

Equilibrium data/parameters for strongly non-linear EP dynamics based on AUG experimental results: EGAMs, RSAEs, BAEs and TAE bursts

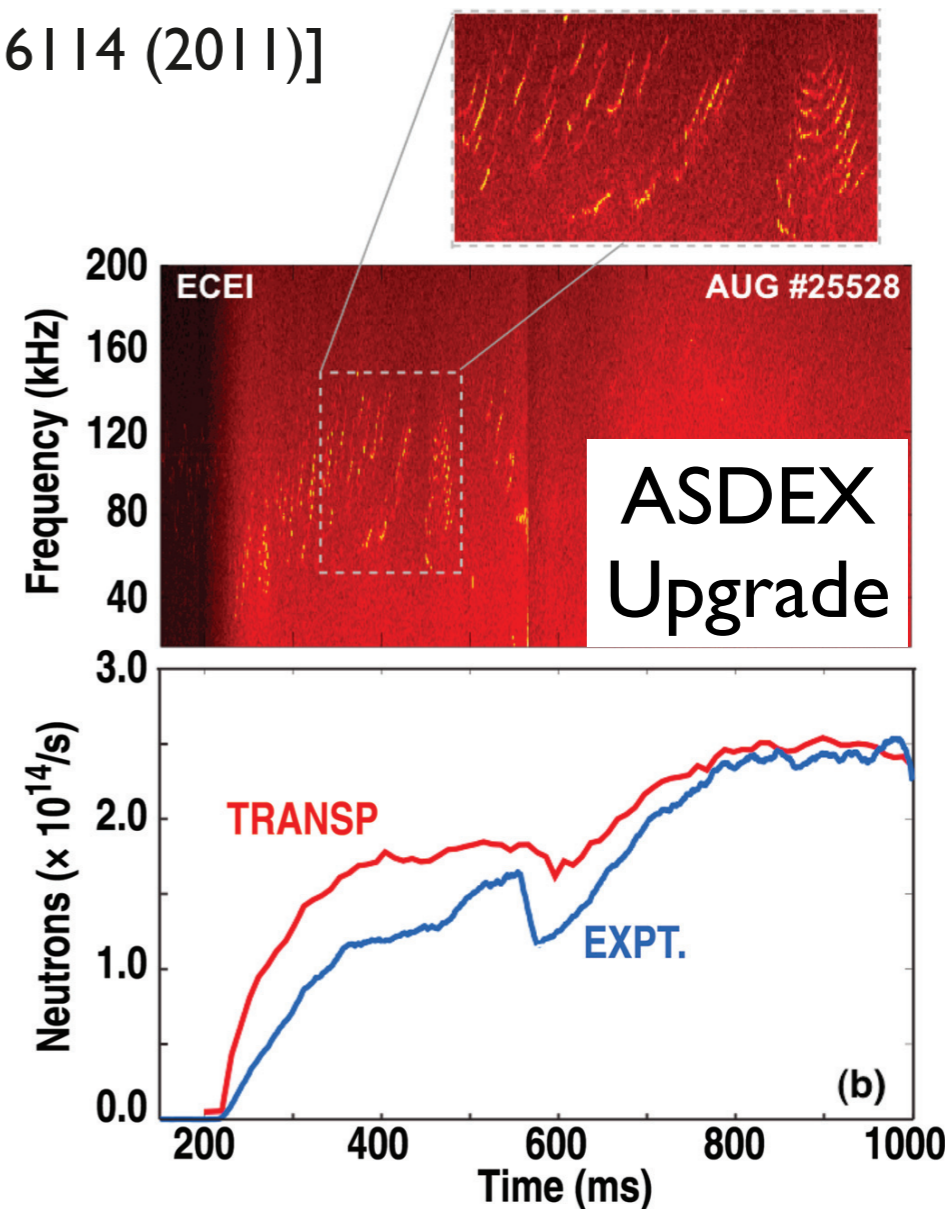
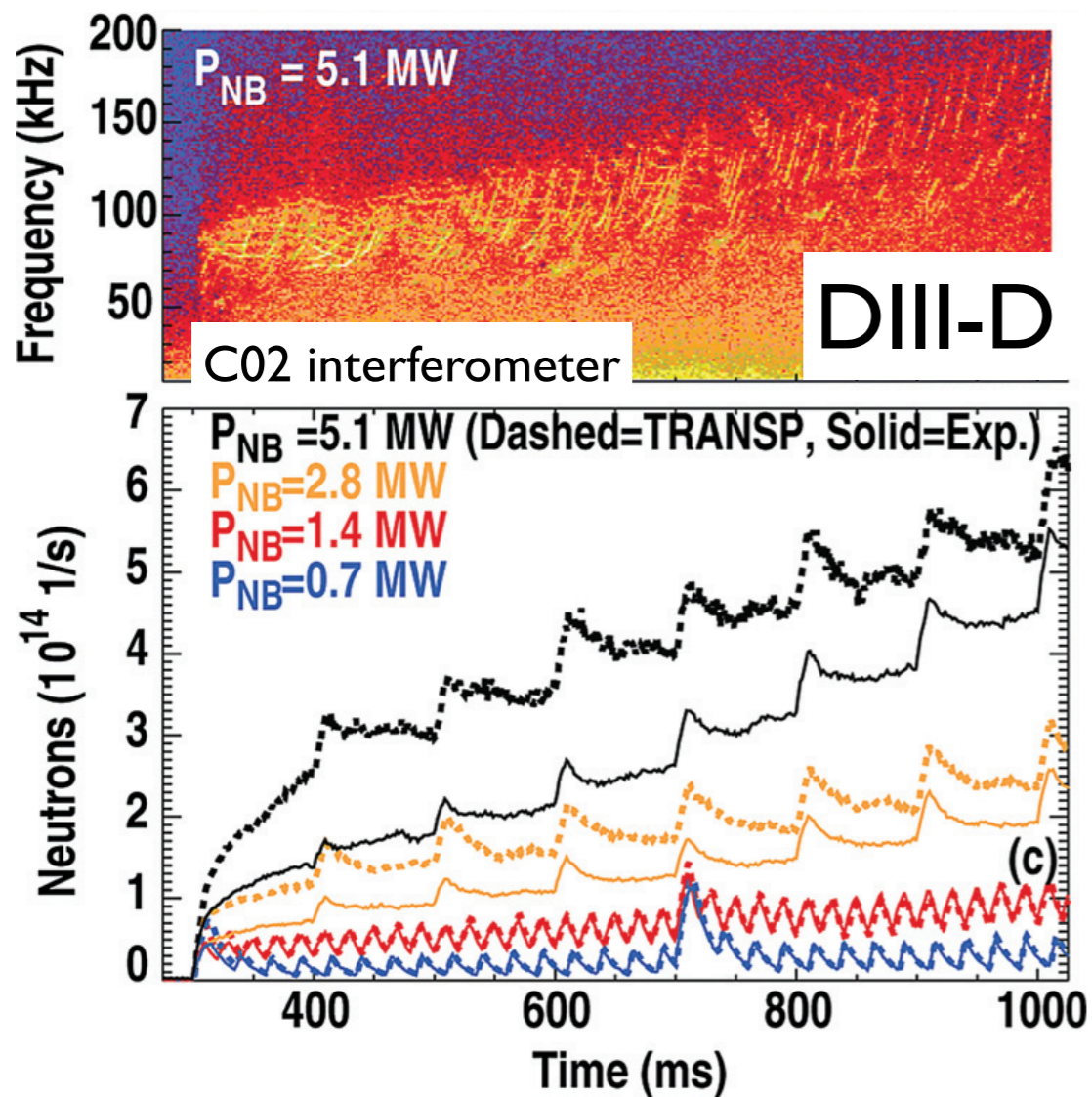
Ph. Lauber, IPP

with input from: C. Di Troia

ASDEX Upgrade Team: M. Maraschek, B. Geiger, V. Nikolaeva, L. Guimares, A. Mlynek, I. Classen, V. Igochine, A. Gude, Ch. Hopf, **M. Dunne,**

'sea' of unstable small amplitude modes for low energy [80/60keV] on-axis NBI drive in ramp-up

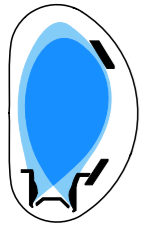
[VanZeeland et al Phys. Plasmas 18, 056114 (2011)]



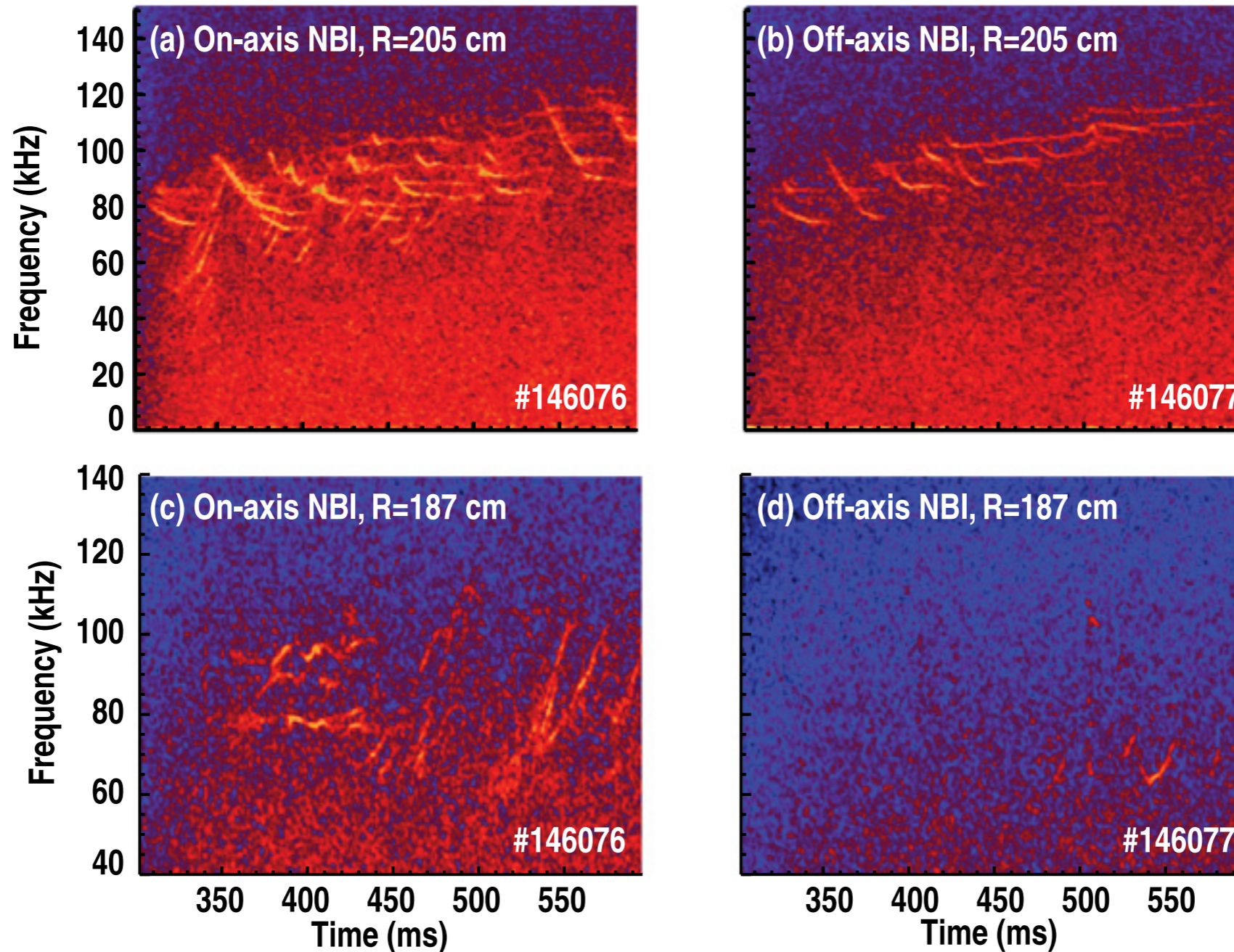
in these cases, neutron deficiency and EP transport well understood and measured (FIDA, FILD), successful comparison to quasi-linear hybrid models (amplitudes given by the experiments)

[R. White et al, 2011]

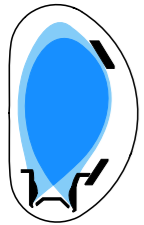
similar to case studied by Vlad et al, 2009, also Todo [2013/14/15], US codes [TAEFL, GTC, GYRO] ITPA DIII-D case [Lauber 2015]



