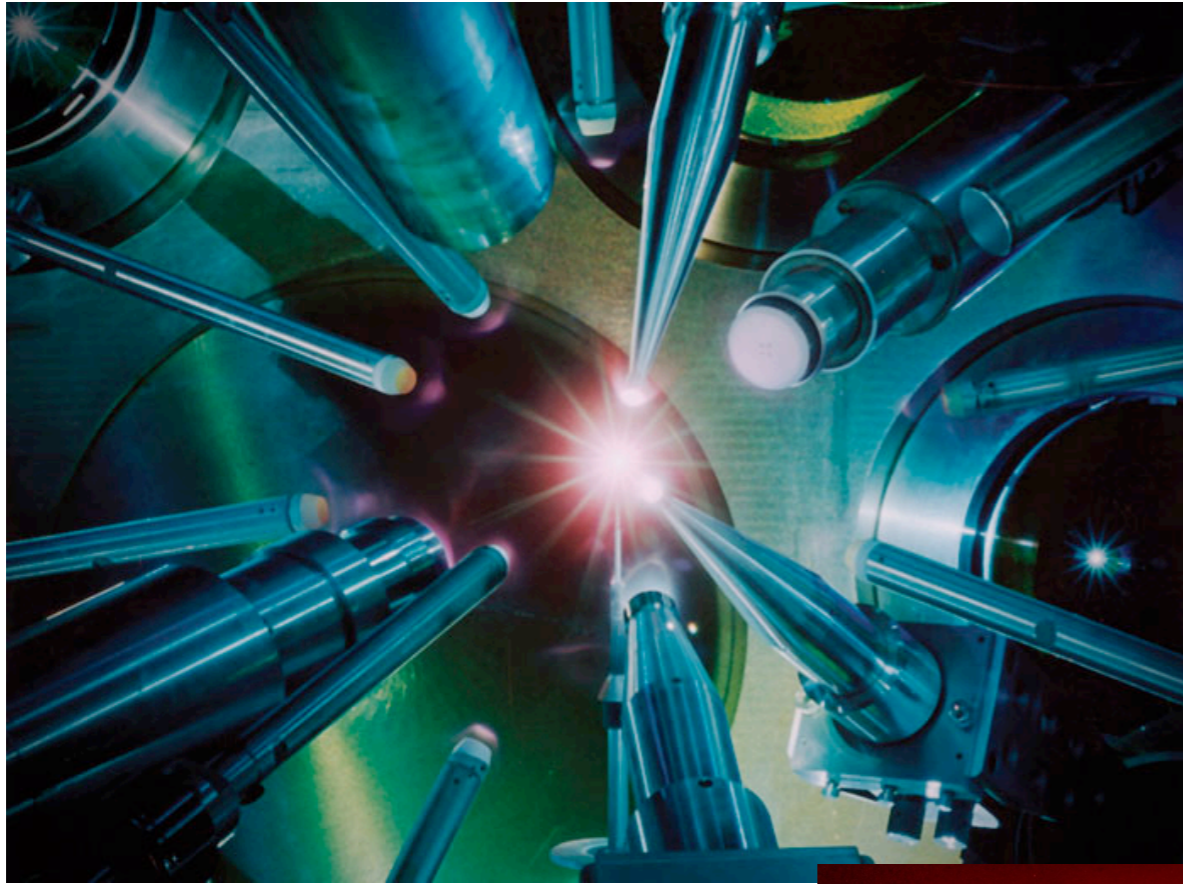
$$\frac{d}{dt} \left(\frac{p}{\rho^\Gamma} \right) = 0,$$

Charged particle motion in inhomogeneous magnetic fields

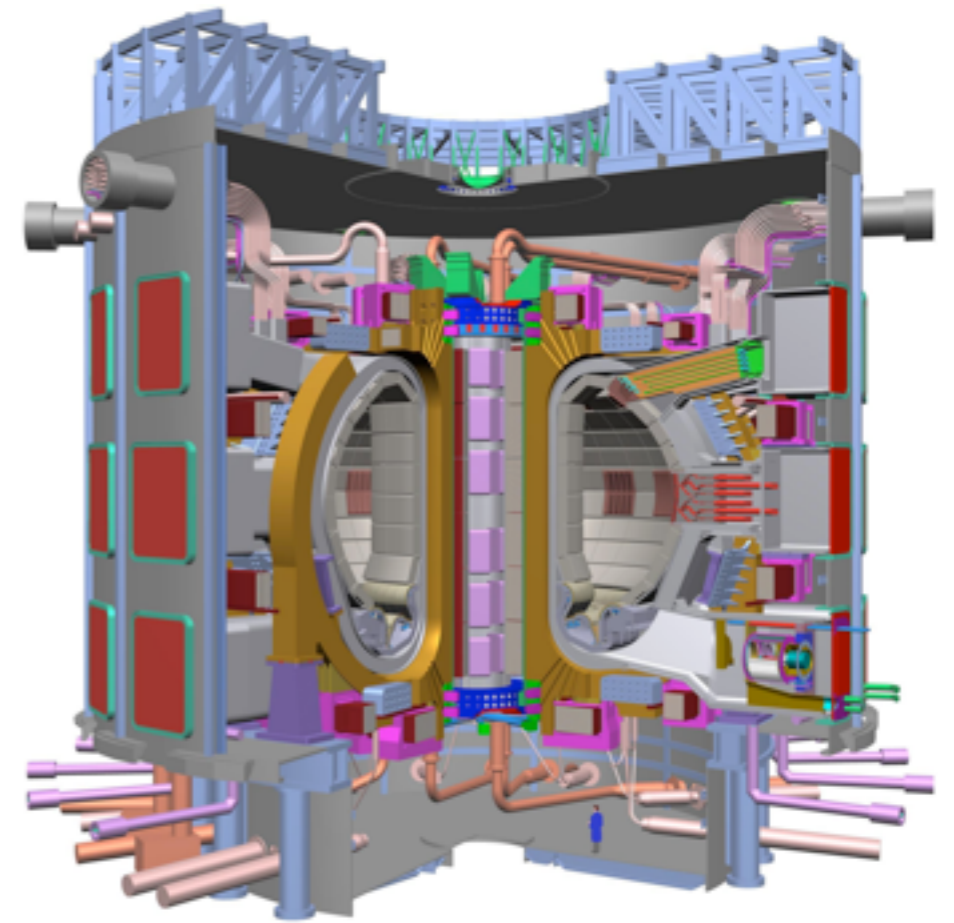


drifts, guiding centre description
(numerical approaches: implicit, explicit, symplectic)

Confinement concepts

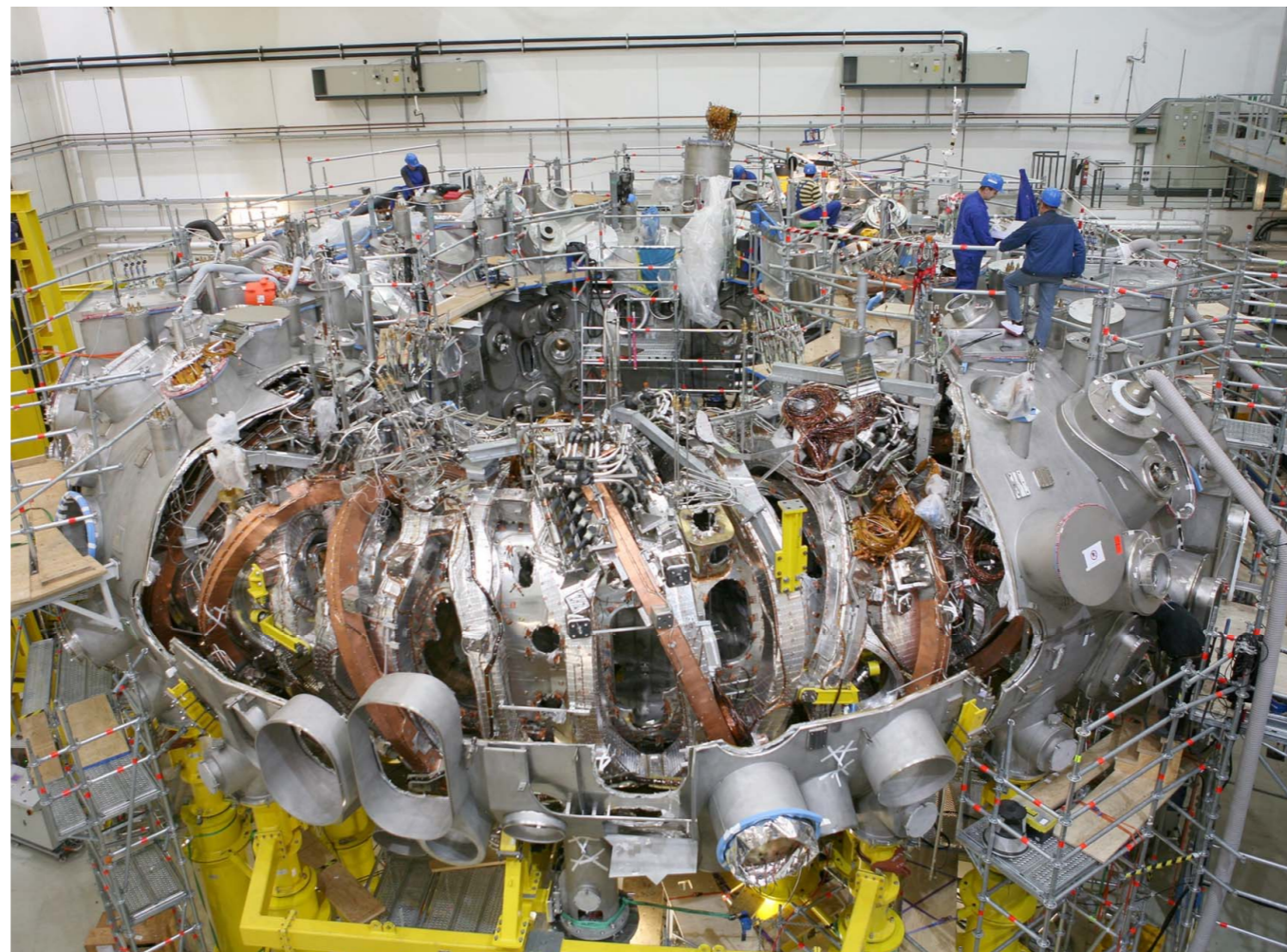


Tokamaks (JET/AUG/IT60SA/ITER/DEMO)

