





## Onset of strongly non-linear energetic particle dynamics at ASDEX Upgrade

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### why co-NBI drive EGAMs at AUG and not at DIII-D?





#### central Te for several EGAM discharges





### #32326(NB only, EGAMs, AEs) vs #32328 (NB+ECRH, no EGAMs, AEs) : profiles







onset conditions for EGAMs and BAEs [Lauber IAEA TM 2013]

#### q≥2,Te,Ti<I.8keV







longer, more stable phases with EGAMs in new experiments





 $\beta_{EP}/\beta_{th}$ >I;Tf ~90 Ti DEMO-like conditions for these two parameters



- the accumulation of W in the core decreases the background Te, Ti and thus reduces the ion Landau damping of AEs and GAMs
- hollow Te profiles form
- large values of  $R_0\nabla\beta_\alpha/\beta_{\text{back}}\,$  can arise
- central ECRH can counteract this accumulation; the increased Landau damping (exponential dependence on Ti) brings the system below excitation threshold (#32328) → threshold typically Ti ≤1.8keV for q>2
- Emax/Ti,th ~90 is comparable to burning plasma parameters (ITER/DEMO: 3,5MeV/30keV)
- opens possibility to study experimentally the interaction between Alfvénic modes, EGAMs i.e. zonal structures and background turbulence: due to the EGAM excitation a direct channel from EPs to n=0 modes can be investigated



strongly non-linear EP dynamics at AUG (in contrast to DIII-D) for sub-Alfvénic neutral beam injection raises the following questions:

- experimental conditions?
- why can system arrive at state well above marginal stability (critical gradient models would not allow for that state)
- study non-linear evolution of various ES and EM modes
- study interaction between modes
- study interaction with turbulence
- study scenarios that match projected DEMO parameters in  $\beta_{fast}/\beta_{th}$  and  $T_{fast}/T_{th}$

obtain confidence in models and codes towards understanding the self-organisation of a burning plasma





• small density feed-forward: more stable transition through  $q=2/q_{edge}=4$  region • optimised beam blips for measuring Ti









- q=2 sawtooth crashes like in 2001-05 JET/JT60U discharges (current holes) is this how advanced scenarios JT-60SA will look like? [A Bierwage, 2016/17]
- EGAMs persistent during these crashes
- almost no q-dependence of f<sub>EGAM</sub> is observed
- surprisingly no variation of EGAM onset frequency constant Ti!

### successful change of plasma position and shape after t>0.8s: mainly different NB deposition position











-3.0 different dynamics for:

- TAEs
- EGAMs
- TAE/EGAM coupling
- q=2 crashes



#### scenario developed further (III)





2nd order EGAM excitation! signature of density perturbation?

DIII-D,Fu [2011]: 2nd order outboard midplane density perturbation is comparable to first-order perturbation

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Time: 0.750 to 1.100 frg: 0.0 to 150.0 nfft: 4096 npad: 0 netp: 512 nrme: 1000 neor: 200



**ASDEX Upgrade** 



signals with rapid chirping (amplitudes, frequency) may cause 'spurious' bicoherence Phase randomized bicoherence probability density function calculation





signals with rapid chirping (amplitudes, frequency) may cause 'spurious' bicoherence **phase randomized bicoherence probability density function** calculation

- High bicoherence around at (55, 55) kHz indicates strongly nonlinear EGAMs. (see spectrogram at ~110 kHz)
- Without filtering interaction with TAEs is not clear
- Filtering shows high, significant bicoherence around (155, 55) kHz
- Indicates the nonlinear interaction between EGAMs and TAEs







### modeling efforts

- linear drive
- radial propagation
- non-linear interaction: wave-wave
- non-linear interaction: wave-particle





- linear local and global GK solver: QN, GKM and GKE, non perturbative mode structures
- continuous spectrum: null-space of 2nd order operator
- in addition to fully numerical solver, new analytical finite orbit width version: Taylor expansion to 2nd order for EPs and thermal ions:

k<sub>r</sub> as parameter or 4th order terms; recovers [Zonca 1998] in appropriate limit