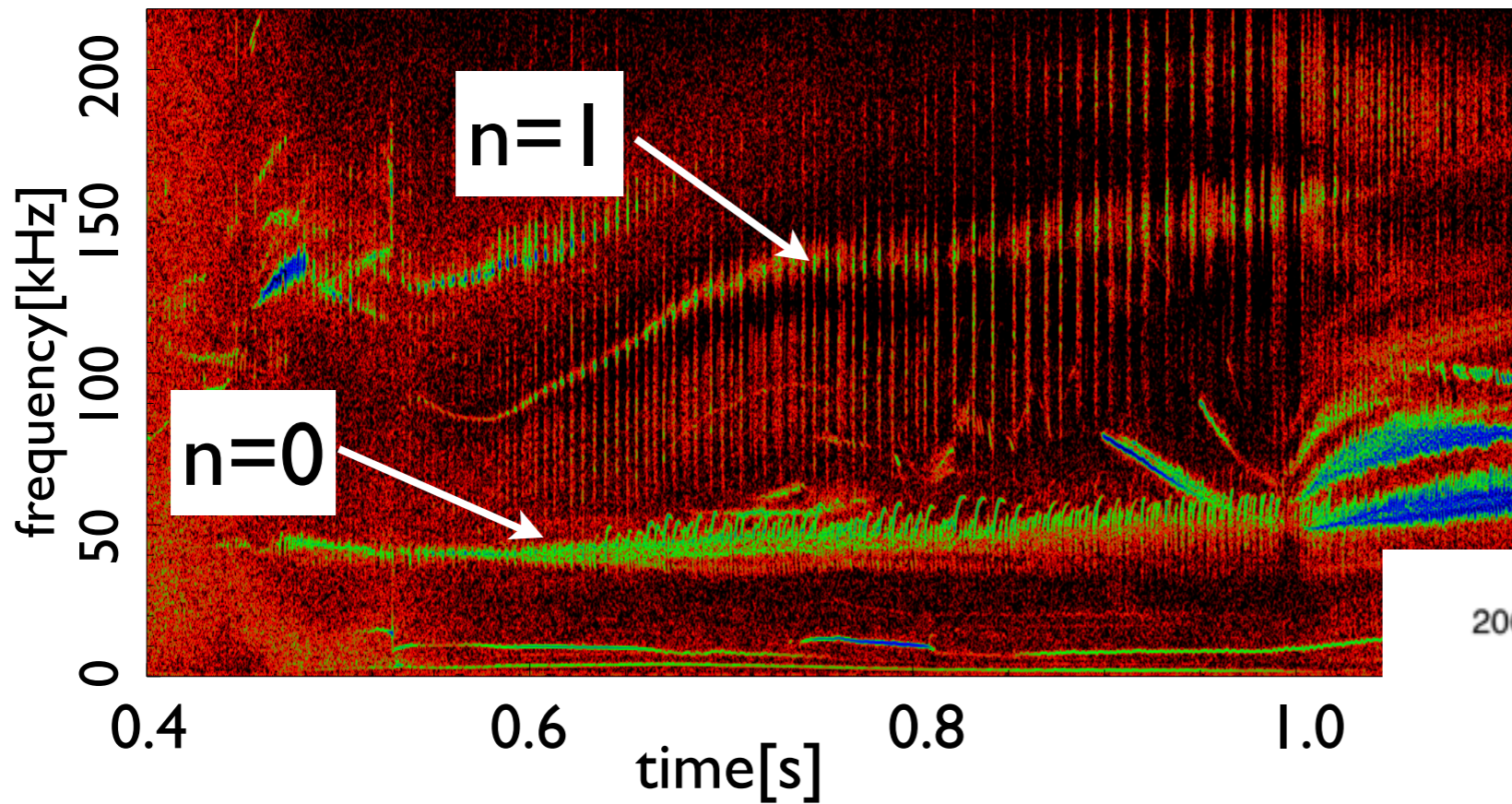
[B. Heidbrink et al , NF 53, 2012]



conclusions: overlap of many small amplitude modes dominates EP transport -threshold to strongly non-linear behaviour? ITER?



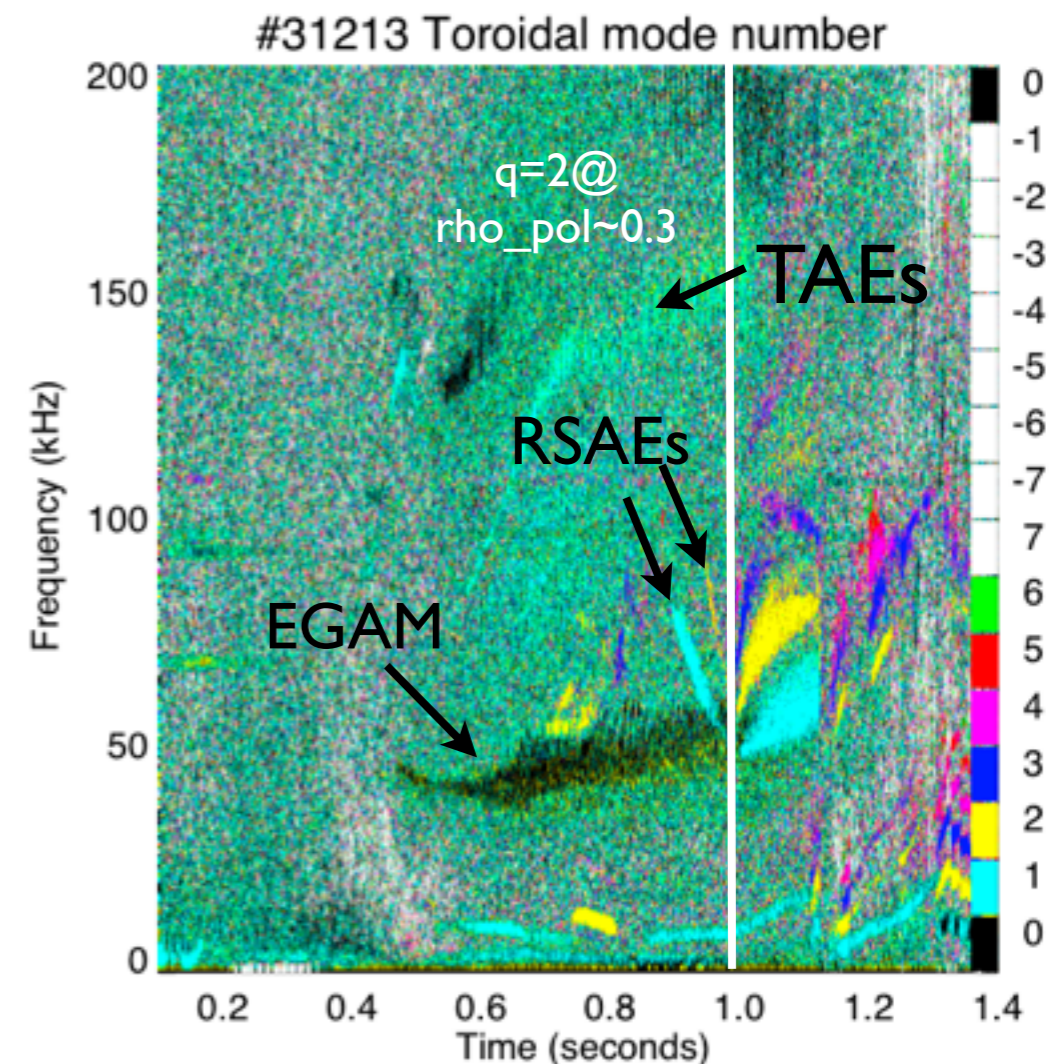
ASDEX Upgrade: early off-axis NBI drive [93keV]: bursting EGAMs, RSAEs and TAE/EPMs



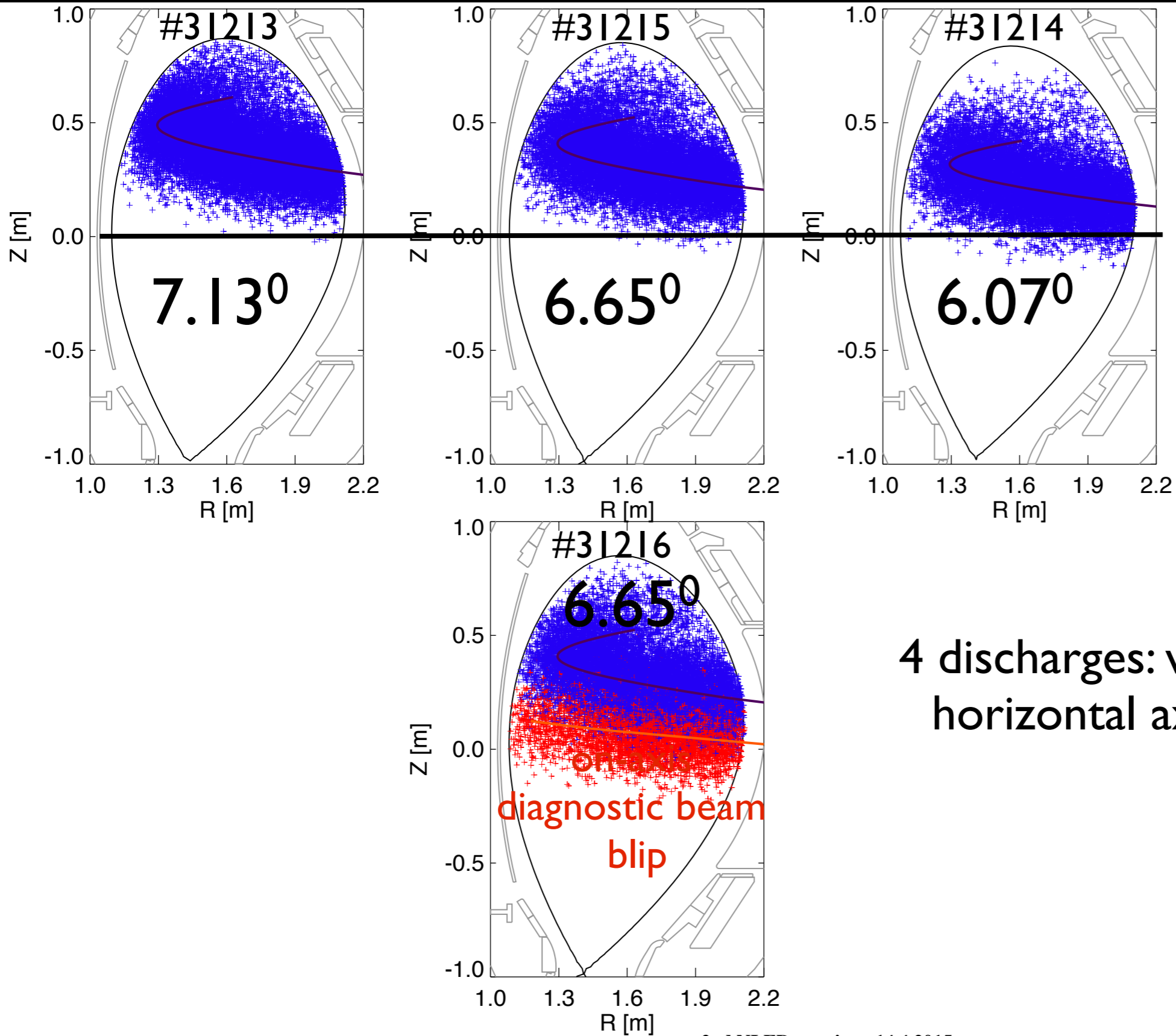
[July 2014]

$B_0=2.2\text{T}$
 $I=0.6-1.0\text{MA}$
 $P_{\text{beam}}=2.5\text{MW}$

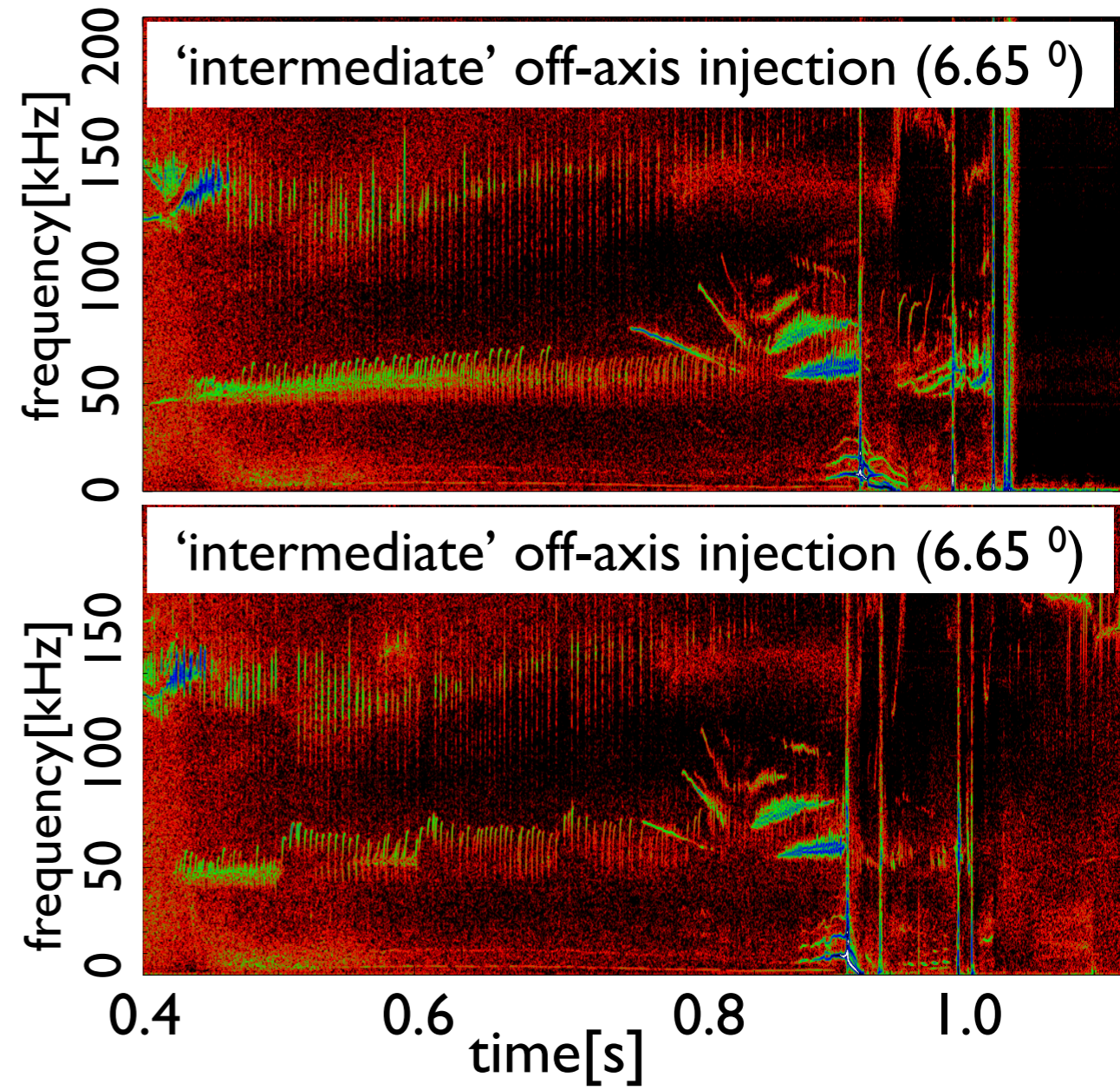
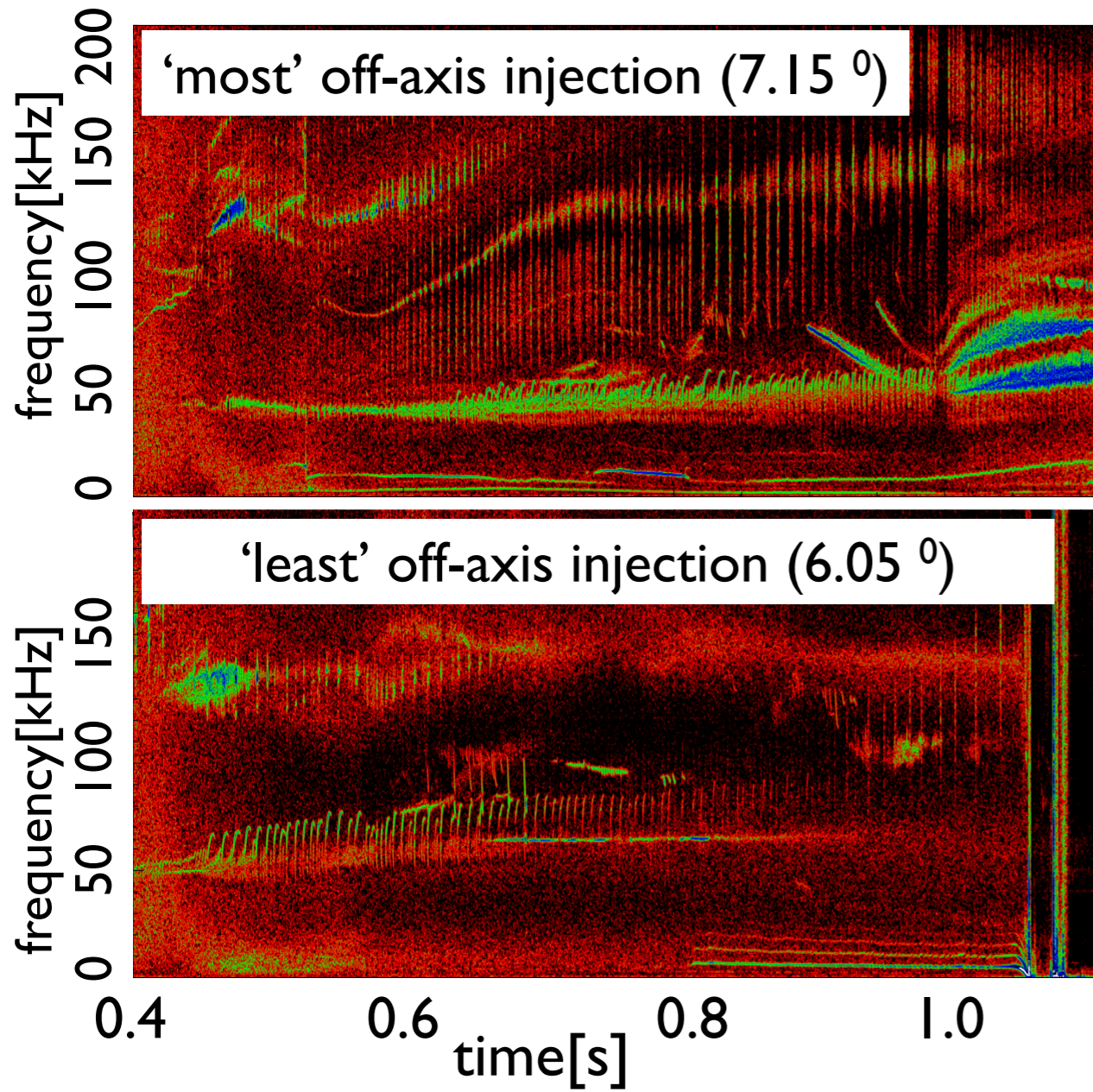
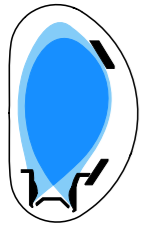
- rather well reproducible scenario
- no 'sea' of Alfvénic modes (TAEs/RSAEs)
- but strongly chirping $n=0$ modes and $n=1$ 'bursts'