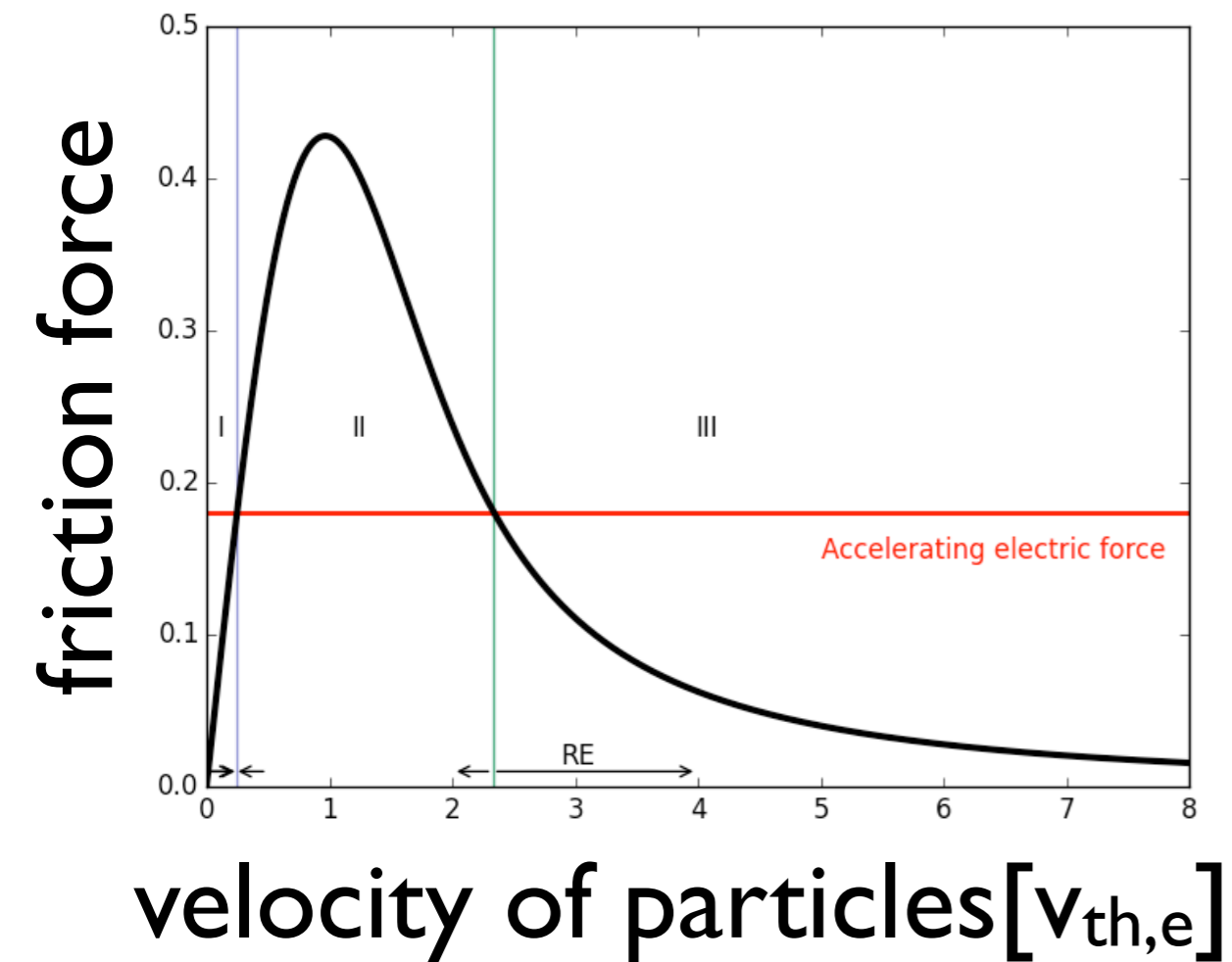
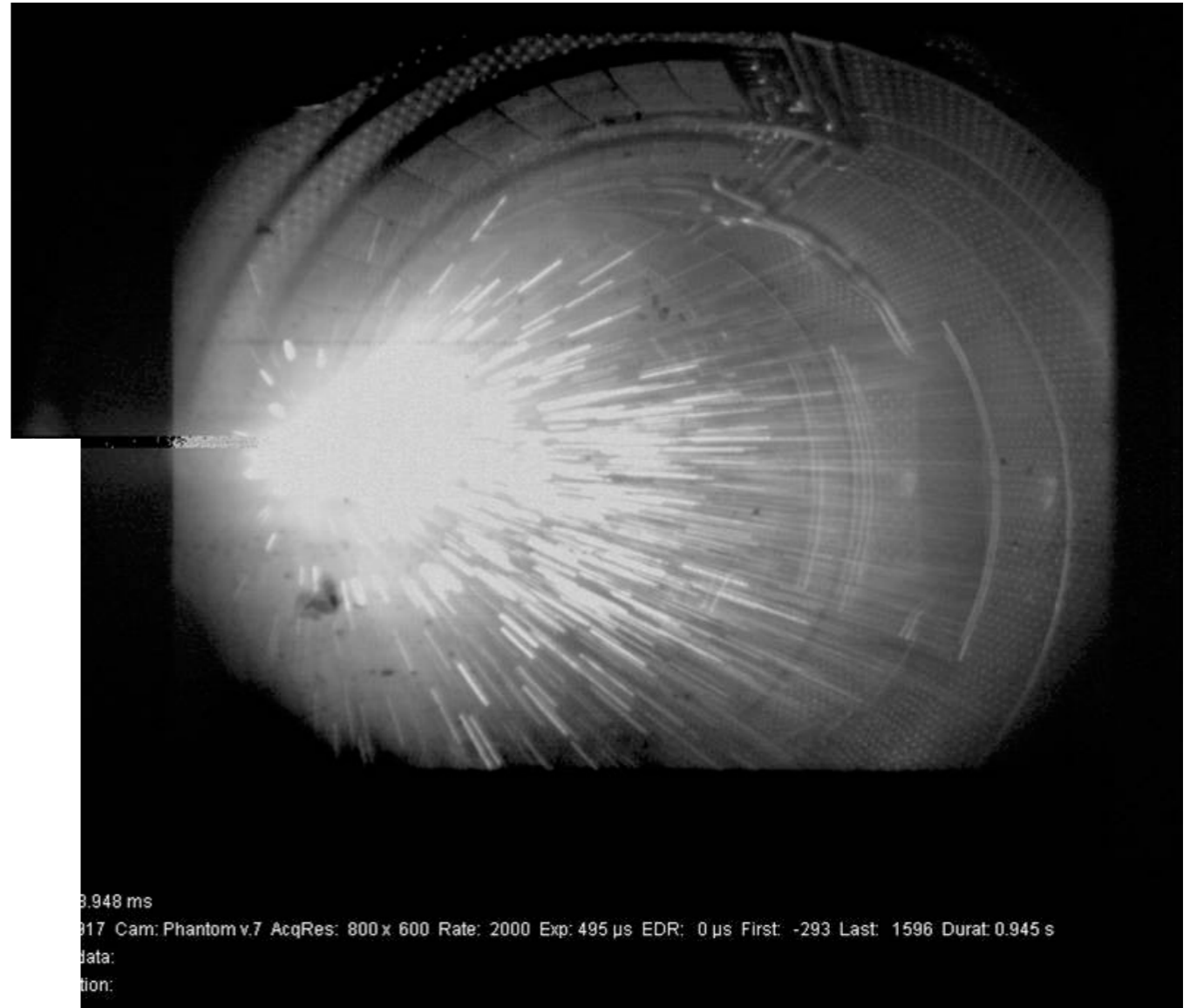