- antenna, nyquist contour and vector iteration solver
- LIGKA\_lib as library for wrapper or other codes (HAGIS)

#### LIGKA model equations













- analytical models point out importance of radial GAM/EGAM propagation [Zonca 2008, Qiu, 2009,Smolyakov 2009, Sasaki, Miki&Idomura 2015, Palermo 2017, etc..]
- nonlinear simulations GYSELA, GTS [Zarzoso, Schneller] and nl-analytical models [Sasaki] emphasise the role of radial GAM/EGAM propagation for turbulence spreading



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experimentally: no (or small) radial propagation but rather radial shrinking of the modes was found





[Horvath, NF 2016]

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### finite orbit width (FOW) dispersion relation for LIGKA



$$\omega^{2} \left(1 - \frac{\omega_{*p}}{\omega}\right) - k_{\parallel}^{2} \omega_{A}^{2} R_{0}^{2} = 2 \frac{v_{thi}^{2}}{R_{0}^{2}} \left(-\left[H(x_{m-1}) + H(x_{m+1})\right] + \tau \left[\frac{N^{m}(x_{m-1})N^{m-1}(x_{m-1})}{D^{m-1}(x_{m-1})} + \frac{N^{m}(x_{m+1})N^{m+1}(x_{m+1})}{D^{m+1}(x_{m+1})}\right]\right)$$

no FOW, circulating particle approximation

[Zonca 1996,2009 Lauber 2009]



[Zonca 1998, Lauber to be published]

- generalised expression (wrt.  $\omega^*$ ) for FOW effects, implemented in LIGKA
- equivalent to EGAM FOW equations: Qiu [2009], Miki & Idomura [2015]
- fast analytical model for FOW effects: solve equations both locally (scan k<sub>r</sub>) and globally
- opens possibility for systematic EPM threshold checks
- extension to non-Maxwellian distribution functions on the way



benchmarking FOW/FLR version with [Zonca e tal 1998,2008]



$$\begin{split} \Lambda_0^2 &\simeq 1 - \frac{q^2}{\Omega^2} \left( \frac{7}{4} + \frac{T_e}{T_i} \right) - \frac{q^2}{\Omega^4} \left( \frac{23}{8} + \frac{2T_e}{T_i} + \frac{T_e^2}{2T_i^2} \right) \quad \text{[ZLR,ZOW]} \\ &+ i\pi^{1/2} q^2 \Omega^3 e^{-\Omega^2} \left[ 1 + (1 + 2T_e/T_i)/\Omega^2 \right]. \end{split}$$

$$\begin{split} \Lambda_0^2 &\to \Lambda_0^2 - (k_\zeta^2 \rho_i^2/2) \left( 3/4 + (q^2/\Omega) S_0(\Omega) \right) \qquad \text{[FLR,FOW]} \\ (q^2(7/4 + T_e/T_i) \gg 1) \\ \frac{3}{4} &+ \frac{q^2}{\Omega} S_0(\Omega) \simeq \frac{3}{4} - \frac{q^2}{\Omega^2} \left( \frac{13}{4} + 3\frac{T_e}{T_i} + \frac{T_e^2}{T_i^2} \right) \\ &+ \frac{q^4}{\Omega^4} \left( \frac{747}{32} + \frac{481}{32}\frac{T_e}{T_i} + \frac{35}{8}\frac{T_e^2}{T_i^2} + \frac{1}{2}\frac{T_e^3}{T_i^3} \right) \\ &- i\pi^{1/2}q^4e^{-\Omega^2/4} \left[ \Omega^5/256 + (1 + T_e/T_i)\Omega^3/32 \right] \end{split}$$