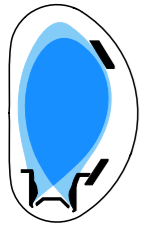


beam injection geometry

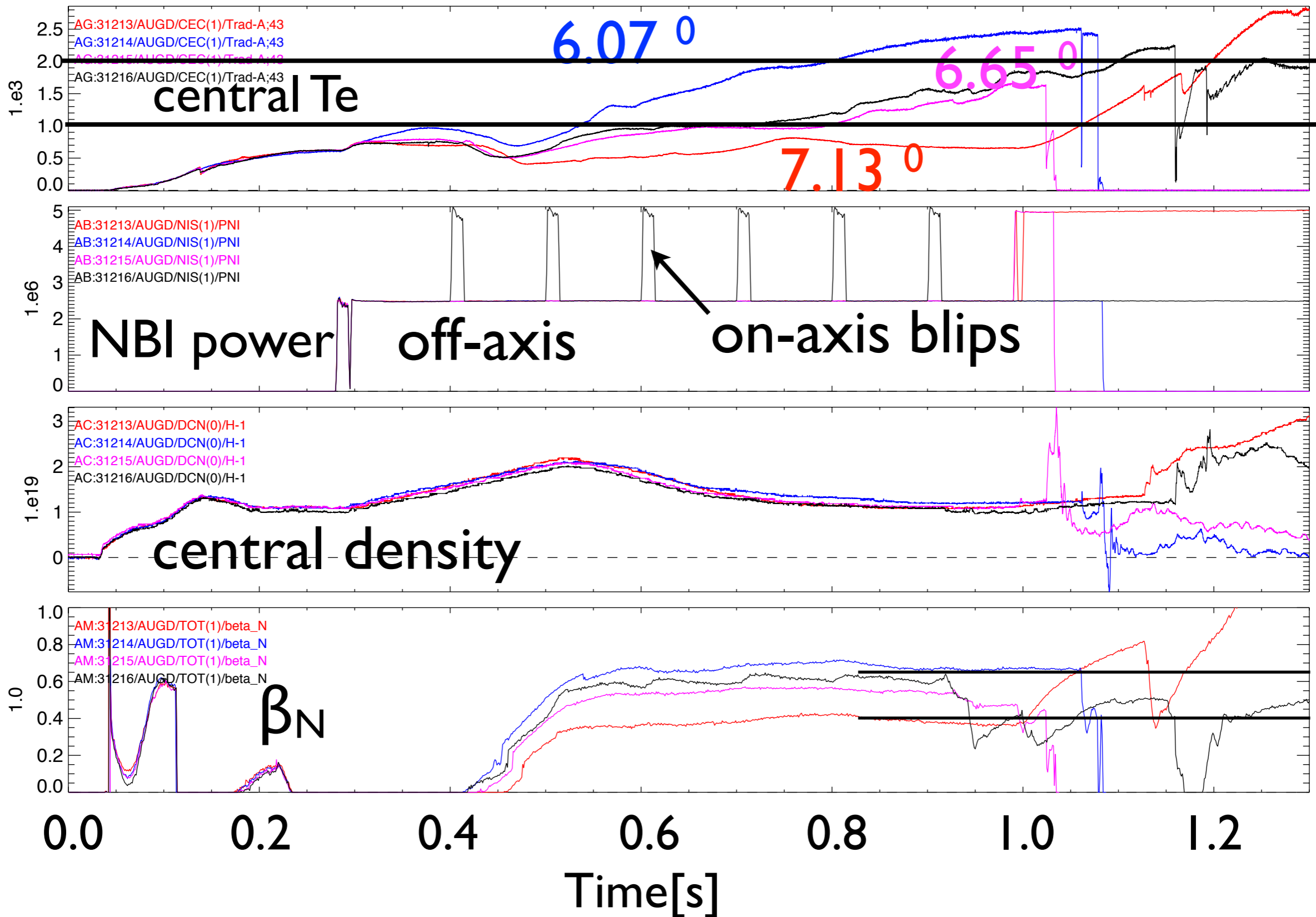


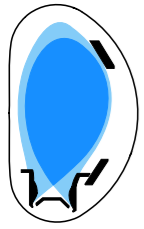
4 discharges: vary angle between horizontal axis and beam-line



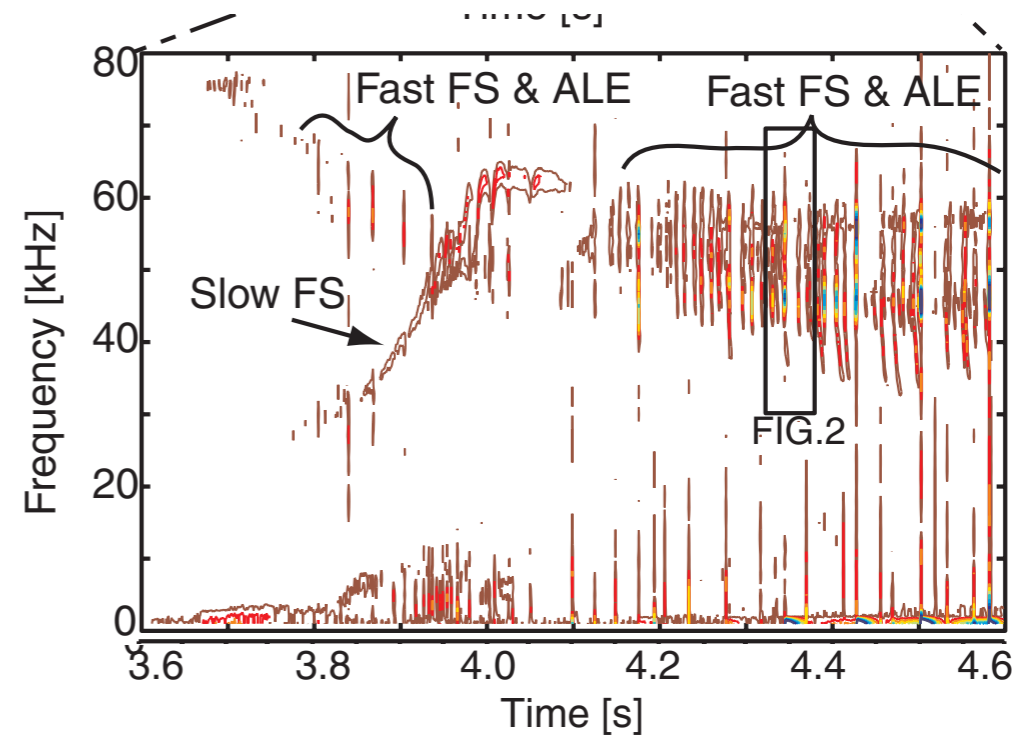
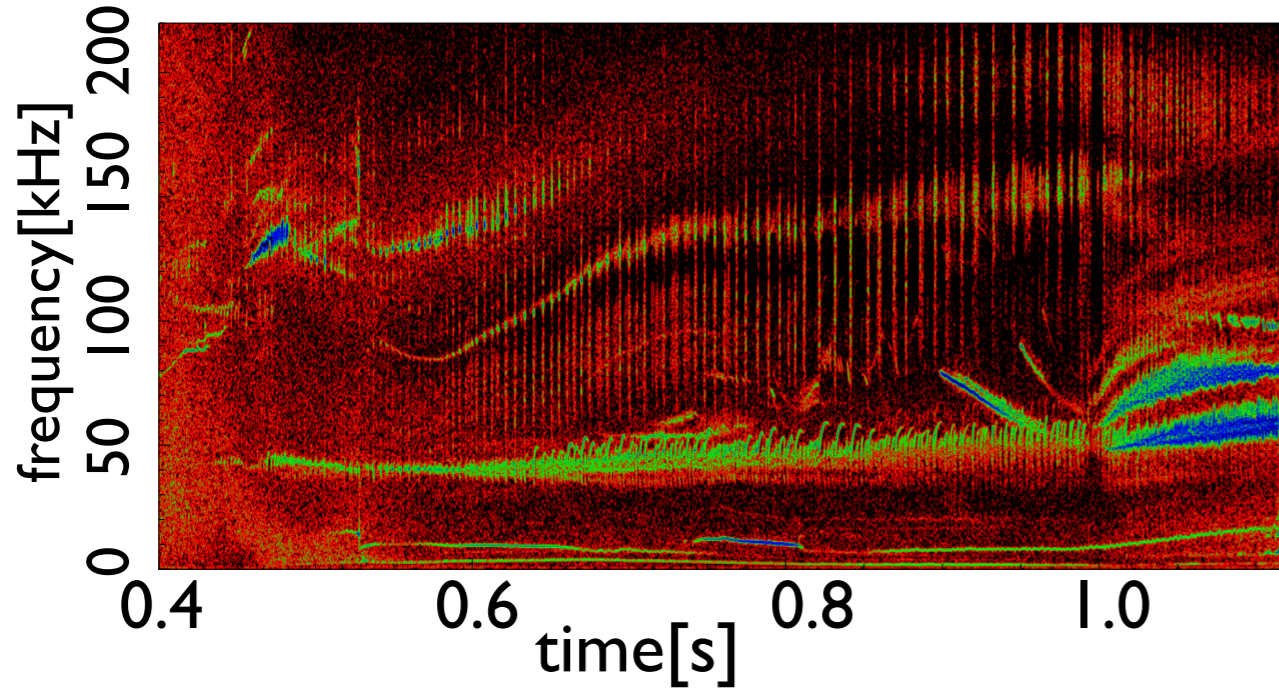


time history of 4 discharges with different NBI injection angles





n=1 TAE bursts seem to have some similarity to 'fast sweeping' and 'ALE' events at JT-60U