Stellarators

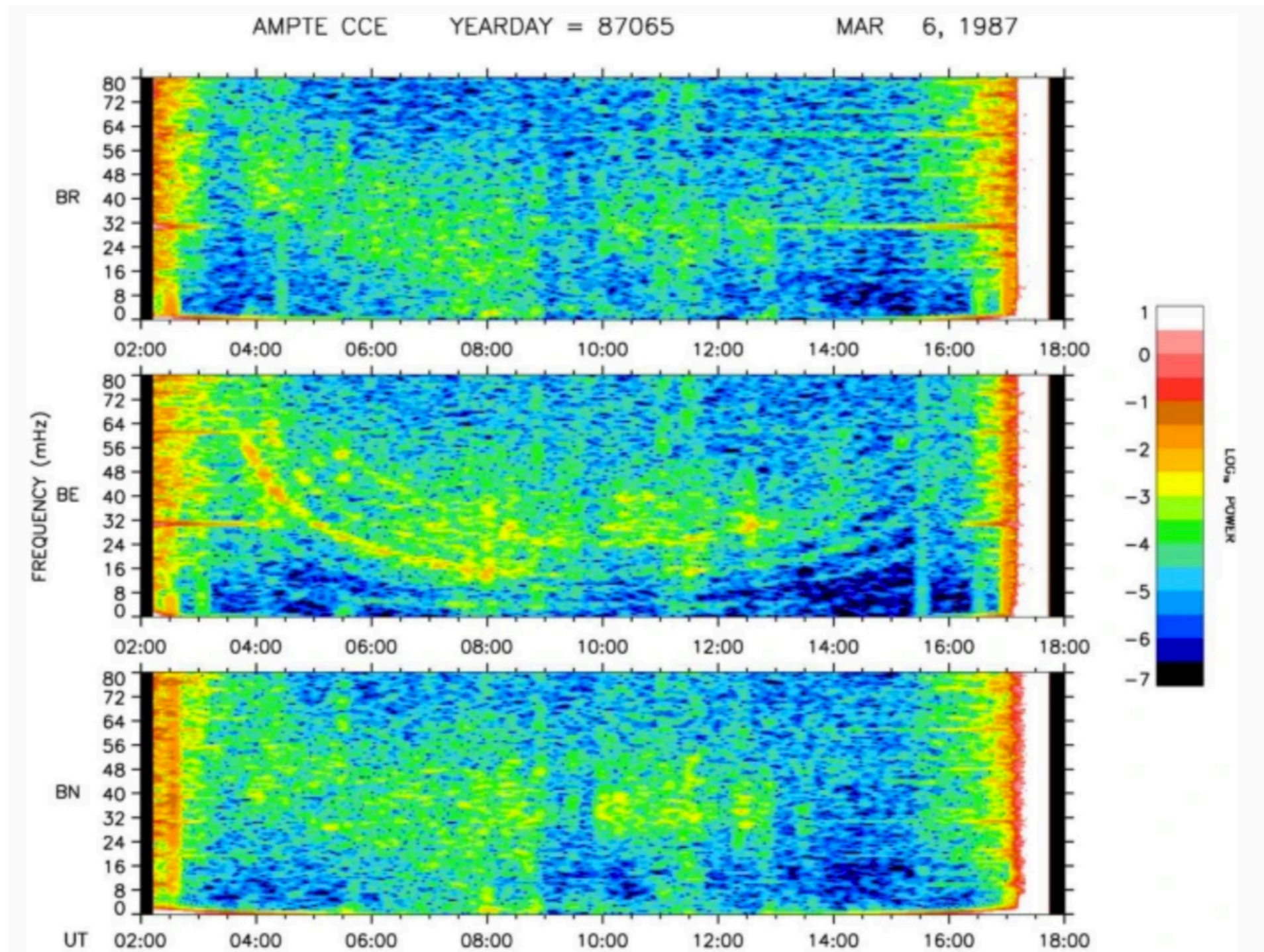
basic concepts
confinement
experiments



Runaway electrons

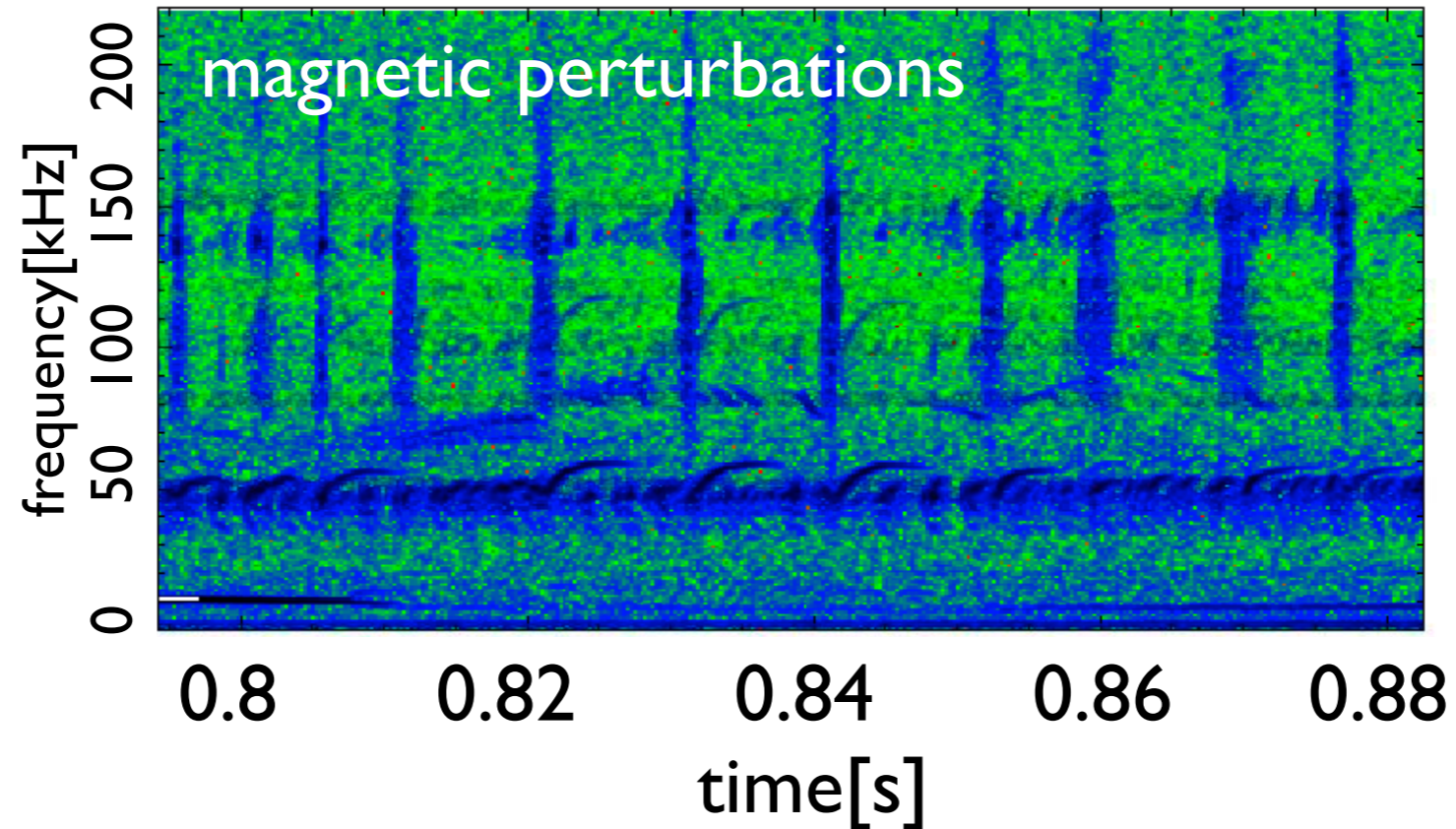


Kinetic Alfvén waves: theory, applications in space/astro/fusion



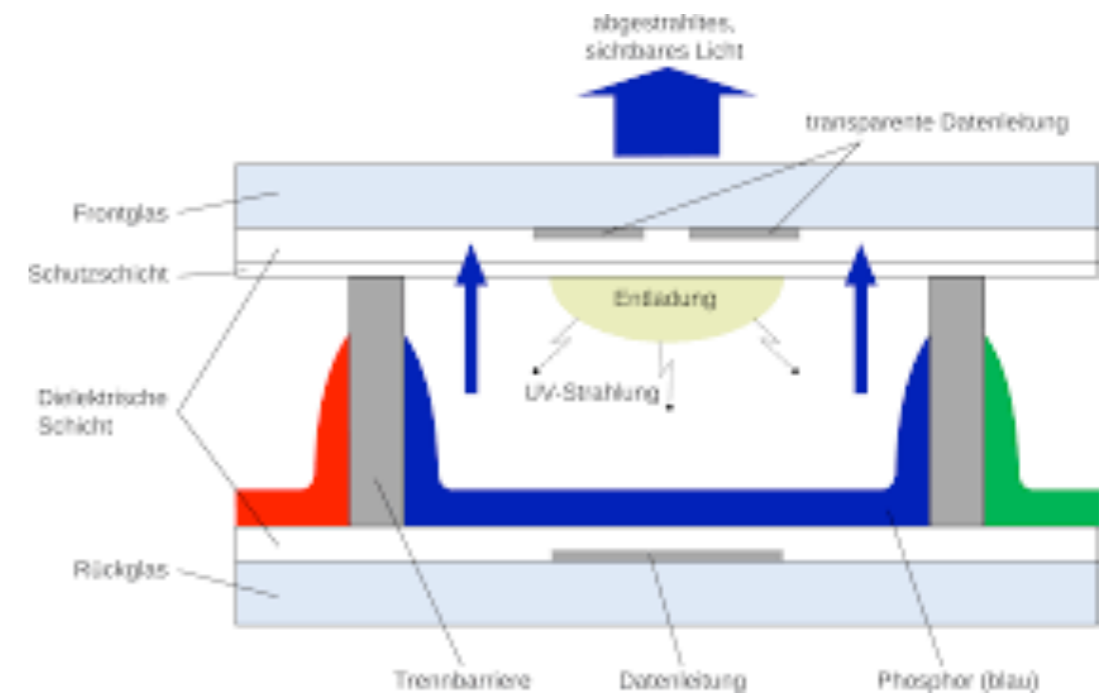
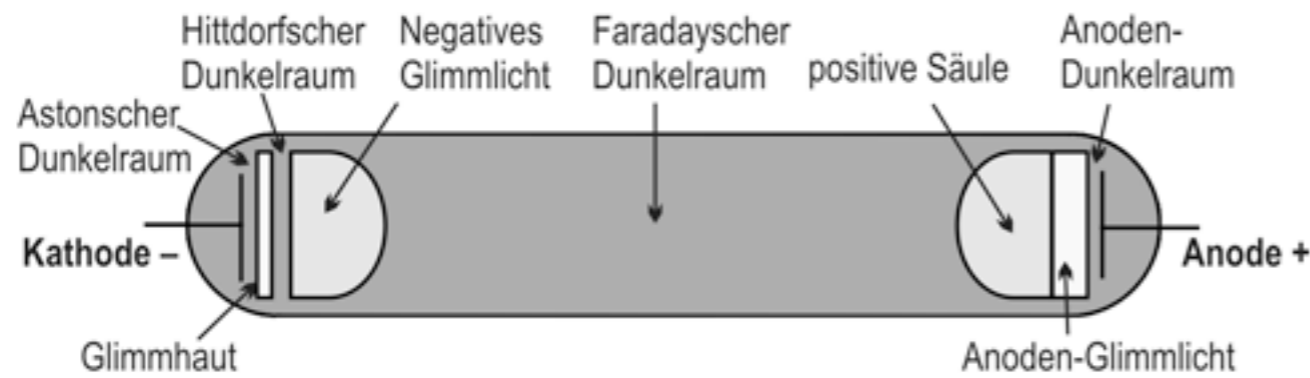
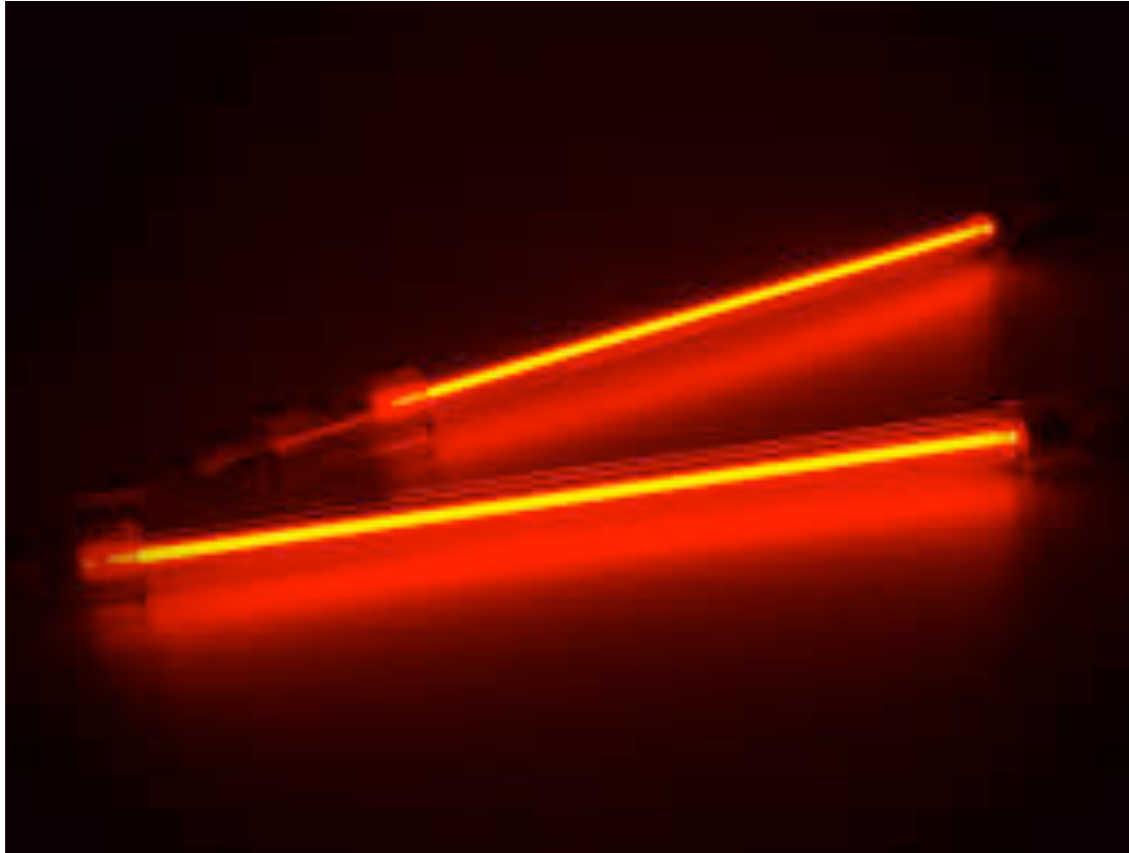
Energetic Ions in Tokamaks