**Ibb** 

### KGAM benchmark: analytical theory vs LIGKA



q=3.25 τ=0.05

$$\begin{split} \frac{3}{4} &+ \frac{q^2}{\Omega} S_0\left(\Omega\right) \approx \frac{3}{4} \frac{q^2}{\Omega^2} \left(\frac{13}{4} + 3\frac{T_e}{T_i} + \frac{T_e^2}{T_i^2}\right) \\ &+ \frac{q^4}{\Omega^4} \left(\frac{747}{32} + \frac{481}{32}\frac{T_e}{T_i} + \frac{35}{8}\frac{T_e^2}{T_i^2} + \frac{1}{2}\frac{T_e^3}{T_i^3}\right) \\ &- i\pi^{1/2}q^4e^{-\Omega^2/4} \left[\Omega^5/256 + (1 + T_e/T_i)\Omega^3/32\right] \end{split}$$



lbb

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radius





 $vg = \partial \omega / \partial kr$  can be determined from this:









for short wavelength: frequency upshift causes reduced damping

### for large Te/Ti: also inward propagation possible:





#### next step: apply quantitatively to EGAM physics (kr is set by EP drift orbit width)

### stable, global kGAM solutions emphasize importance of GAM continuum[LIGKA], exp. f<sub>EGAM</sub> close to f<sub>GAM</sub>





# Effects of velocity anisotropy on the excitation of EGAMs;



Work in progress; To be submitted to Nuclear Fusion: Conf. series , I Chavdarovski, M Schneller, Z Qiu, A Biancalani;

- The local dispersion relation of energetic -particle-induced Geodesic acoustic mode (EGAMs) is derived for both circulating and trapped particle beam with single pitch angle slowing down and Maxwellian distribution.
- Solutions of the local dispersion relation for each case give the spectrum, the growth rate and the threshold of excitation as function of the slowing down critical energy (for the trapped) and the pitch angle (for circulating).
- Sample result: dispersion relation for trapped EP with slowing down distribution, as function of the bounce frequency and Λ. The logarithmic term gives the drive.

$$0 = -1 + \frac{\omega_G^2}{\omega^2} + \frac{\pi}{8} N_b \frac{1}{\epsilon \Lambda_0 B_0} \left[ \frac{3 - 2\Lambda_0 B_0 (1 - \epsilon)}{2(1 - \Lambda_0 B_0 (1 - \epsilon)^{3/2})} \log\left(1 - \frac{\omega_b^2}{\omega^2}\right) + \frac{1}{1 - \Lambda_0 B_0 (1 - \epsilon)^{1/2}} \cdot \frac{\omega_b^2 / \omega^2}{1 - \omega_b^2 / \omega^2} \right]$$

### to be implemented together with anisotropic shifted Maxwellian into LIGKA





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### Saturation of EGAMs due to waveparticle nonlinearity

Work by A Biancalani, I Chavdarovski and Z Qiu. Currently on IPP pinboard.

- Only wave-particle nonlinearities are considered.
- The EGAM saturates mostly due to flattening of the EP profile in velocity space.
- Quadratic scaling of the saturated electric field with the linear growth rate is found, similarly to the beam-plasma instability:  $\delta \bar{E}_r = \alpha_2 \gamma_L^2$



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- The EP bounce frequency in the mode potential well is proportional to the linear growth rate :  $\omega_b = 2.7 \gamma_L$ . This factor is 3.2 for beam-plasma system.
- Near the saturation we observe a transition from adiabatic to non-adiabatic chirping, where the time derivative of the mode frequency is of the same order as the squared bounce frequency.
- Non-adiabatic chirping is due to non-perturbative response of the resonant particles, experiencing trapping and de-trapping in the mode.