JT-60U: K. Shinohara et al, 2002-2004

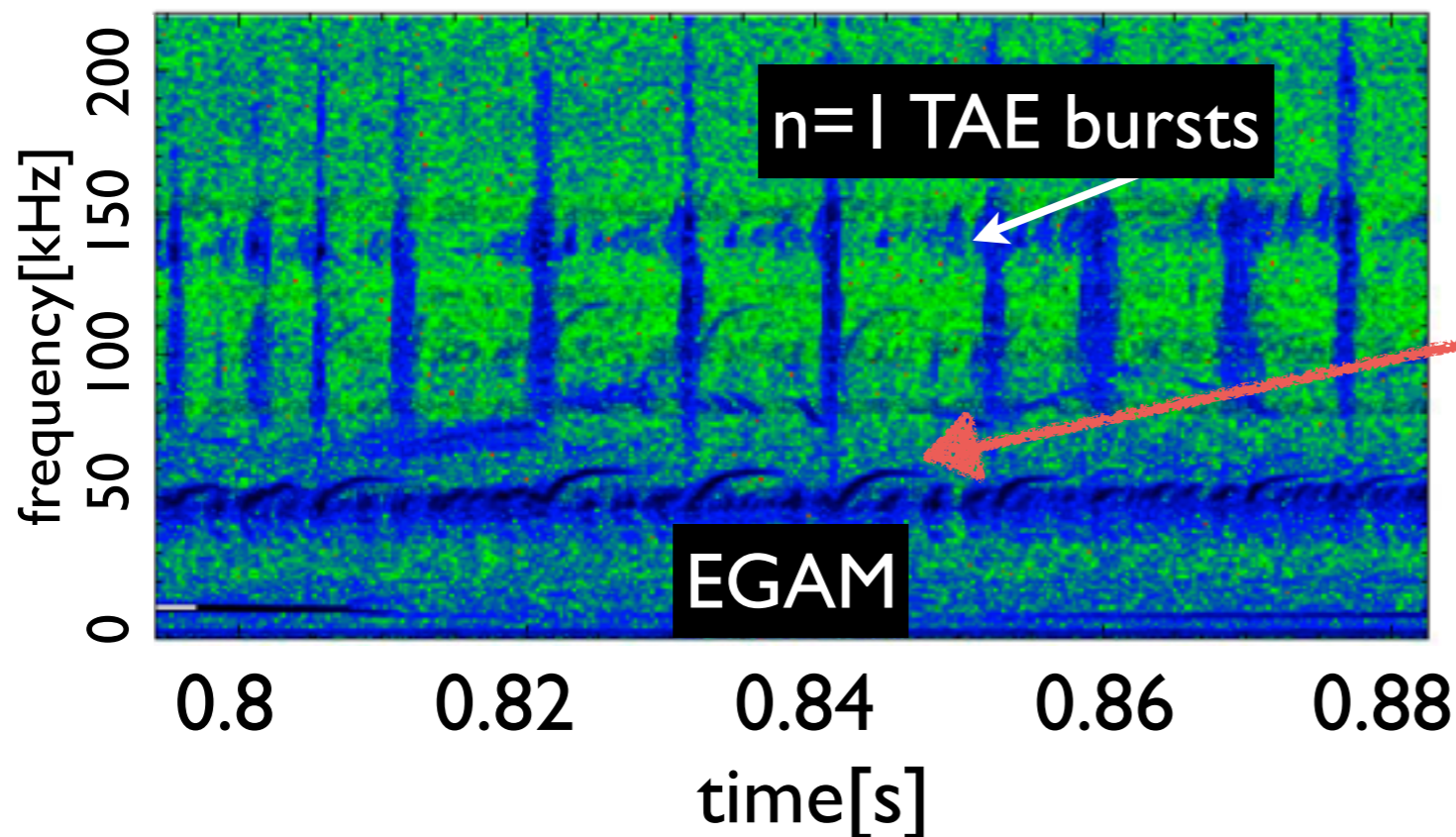
JT-60U: $v_f/v_{A0} \sim 1.3$; NB: 350keV

DIII-D: $v_f/v_{A0} \sim 0.4$; NB: 80keV

AUG: $v_f/v_{A0} \sim 0.45$; NB: 93keV

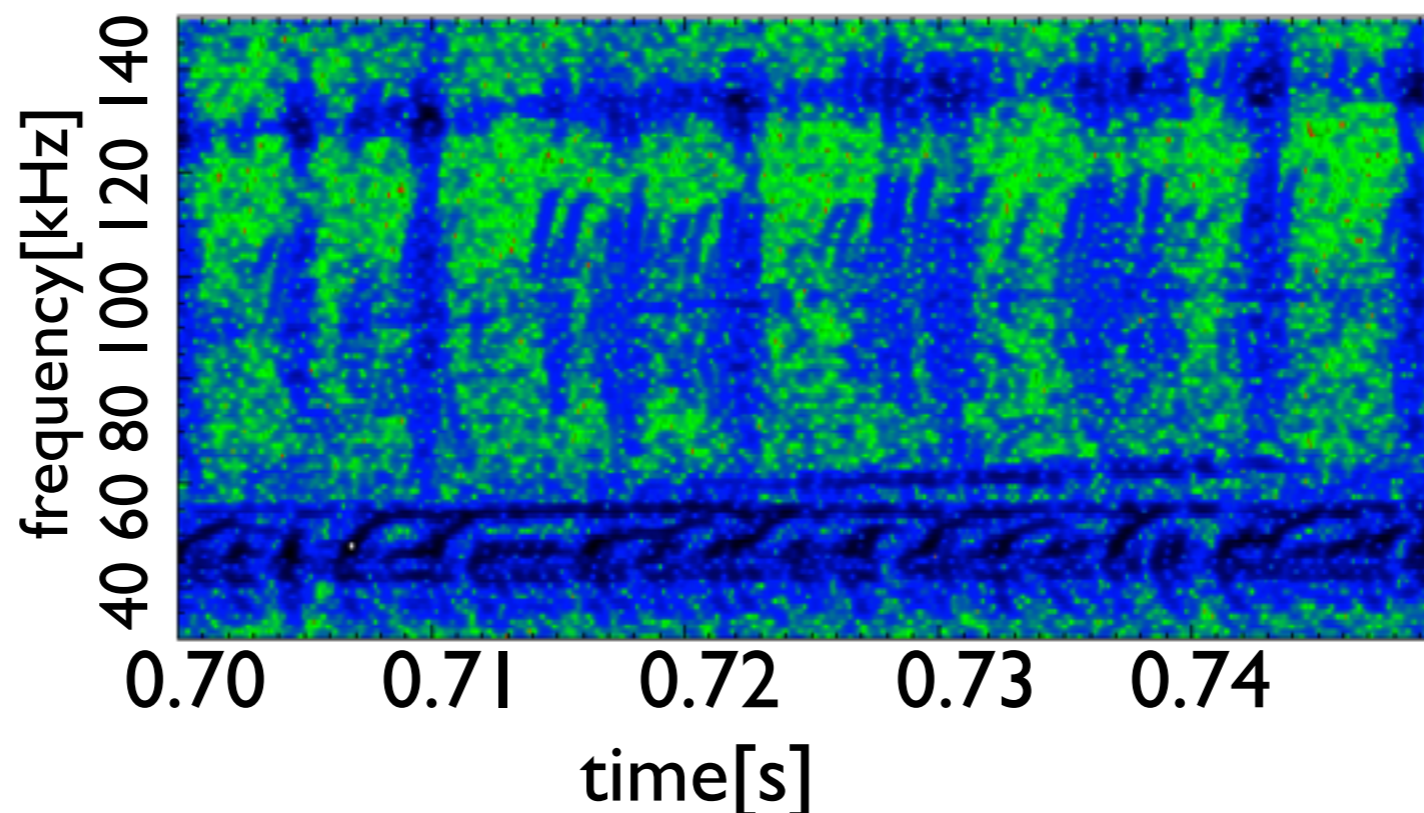
additional new physics: n1 EGAM interaction, drive at $v_A/3$

phase space coupling: $n=1$ TAE bursts seem to trigger EGAMs



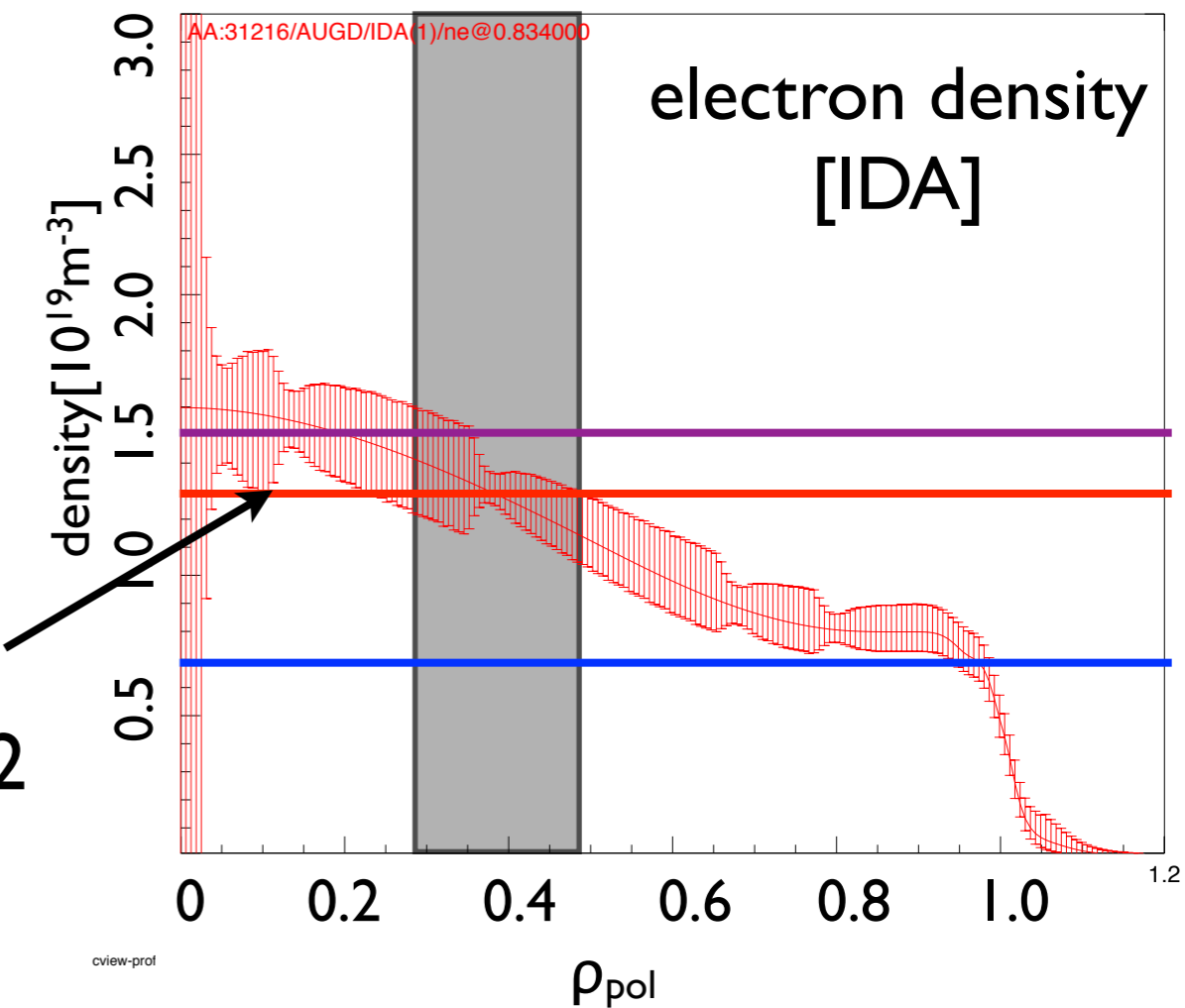
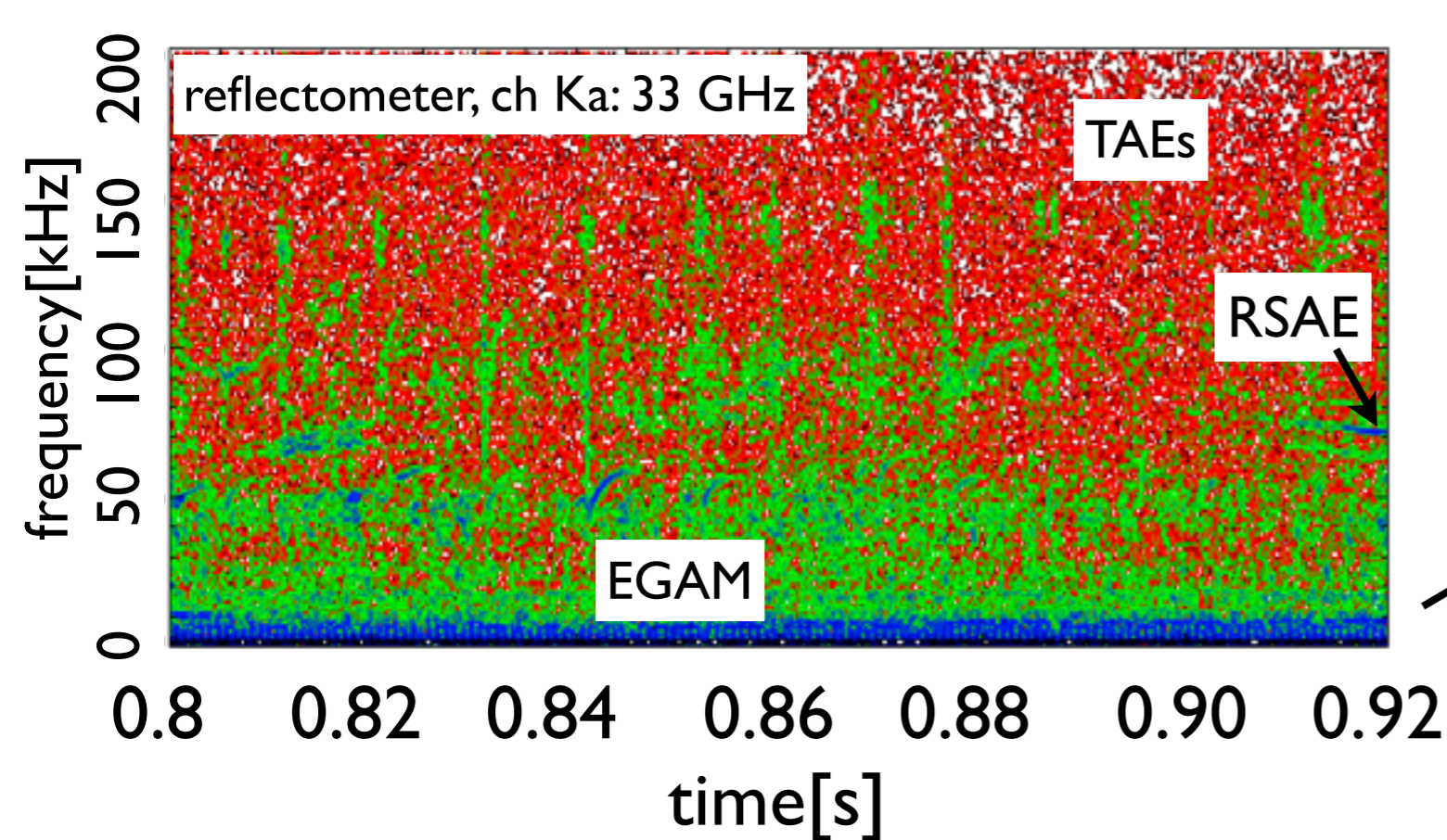
experimental evidence that EGAMs change radial width during $n1$ evolution [L Horvath, master thesis 2015]

modes around 100kHz: probably modes propagating in the electron diamagnetic direction ($2 f_{rot}$)