Alfvén waves, resonant interaction, non-linear saturation



1. theoretical framework
2. experiment

Low temperature plasmas: principles and applications

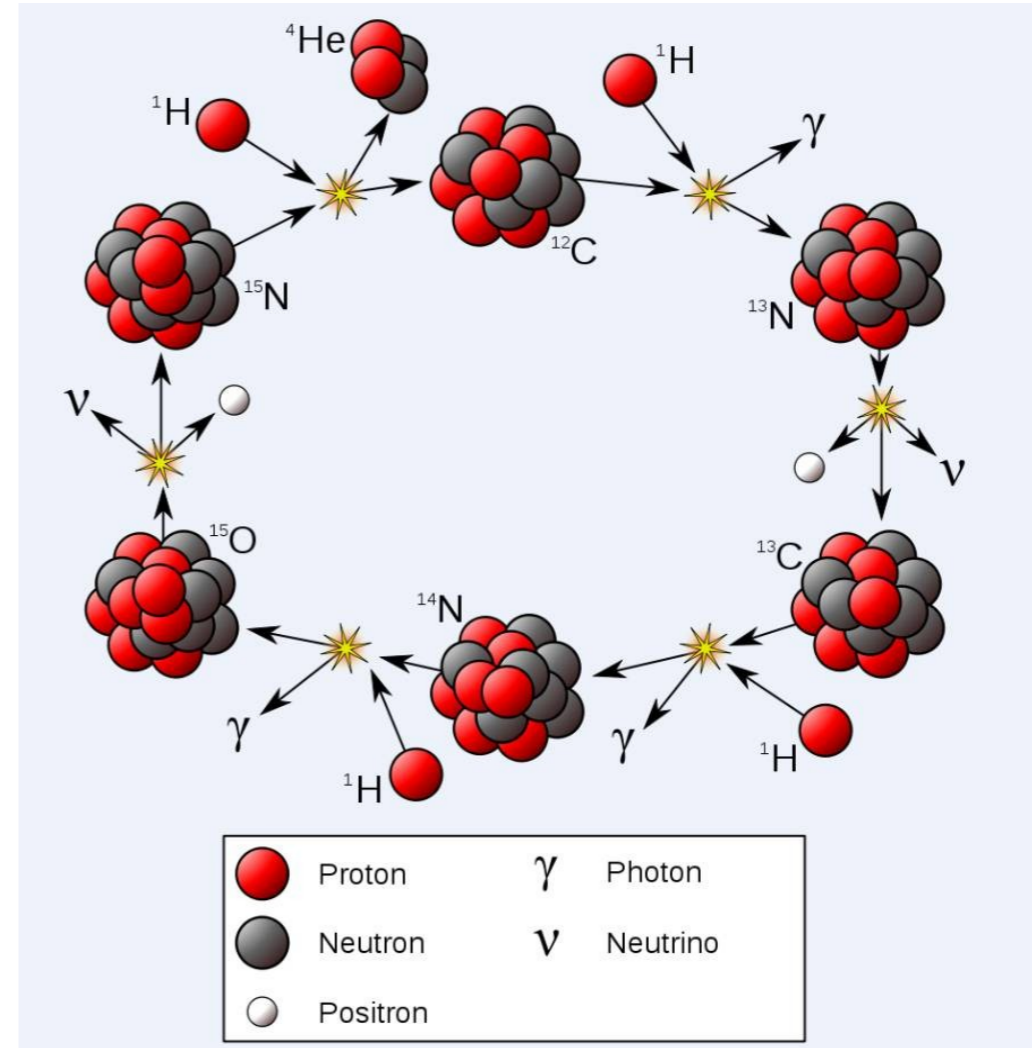
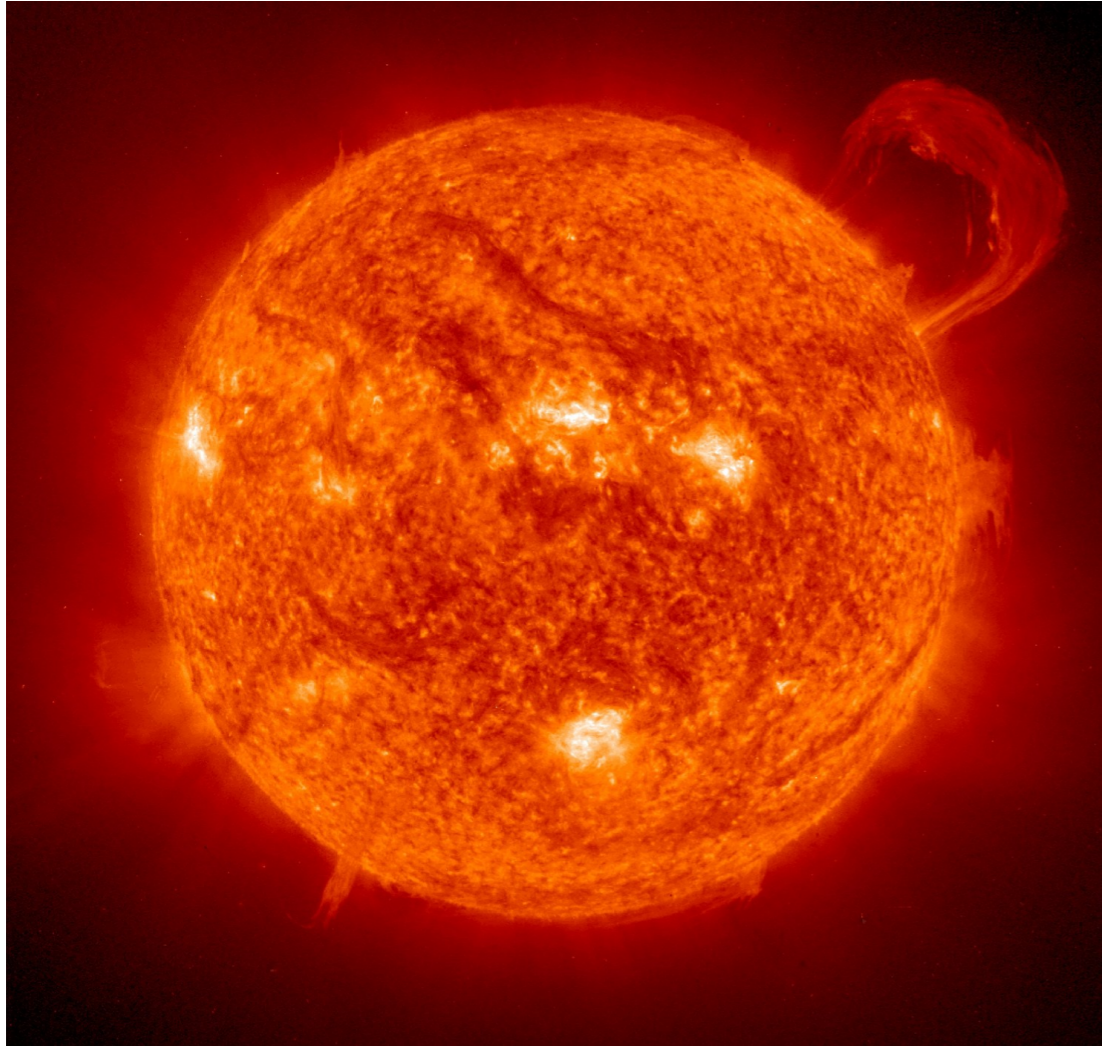