### Non-linear generation of Zonal Flow and EGAMs second harmonic;

Published work Z Qiu , I Chavdarovski, A Biancalani and J Cao,

Physics of Plasmas 24, 072509 (2017);

- Both second harmonic and ZFZFs can be driven by EGAMs, with the finite orbit width (FOW) effects playing a dominant role in the nonlinear couplings, contrary to the thermal where toroidicity dominates
- The contribution of resonant EPs to the cross-section of the nonlinear couplings dominates that of the thermal plasma.
- The generated ZFZF has a radial scale length half of the pump EGAM and growth rate double the size.
- For GAMs perpendicular and parallel non-linearities cancel out , but present EP will dominate (Fu 2011). Second harmonic is generated with *I=1,2* resonances and radial scale half of the primary EGAM.

$$b_{S}\hat{\mathscr{E}}_{EGAM}(\omega_{S})\hat{\Phi}_{S} = -\frac{ik_{r}T_{i}}{n_{0}m\Omega}\left\langle\frac{\partial F_{0h}}{\partial E}\frac{3\omega\hat{\omega}_{d}^{2}}{\left(\omega^{2}-\omega_{lr}^{2}\right)\left(\omega_{S}^{2}-\omega_{lr}^{2}\right)}\right\rangle$$
$$\times\frac{\hat{\Phi}_{G}\hat{\Phi}_{G}}{r}$$
$$\hat{\mathscr{E}}_{EGAM}(\omega_{S}) \equiv -1 + \frac{\omega_{G}^{2}}{\omega_{S}^{2}} + \frac{T_{i}}{n_{0}m_{i}b_{S}}\left\langle\frac{\partial F_{0h}}{\partial E}\sum_{l=\pm 1,\pm 2}\frac{J_{l}^{2}(\hat{\Lambda}_{S})\omega_{S}}{\omega_{S}-l\omega_{lr}}\right\rangle$$





interesting and unique set of experimental AUG data:

- more stable EGAM scenarios developed
- non-linear coupling of EGAM and TAEs found
- second order EGAMs measured
- shrinking of mode structure found

modeling progress:

several linear and non-linear key physics items are being addressed

allows for physics understanding and development of fast reduced models while moving towards complete non-linear EM simulations





#### extra material







### Euterpe: [M. Cole] elliptic eq

ion diamagnetic propagation exp: electrondiamagnetic propagation!



LIGKA [Lauber] AUG eq ω=0.238 ωA, γ/ω=1.5%







experimental difference: longer (30ms) on-axis pre-heating phase leads to very flat q-profile

> note: BAE/EPM and GAM onset frequency almost constant BAE/EPM onset above BAE continuum accumulation point (toroidal rotation ~10kHz)









### include small but finite kr (background only): BAEs and KBAEs

finite Larmor radius and orbit width effects



#### determine $\partial \omega / \partial kr$ and match for discrete KBAE solutions

at BAE accumulation point ( $\rho$ =0.41): ~0 \*left from q=1 surface ( $\rho$ =0.355): <0 + right from q=1 surface ( $\rho$ =0.459): >0 X



EP excitation and lowfrequency modes to be studied



[C. di Troia 2015]



#### Equilibrium Distribution Function

- Constants of Motion at leading order if  $B = \nabla \psi \times \nabla \phi + F \nabla \phi$ 
  - toroidal canonical momentum:  $\mathcal{P}_{\phi} = \psi + (F/\omega_c)v_{\parallel}$
  - energy:  $w = v^2/2 = v_{\parallel}^2/2 + \mu |B|$
  - $\lambda = \mu/w = (1 v_{\parallel}^2/v^2)/|B|$ , being  $\mu = v_{\perp}^2/(2|B|)$
- Parameters
  - $\triangleright \mathcal{N}$

is the overall regularization factor are the parameters of the *Gamma* distribution in energy (at  $\triangleright \mathcal{N} = 1.2 \times 10^{17} \, \mathrm{m}^{-3}$ ▷  $T_w = 31.13$  keV,  $\land \alpha_w = 1.0,$ ▷  $\mathcal{P}_{\phi 0} = 0.035$  Wb,  $\triangleright \Delta_{P_{\phi}} = 0.02 \text{ Wb}$ ▷  $\lambda_0 = 0.08 \text{ T}^{-1}$  $\triangleright \Delta_{\lambda} = 0.12 \text{ T}^{-1}$ ▷  $w_b = 93$  keV,  $\triangleright w_c = 15 \text{ keV}$ 