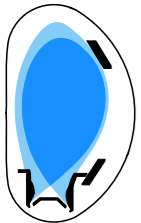
from reflectometry (hopping frequency) and soft-X-ray measurements:
 EGAMs, TAEs, RSAEs and intermediate frequency modes are visible in the same channels \Rightarrow similar radial location at $\rho_{pol} \sim 0.2-0.5$,

EGAMs more core localised (0.1-0.5), TAEs more outside (0.3-0.7)



[V. Nikolaeva, L. Guimares]

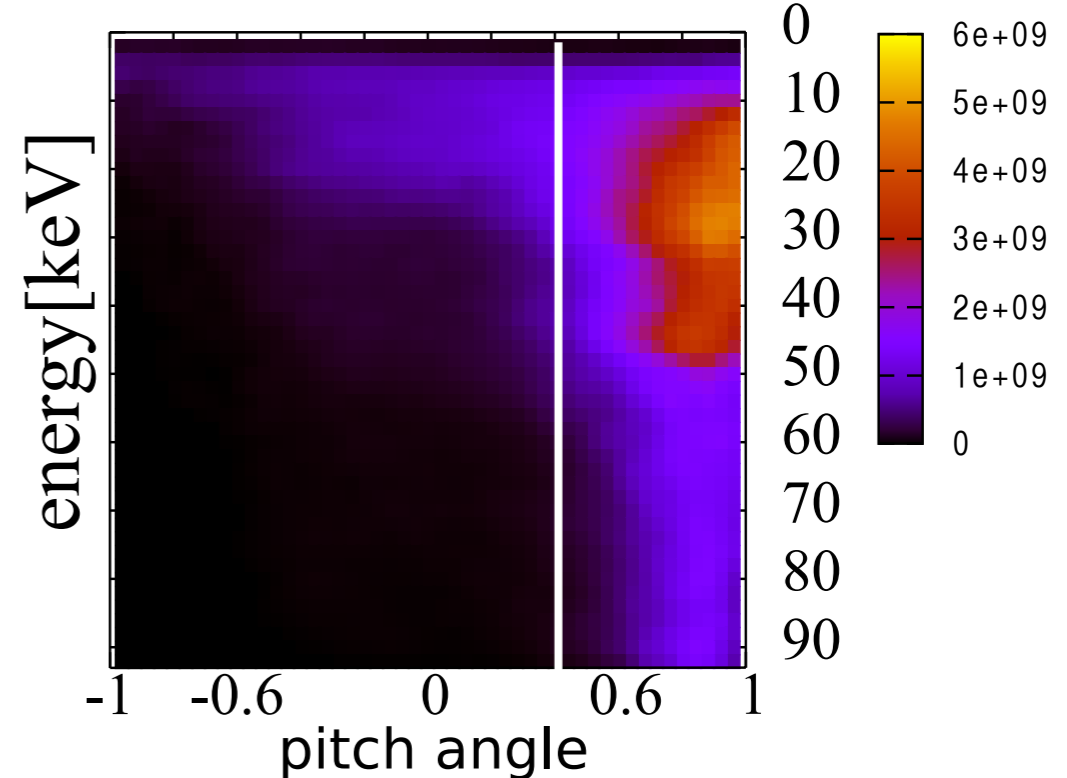
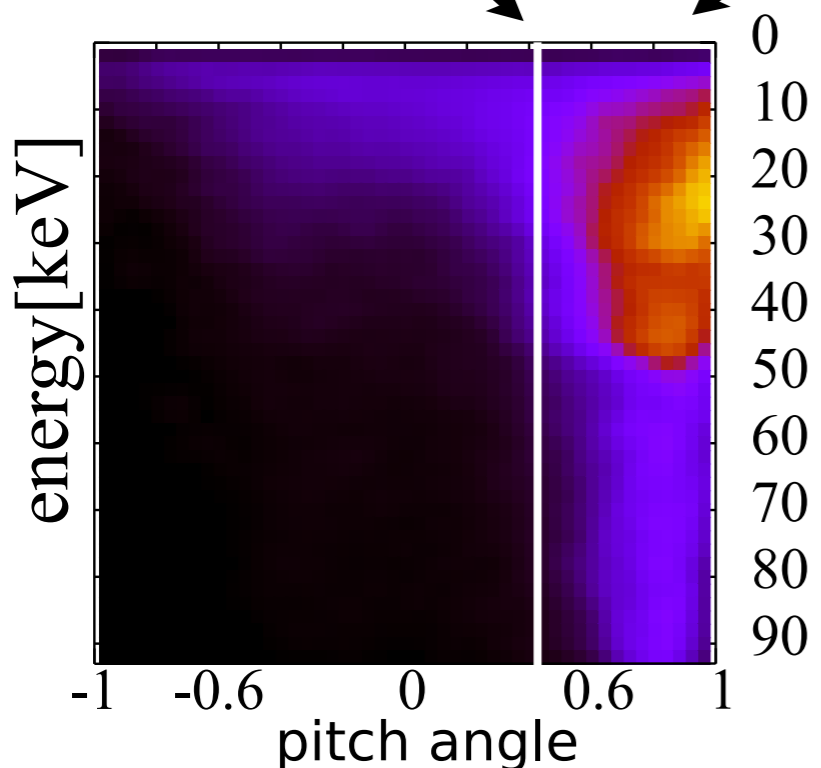
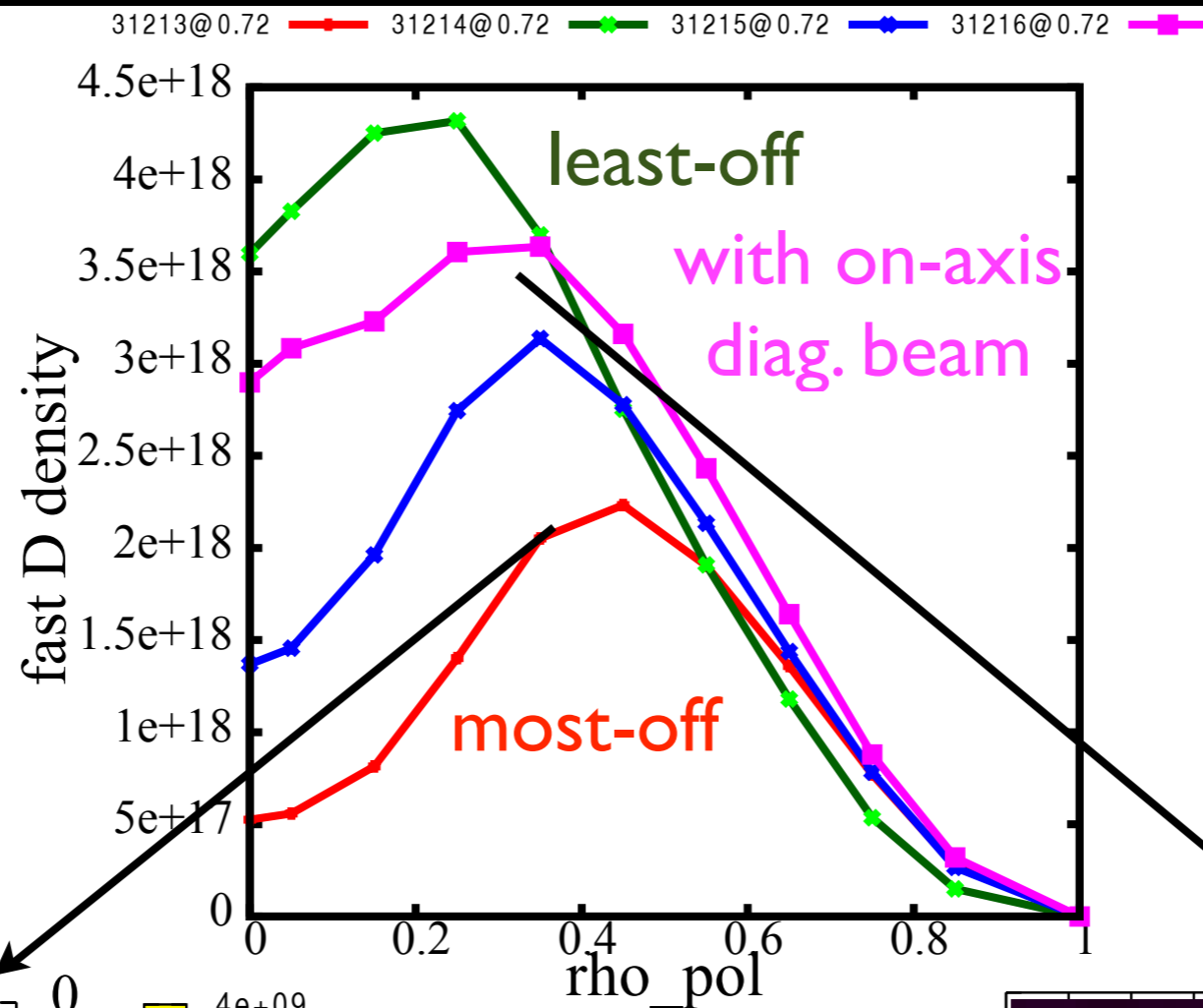