Plasma Thrusters



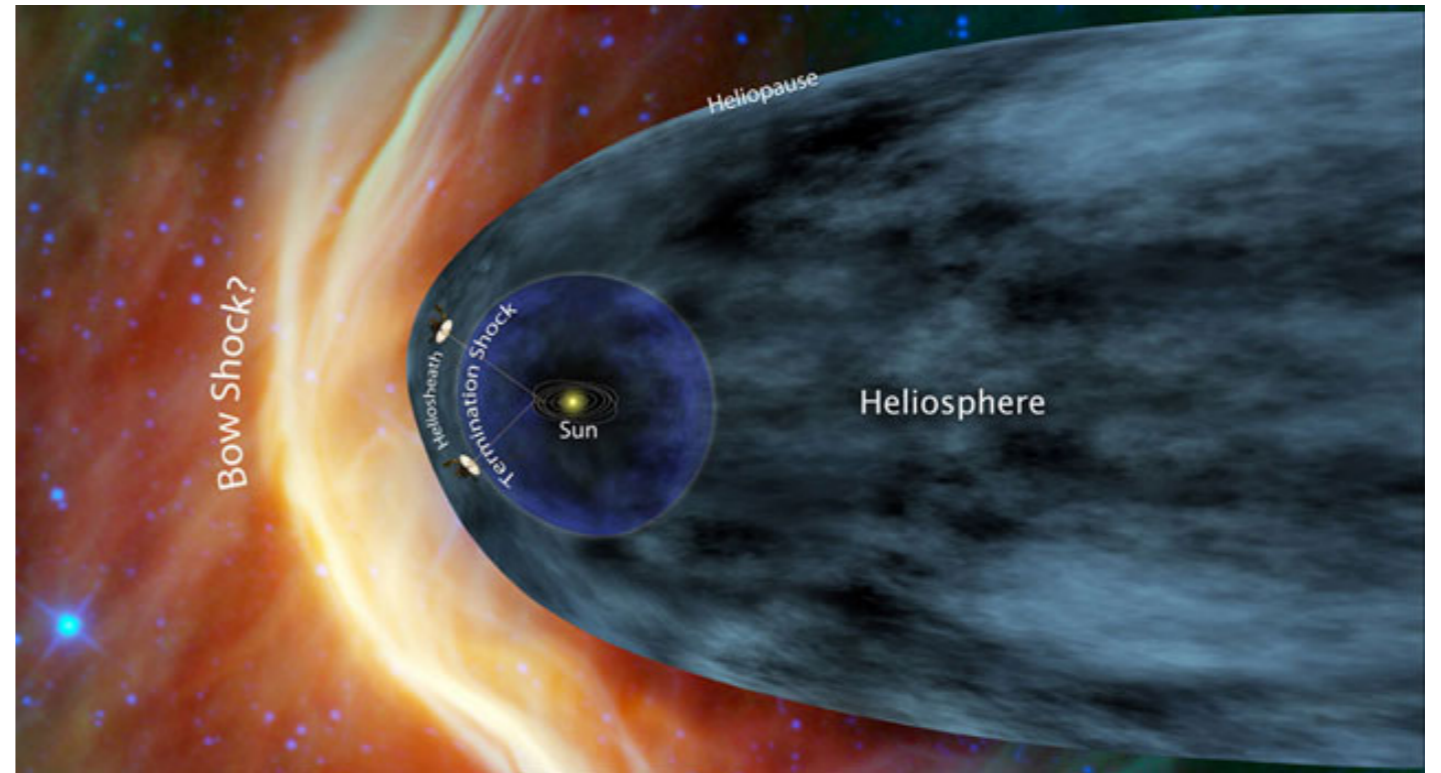
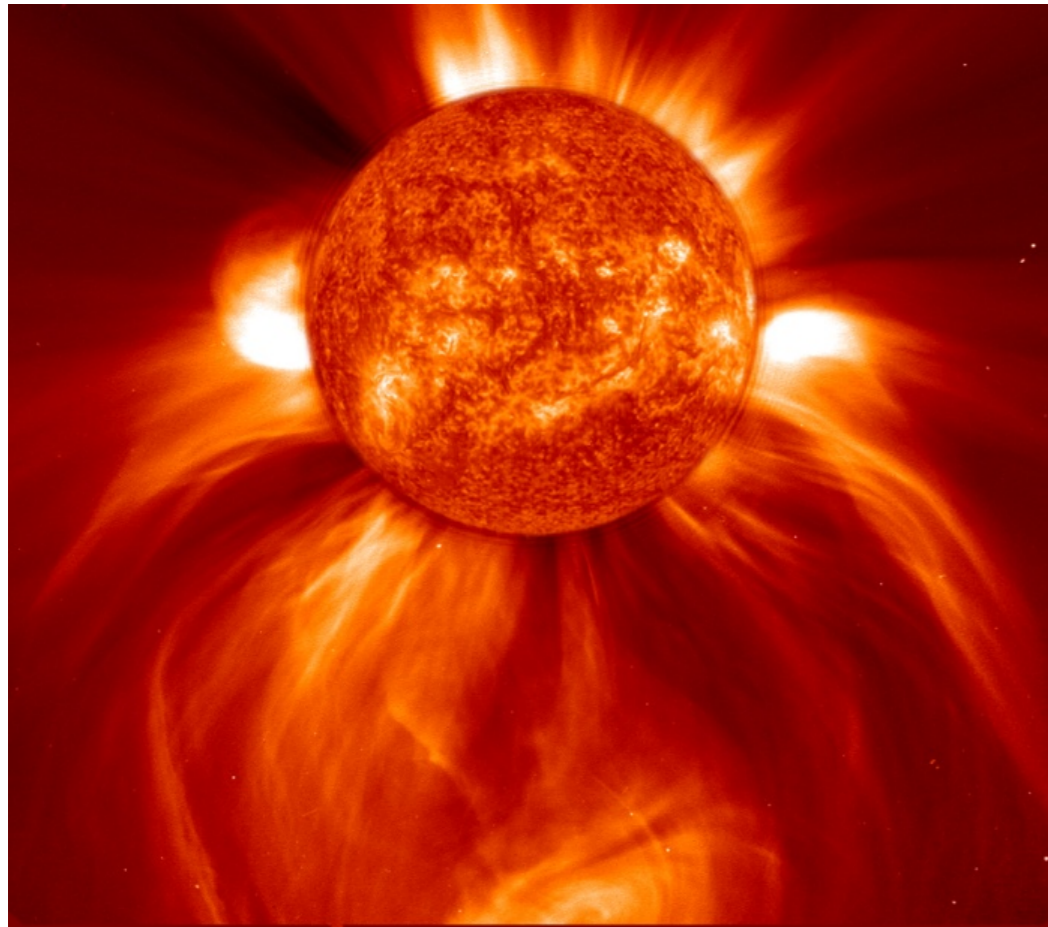
Photo: © ESA

The Sun



nuclear processes, solar structure, solar equilibrium and equations of state, radiation transport, convection

The solar Corona and the solar wind

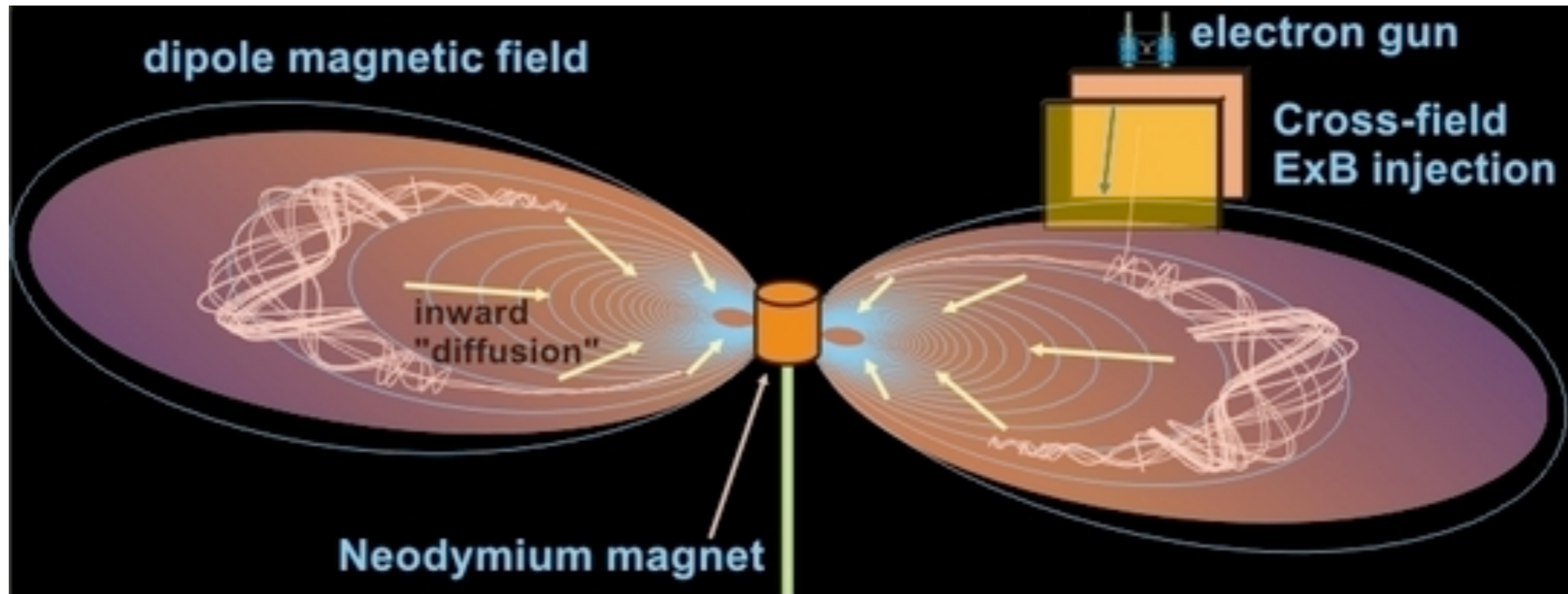


NASA, easa/SOHO

solar structure, magnetic fields in the sun(dynamo),
solar spots, the corona heating problem