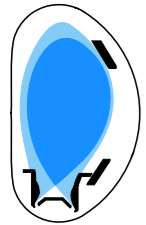


NBI distribution function

radial dependence:
off-axis maximum

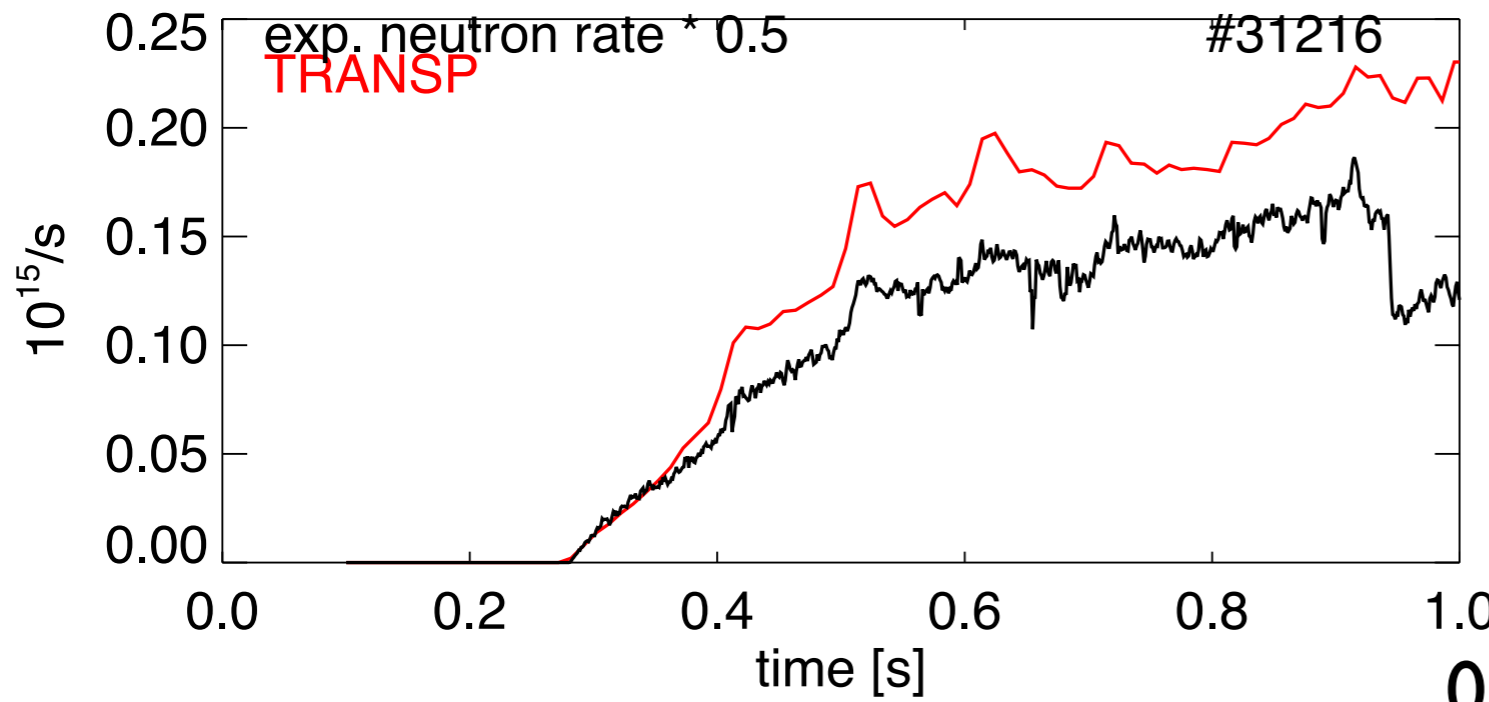
trapped/
passing boundary





ASDEX Upgrade

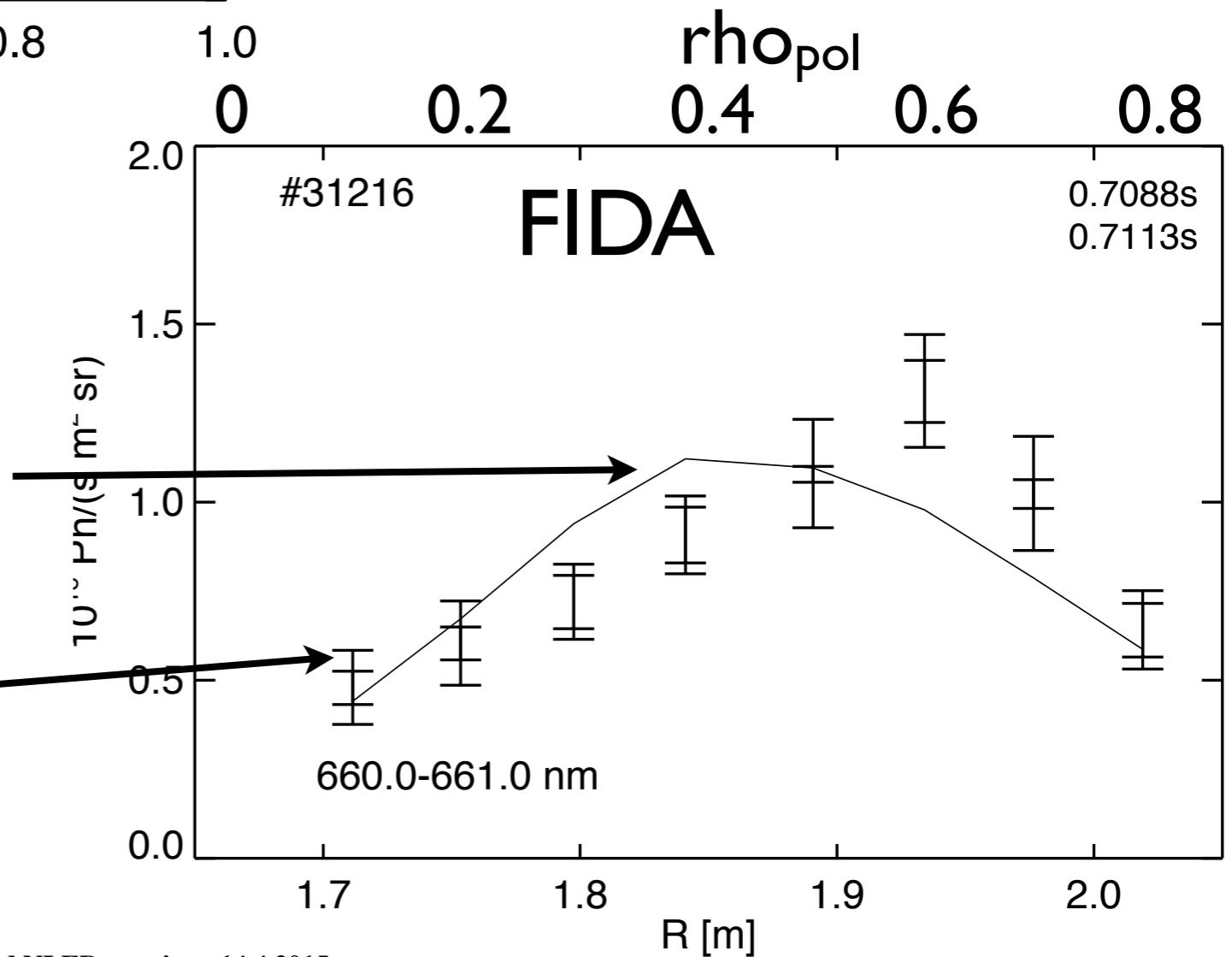
experimental EP transport



[B Geiger]

forward simulation

measurements

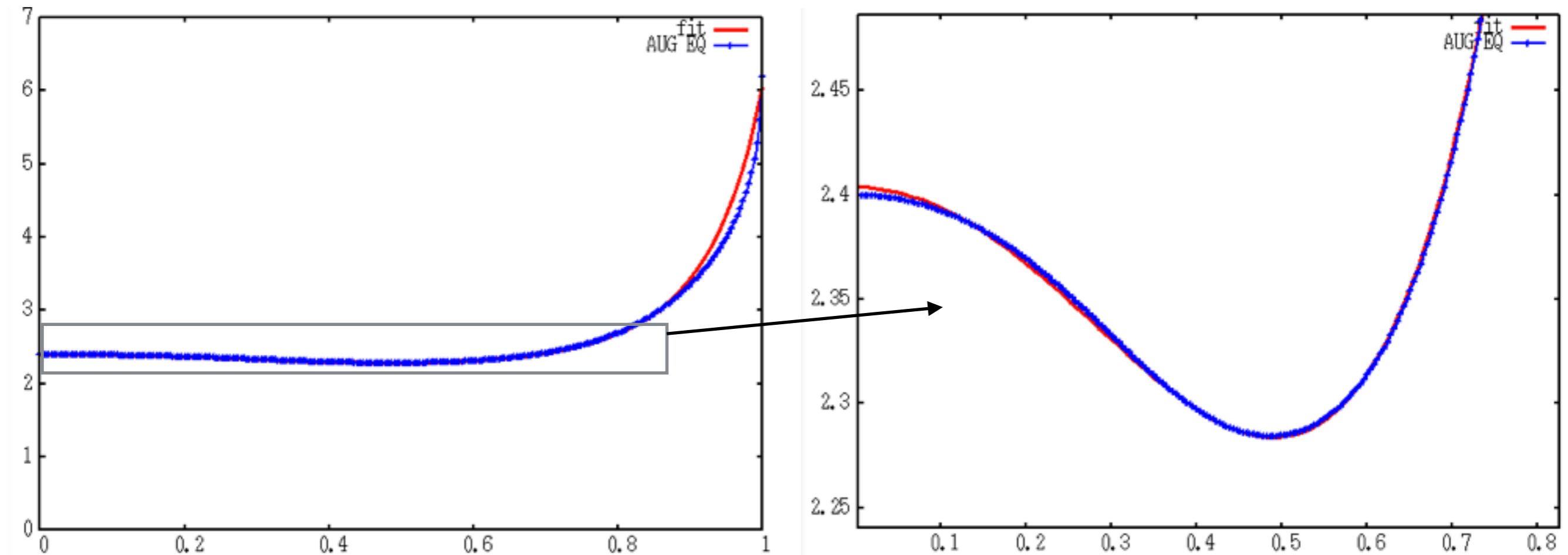


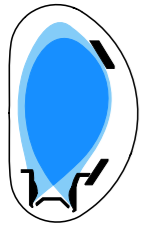
x: sqrt (normalised poloidal flux)

$$q(x) = a + b x^2 + c x^4 + d x^6$$

a = 2.40367; b = -0.994915; c = 2.06782; d = 2.61738

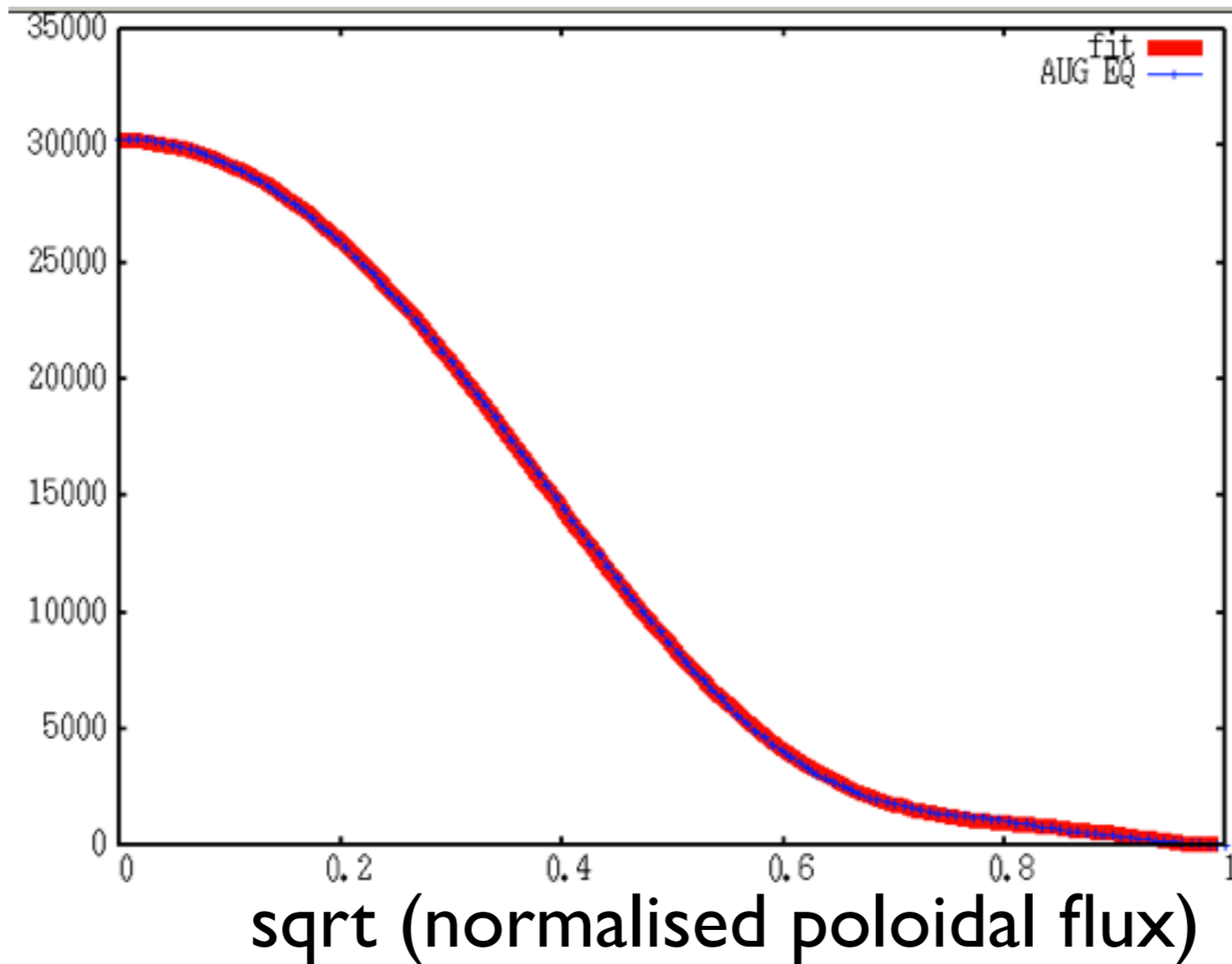
simplify?





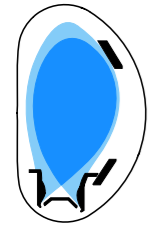
total equilibrium pressure

[kPa]



$$f(x) = a + b x^2 + c x^4 + d x^5 + g x^6 + h x^7$$

- a = 30195.9
- b = -102303
- c = -381943
- d = 1.65216e+06
- g = -1.85843e+06
- h = 660312



$$R_{\text{geo}} = 1.620\text{m}$$

$$B_{\text{geo}} = 2.248\text{T}$$

$$R_{\text{mag}} = 1.666\text{m}$$

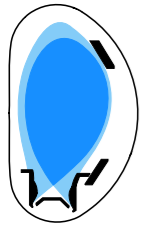
$$B_{\text{mag}} = 2.208$$

$$a [\text{m}] = 0.482\text{m}$$

$$\varepsilon = a/R_{\text{geo}} = 0.297$$

$$\beta_{\text{tot,axis}} = 1.3\%$$

- 1.step: take mid-radius values for background $T_{i,e}$; on axis flat density
- 2.step: use density profile
- 3.step: use profile for $T_i = T_e$
- 4.step: use different profiles for T_i, T_e



use parameters at mid radius:

$$n_e = n_D = 1.265 \cdot 10^{19} \text{ m}^{-3}$$

$$T_e = T_i = 1.6 \text{ keV}$$

$$f_{A0} = 909.209 \text{ kHz}$$

$$v_{a0} = 9.54 \cdot 10^6 \text{ m/s}$$

dimensionless parameters: $v_{th}^2 = 2T/m$

$$a/\rho_D = 115$$

$$a/\rho_{fast D} = 15 \text{ [93 keV]}$$

$$\beta_{kin,back,on-axis} = 0.3\%$$

$$v_{th,e}/v_{a0} = 2.49$$

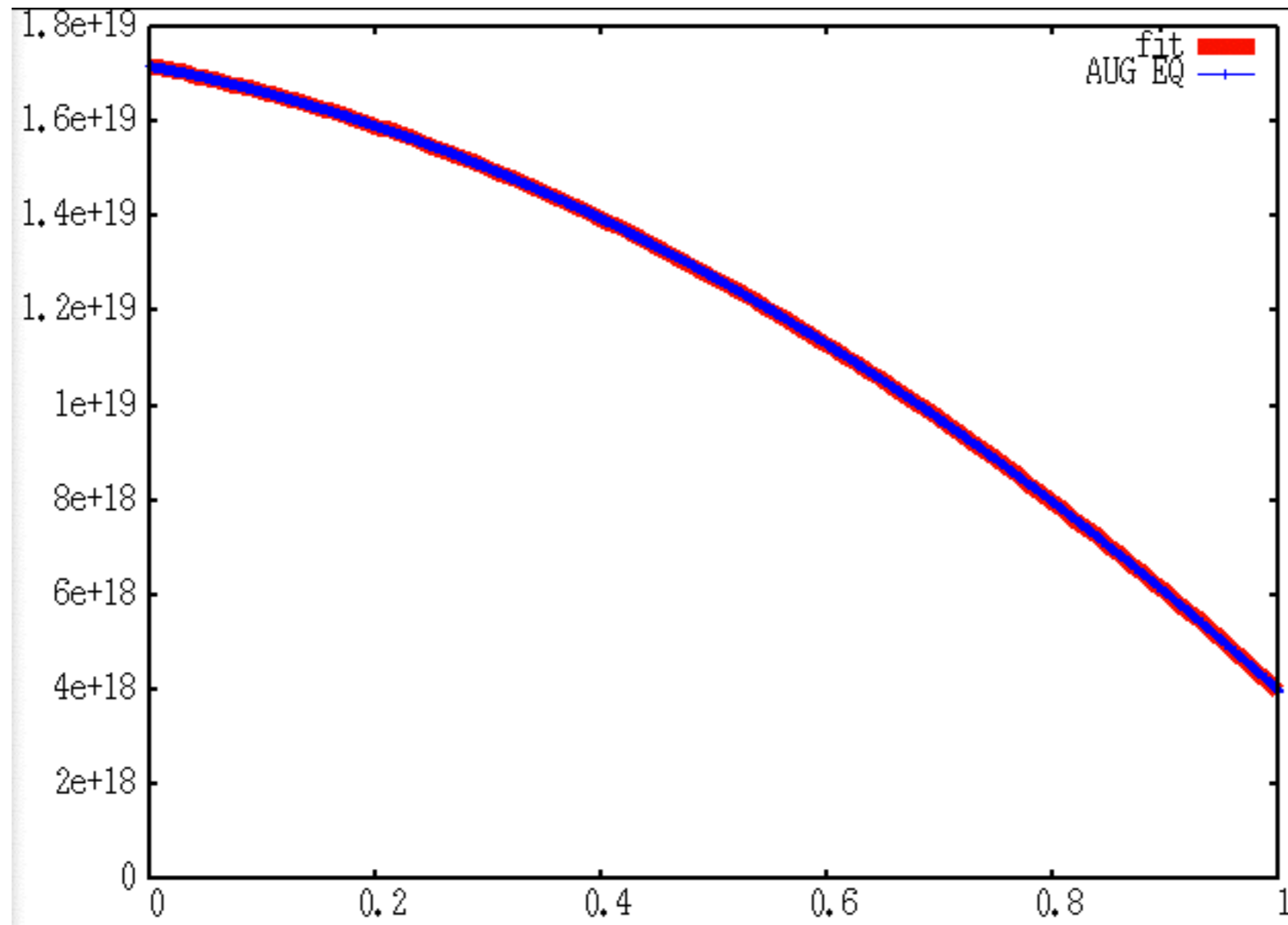
$$v_{th,D}/v_{a0} = 0.04$$

$$v_{th,fast D}/v_{a0} = 0.312 \text{ [93keV]}$$

x: sqrt (normalised poloidal flux)

electron density [m^{-3}]: $n_e(x) = a + b x + c x^2 + d x^4$

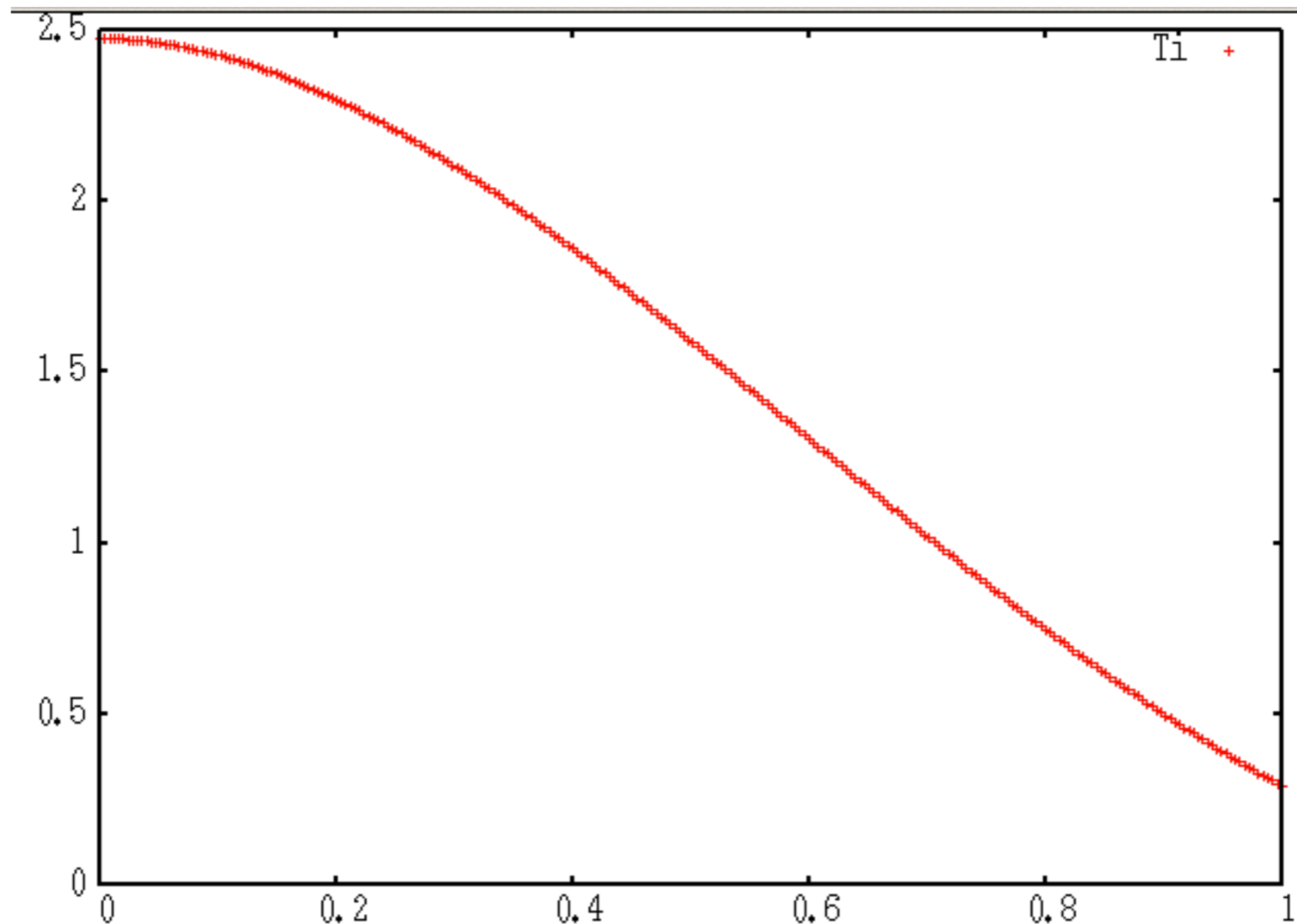
$a = 1.71587\text{e}+19$ $b = -4.52225\text{e}+18$; $c = -8.85926\text{e}+18$ $d = 1.98778\text{e}+17$



x: sqrt (normalised poloidal flux)

deuterium temperature[eV]: $T_i(x) = a + b x^2 + c x^3 + d x^4$

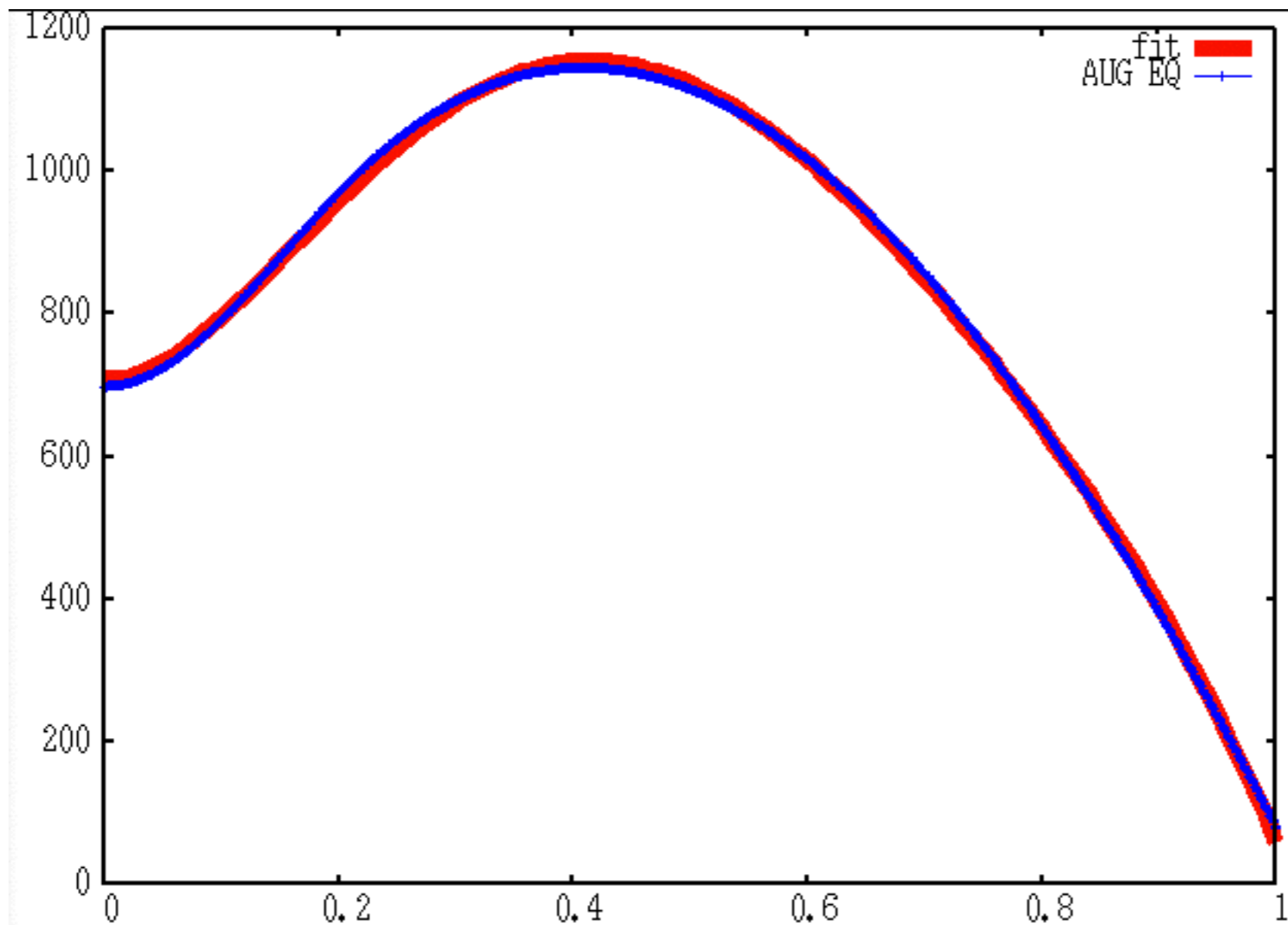
$a = 2474.1$; $b = -5128.7$; $c = 3417.9$; $d = -473.3$

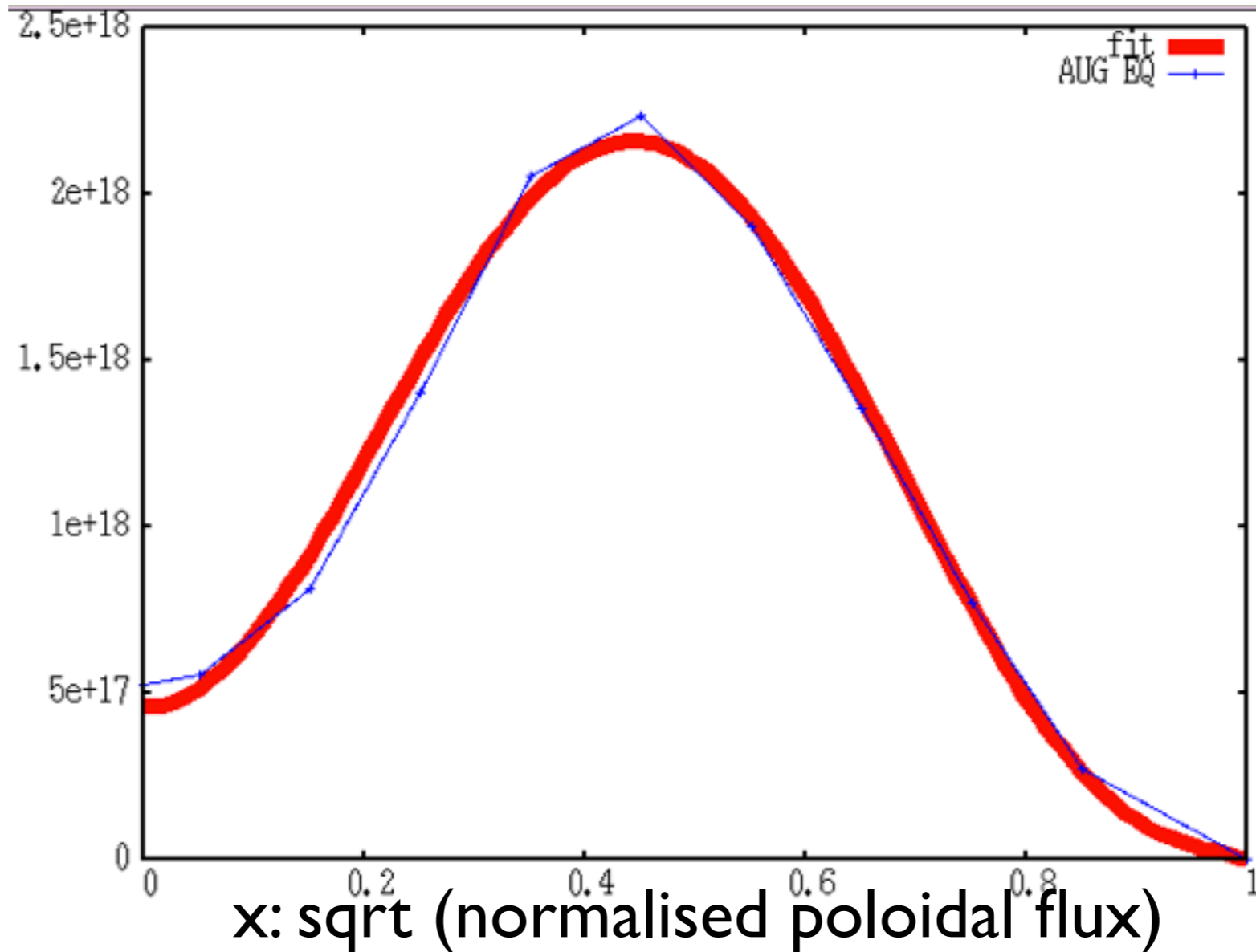
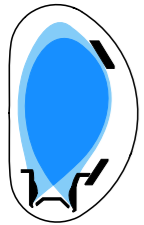


x: sqrt (normalised poloidal flux)