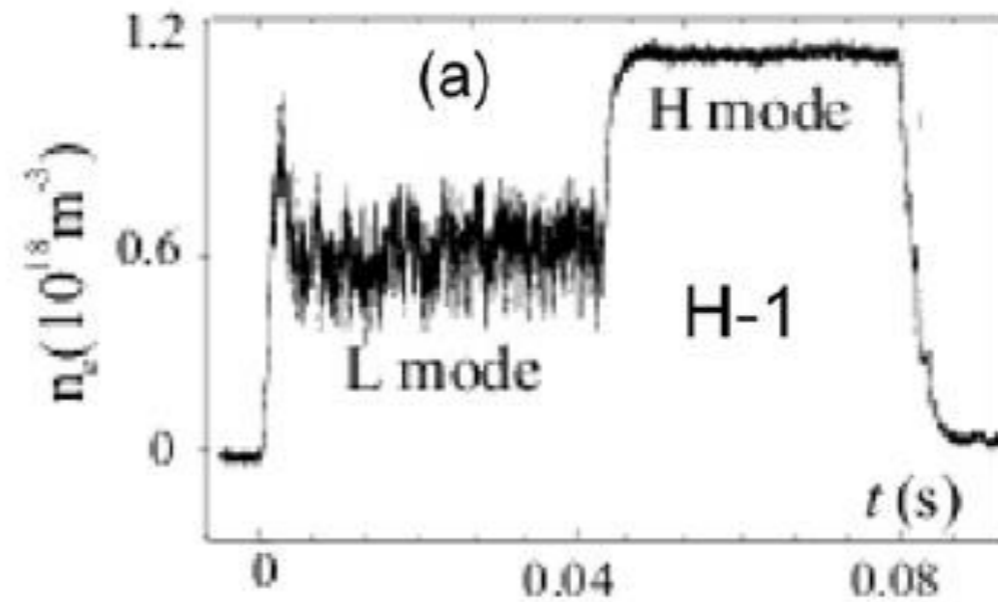
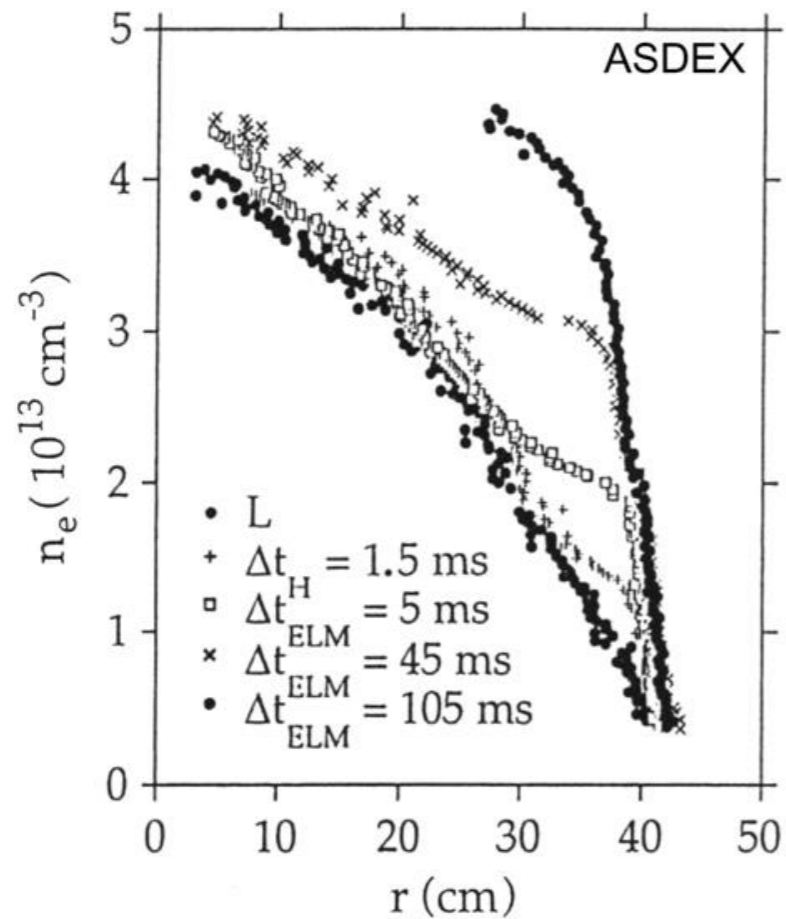
history, origin, Chapman model (static), Parker Model,
interaction with earth magnetic field

Electron-Positron Plasmas



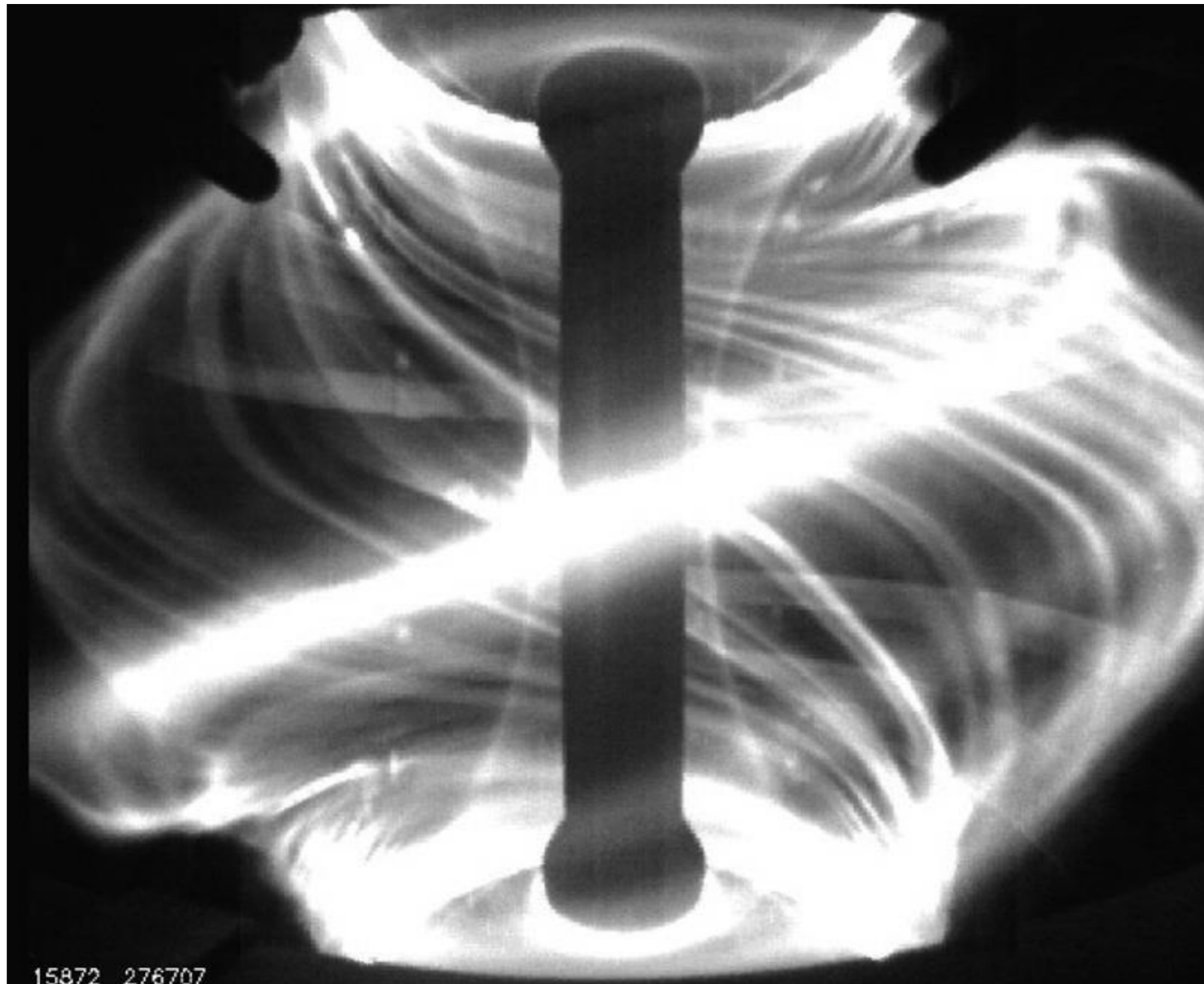
theoretical properties, experimental setup