electron temperature [eV]: $T_e(x) = a + b x^2 + c x^3 + d x^4 + e x^5$

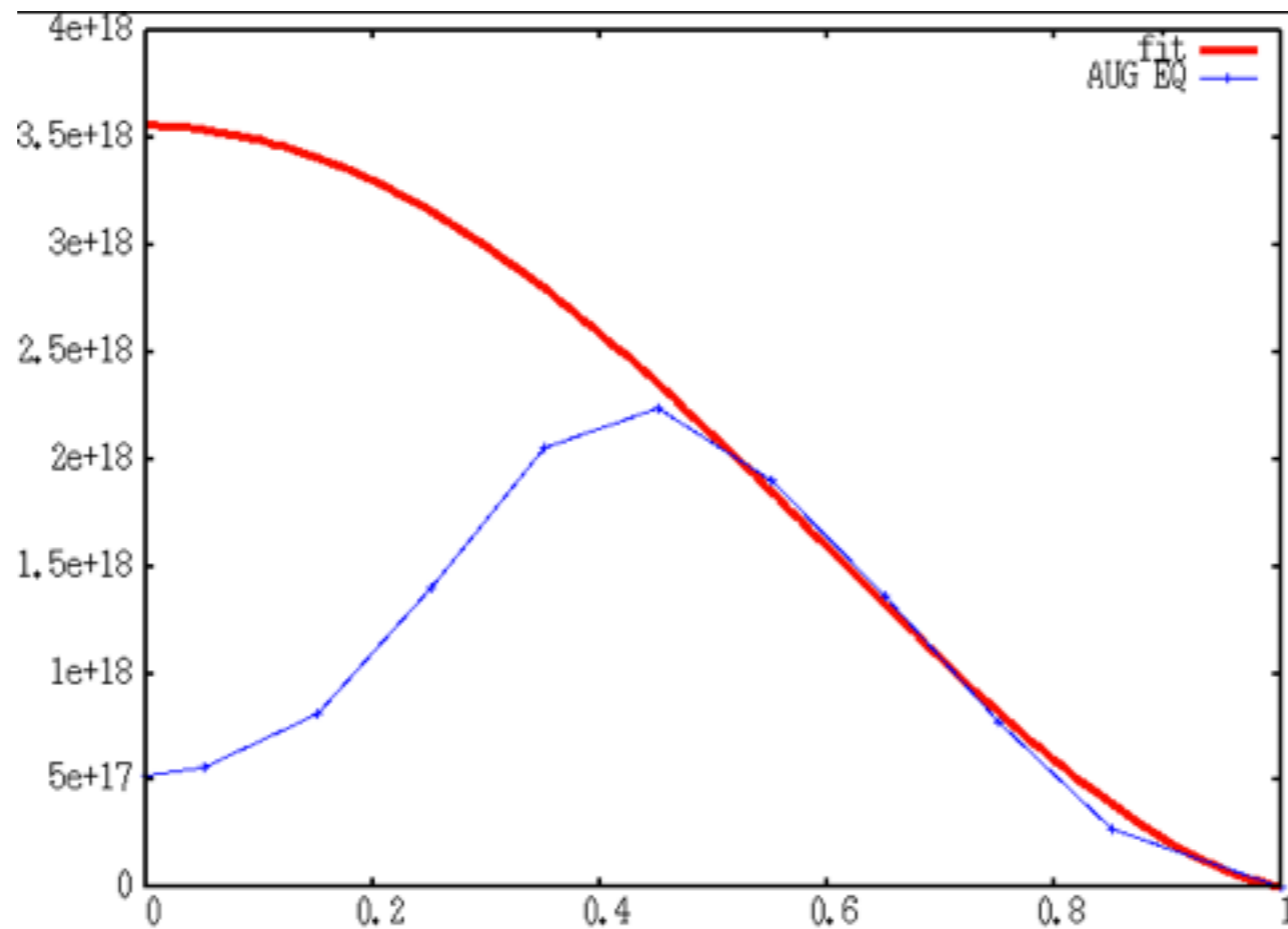
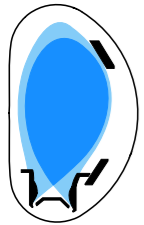
$a = 707.419; b = 11909.8; c = -34439.8; d = 33868.6; e = -11986.6$





$$nf(x) = a + b x^2 + c x^6 + e x^4 + g x^5$$

$a = 4.58182e+17$; $b = 2.30149e+19$;
 $c = -7.56176e+19$; $e = -1.4614e+20$; $g = 1.98287e+20$



$$n_{fD}(x) = a + b x^2 + c x^4$$

$$a = 3.552e+18;$$

$$b = -6.52986e+18;$$

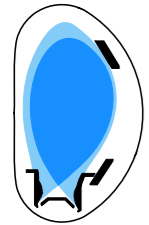
$$c = 2.97864e+18$$

suggestion for starting point (TAE, RSAE, BAE):

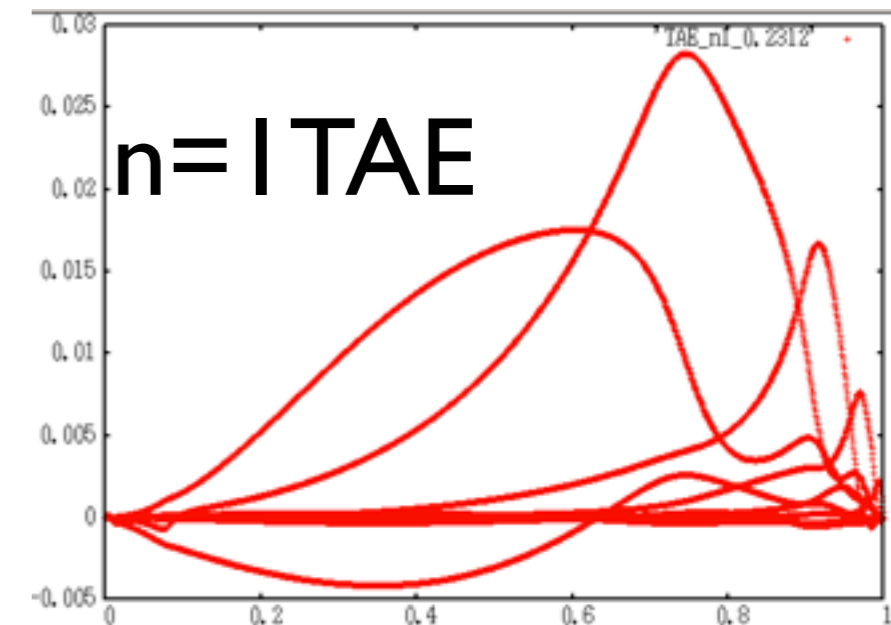
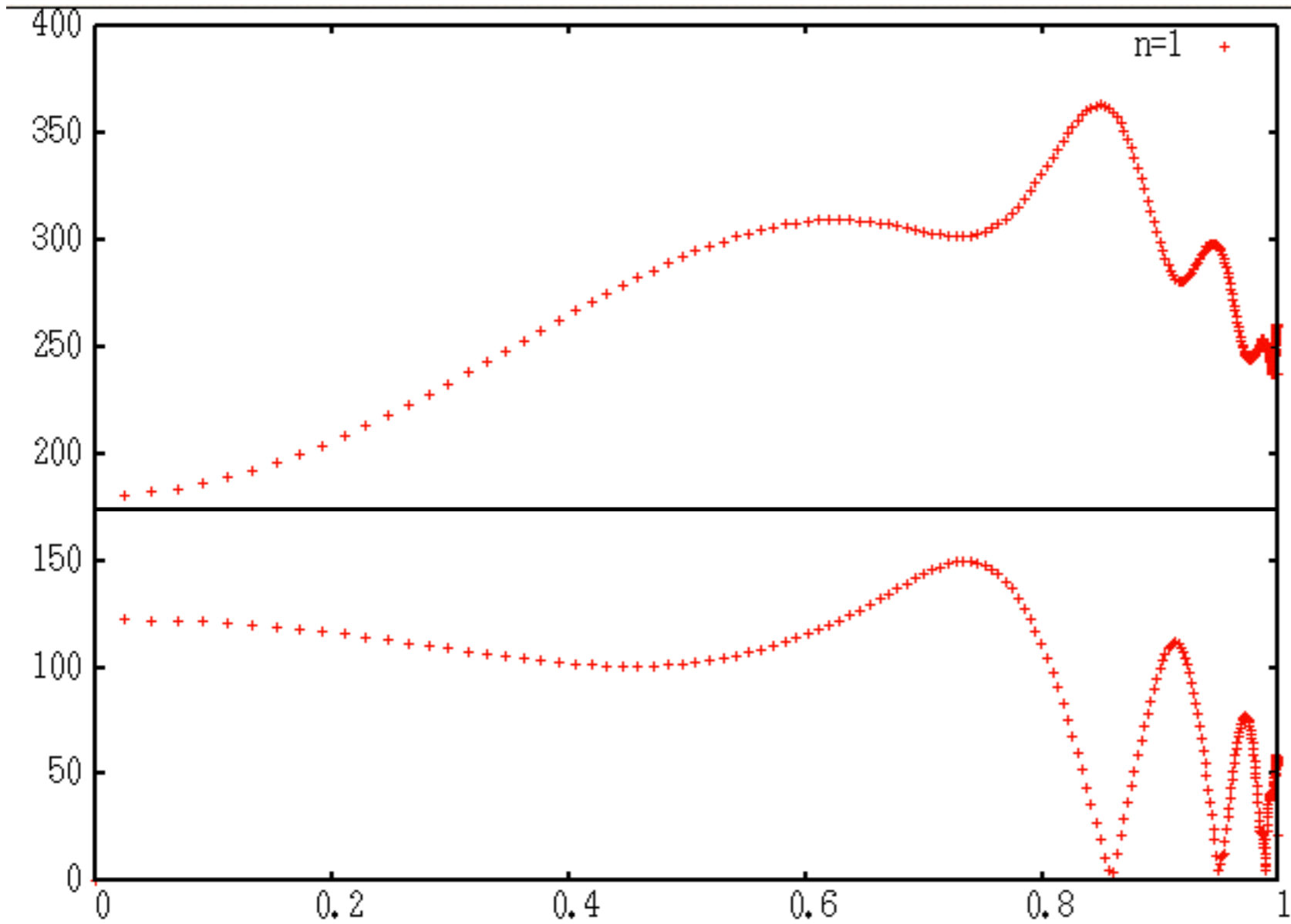
change (Maxwellian) temperature; 'effective' $T_{fast,D} \sim 30\text{keV}$ ($T_f/T_D = 18.75$)

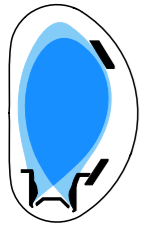
representation of anisotropy (EGAM) to be determined together with Claudio
 is Claudio's representation [PPCF, 54, (2012) 105017] ok?

$$f_{eq} = \frac{\mathcal{N}(w/T_w)^{\alpha_w}}{\sqrt{2\pi}w^{3/2}} \exp\left[-\left(\frac{\mathcal{P}_\phi - \mathcal{P}_{\phi 0}}{\Delta P_\phi}\right)^2\right] \exp\left\{-\frac{w}{T_w}\left[1 + \left(\frac{\lambda - \lambda_0}{\Delta\lambda}\right)^2\right]\right\}$$

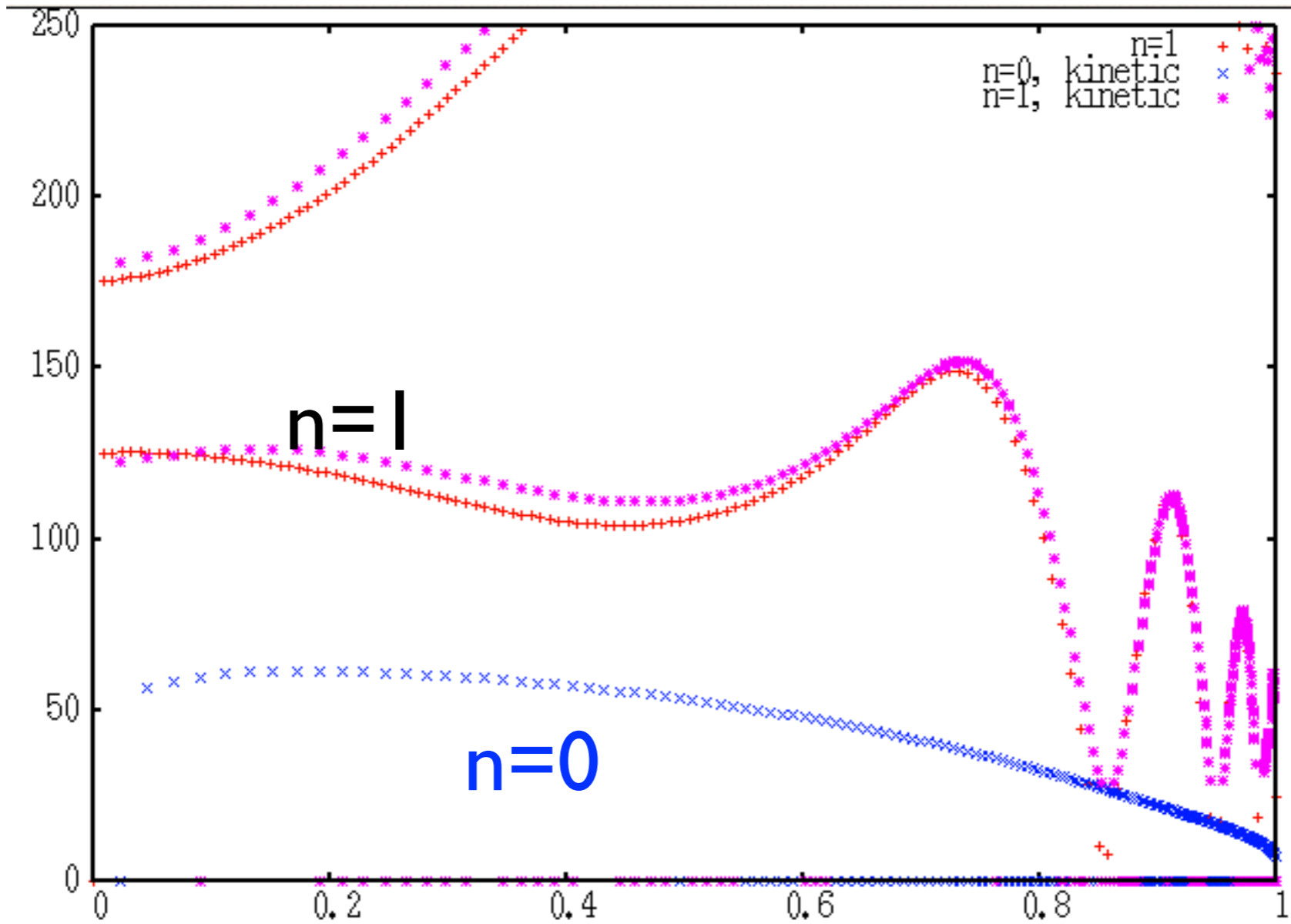


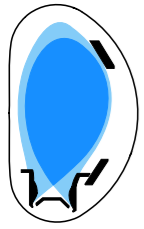
ideal $n=1$ SAW spectrum





kinetic spectra





- AUG (and DIII-D?) seem to be close to regimes with strongly nonlinear EP dynamics (like spherical Tokamaks or JT-60U)
- TAE/EPM bursts instead of several marginally stable modes - where is the transition?
- experimentally: reduced beam voltage and NBI power scans to be performed in 2015
- are proposed parameters possible for NLED codes? simplifications needed? different representation?
- EP parametric distribution function to be determined (next weeks)
- scale to burning plasmas: change ρ^* - mode numbers