Confinement regimes

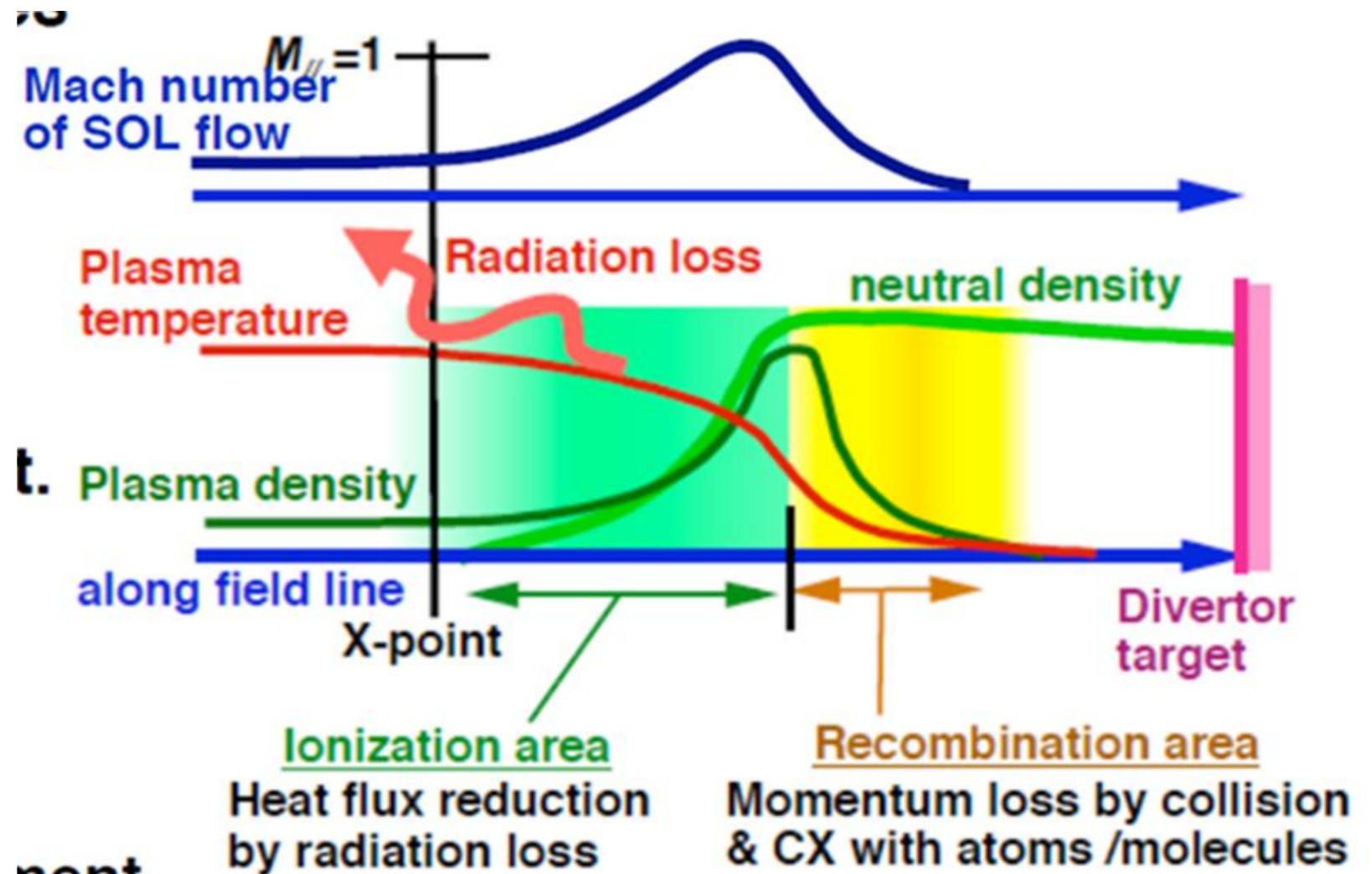
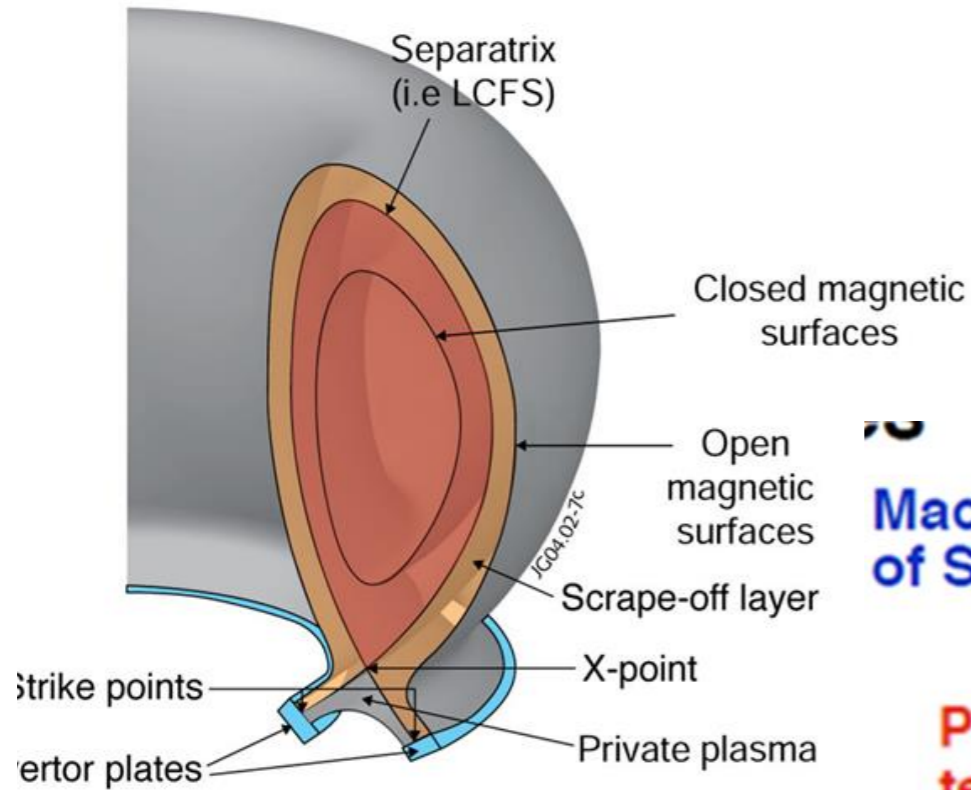


L-mode, H-mode, I-phase, I-mode, QH-mode, EDA-H-mode, Super-H-mode
turbulence, magnetic confinement, transport barriers

H-mode ELMs

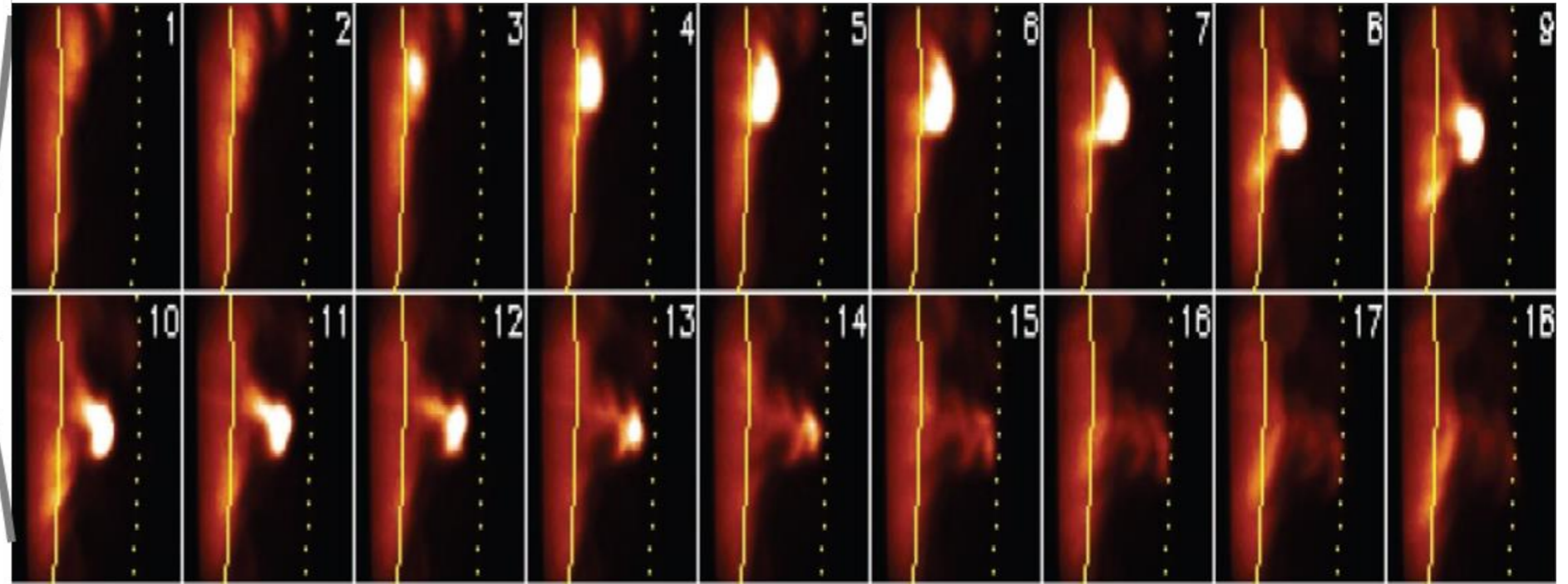
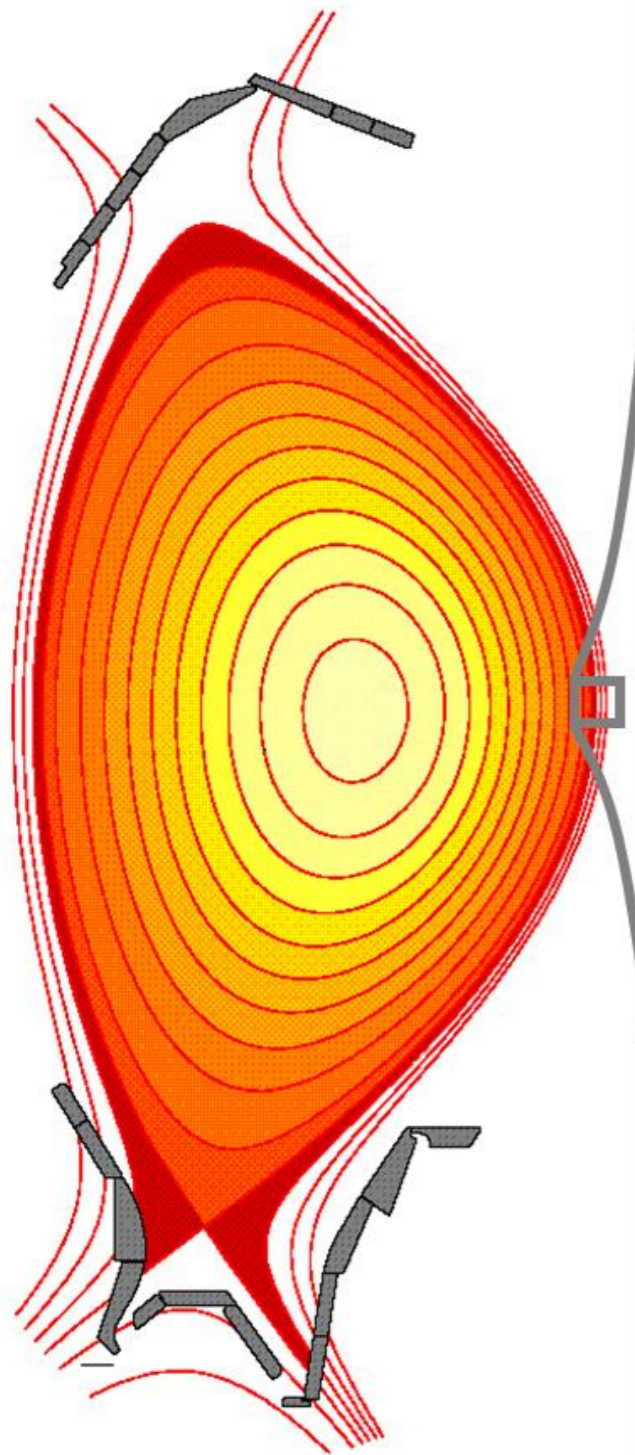


Divertor detachment



ment

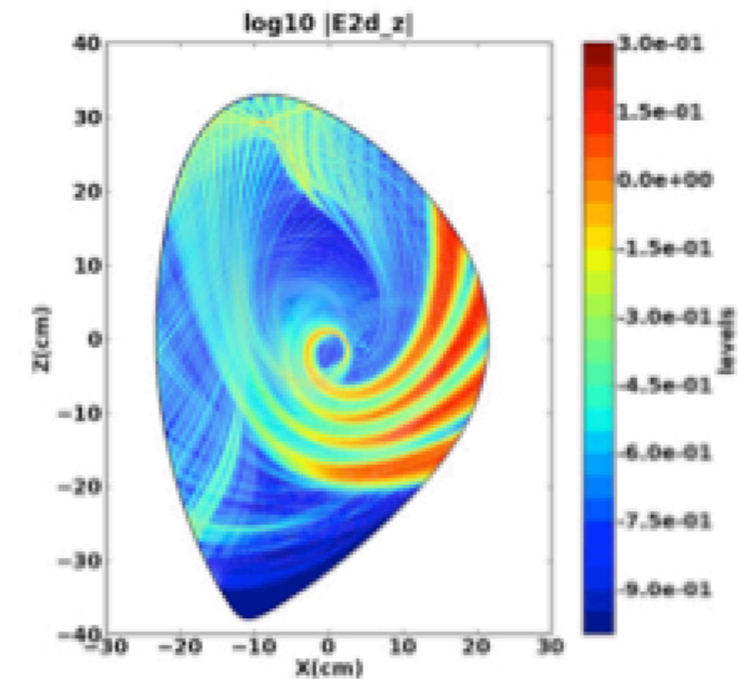
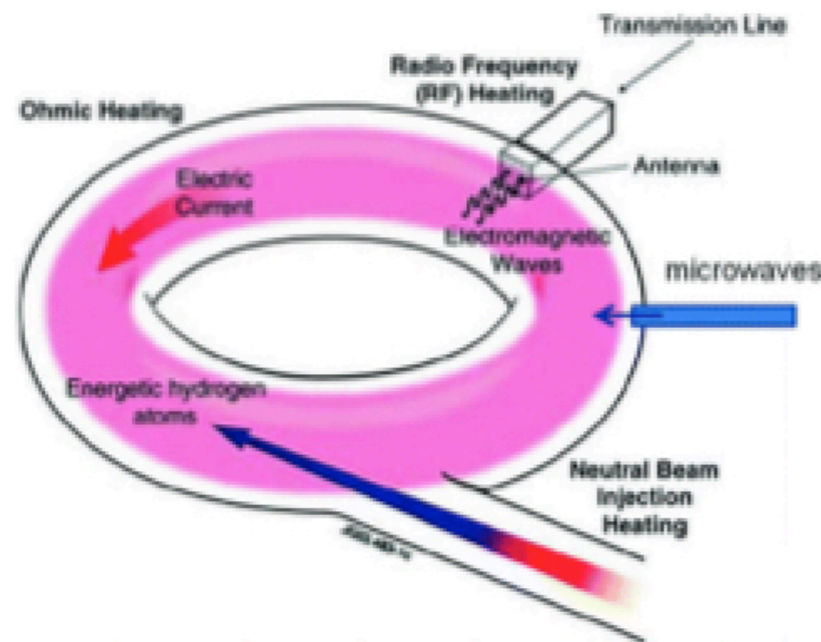
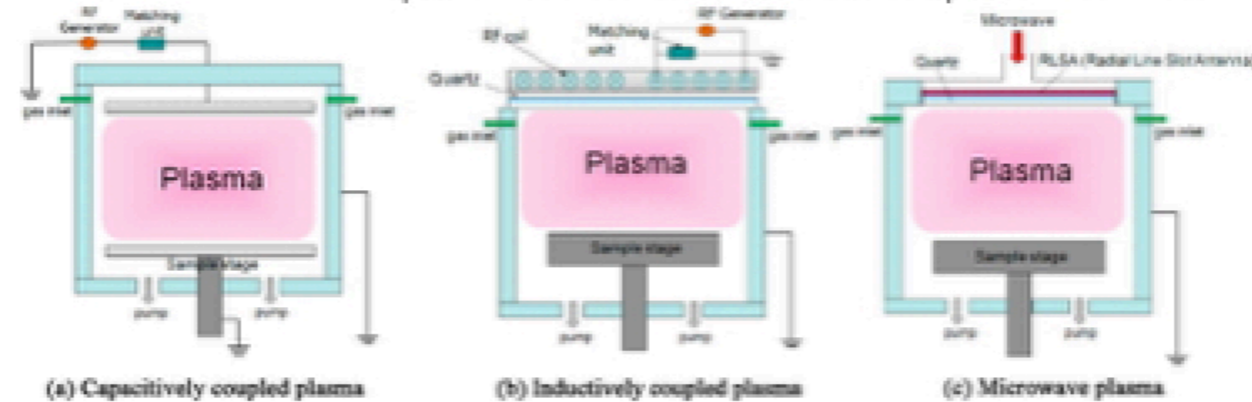
Plasma filaments in the scrape-off layer



+ Plasmadiagnostiken...

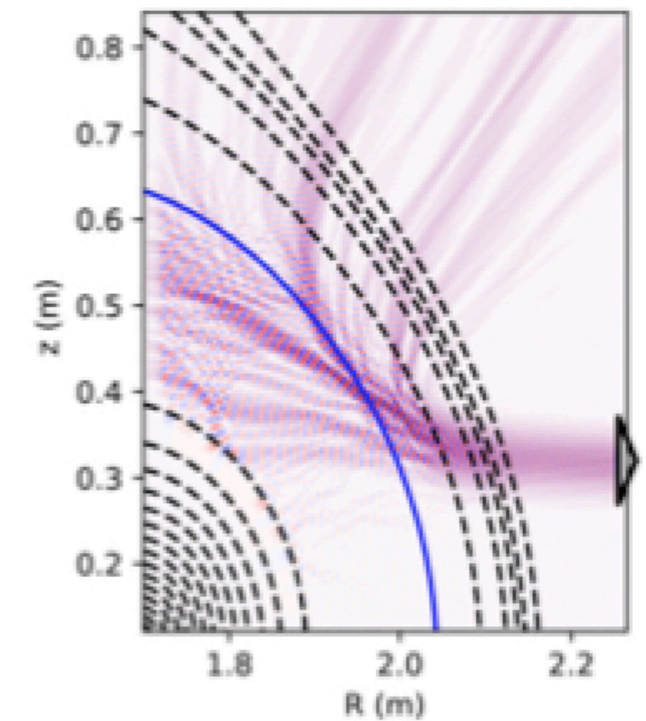
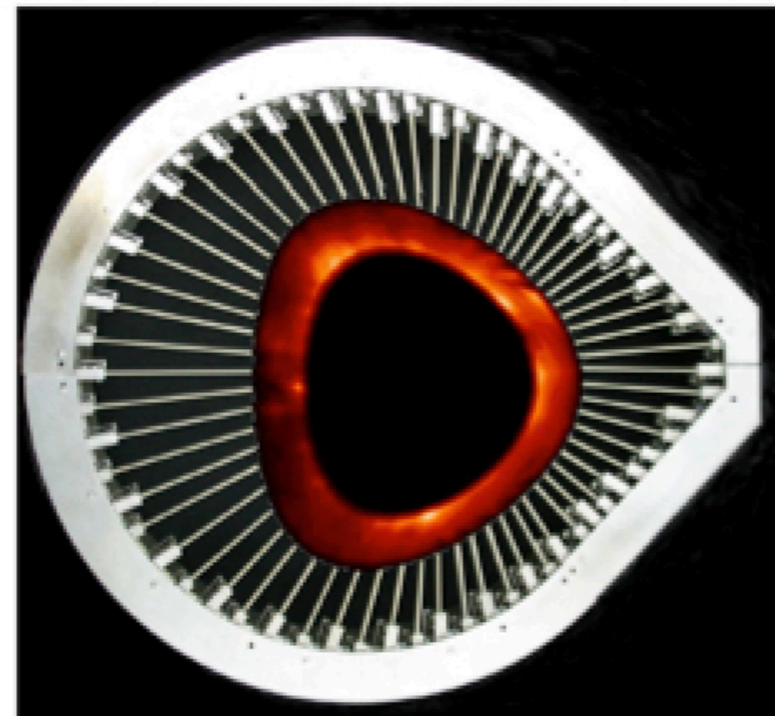
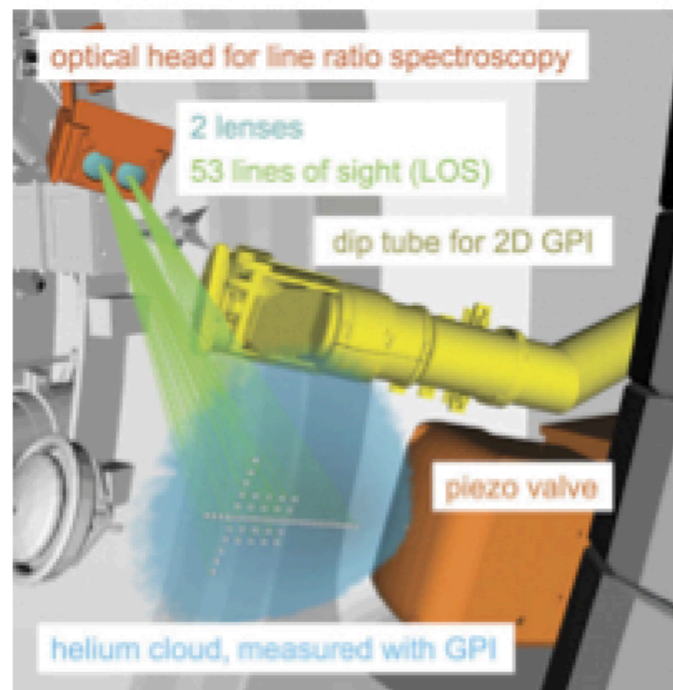
Plasma Heating

<https://www.sciencedirect.com/science/article/pii/S0927796X14000114>



Dolan T.J. (2013) Plasma Heating and Current Drive. Springer, London.
https://doi.org/10.1007/978-1-4471-5556-0_5

J.C. Wright, Phys. Plasmas 16, 072502 (2009)



C. Lechte et al., 2020 *Plasma Sci. Technol.* 22 064006