- $\triangleright$   $T_w, \alpha_w$ fixed  $\lambda$ )
- $\triangleright \mathcal{P}_{\phi 0}, \Delta_{P_{\phi}}$
- $\triangleright \lambda_0, \Delta_\lambda$

 $\triangleright w_b, w_c$ 

are the parameters of the Normal distribution in  $\mathcal{P}_{\phi}$ are the parameters of the *Normal* distribution in  $\lambda$  (at fixed energy) are the parameters of the *SlowingDown* distribution in energy

$$f_{\text{Ref1}}(\mathcal{P}_{\phi}, w, \lambda) = \mathcal{N}\frac{(1+\lambda/\lambda_{0})(w/T_{w})^{\alpha_{w}}}{\sqrt{2\pi}} \frac{\mathrm{H}(w_{b}-w)}{w^{3/2}+w_{c}^{3/2}} \times \exp\left[-\frac{(\mathcal{P}_{\phi}-\mathcal{P}_{\phi0})^{2}}{\Delta_{P_{\phi}}^{2}}\right] \exp\left[-\frac{w}{T_{w}}\left(\frac{\lambda-\lambda_{0}}{\Delta_{\lambda}}\right)^{2}\right]$$

### first: shifted Maxwellian for EGAM benchmark

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### ideal n=1 SAW spectrum









 Bicoherence analysis is a widely applied method for identifying quadratic nonlinear interactions of stationary processes [1]

$$b^{2}(f_{1},f_{2}) = \frac{|\mathbf{A}[X(f_{1})X(f_{2})X^{*}(f_{1}+f_{2})]|^{2}}{\mathbf{A}[|X(f_{1})X(f_{2})|^{2}]\mathbf{A}[|X^{*}(f_{1}+f_{2})|^{2}]}, \text{ with } \mathbf{A}[Y] \coloneqq \lim_{N \to \infty} \frac{1}{N} \sum_{i=1}^{N} Y_{i},$$

- (b) Instationary (c) (a) Stationary (a) Stationary Signals exhibiting rapidly Not coupled Not coupled Phase-coupled process changing amplitudes or Im 2 Probability density frequencies (typical in the Confidence case of strongly driven level EPMs) may cause high b<sup>c</sup> (α≈0.9) bm bicoherence even without phase coupling 0.2 0.8 Re Re 0.0 0.4 0.6 1.0
- Phase randomized bicoherence probability density function calculation for each (f1,f2) point, which will describe a random process without any phase coupling.
- We can set an α confidence level, which will define a b<sub>c</sub> critical bicoherence value (at each (f<sub>1</sub>,f<sub>2</sub>) point)

ne 
$$\alpha = \int_{0}^{b^{c}} \rho(b) db$$

**Bicoherence** 

 Bicoherence values higher than b<sub>c</sub> are from non-phase coupled process with 1-α probability. Confidence level can be used as a filtering parameter, by only plotting bicoherence values higher than b<sub>c</sub>, thus eliminating probable false positives PP

### include small but finite kr (background only): finite Larmor radius and orbit width effects





#### discharges with the same total heating power





Time: 0.200 to 1.500 fre: 0.0 to 500.0 efft: 4046 read: 0 addr 512 renue: 1000 n



### 32326(NB) vs 32328 (NB+ECRH) : profiles





### PP

Comparison of:

discharge with and without EGAM/TAE/ W accumulation

very different turbulent spectrum magnetic coils HF side

although R/L<sub>Ti</sub> is quite similar

is this due to modes?



G Shot: 32327 : MHD : C09-16 npts: 259990

Time: 0.200 to 1.500 frq: 0.0 to 500.0 nfft: 4096 npod: 0 natp: 512 nmme: 1000 neur: 200





#	EGAM/BAE/	NBI	angle	behav	later heating		1	В
<u>27923</u>	<mark>y/y</mark> /y/n	2:0.35-0.5;3:0.38-0.59;80.59-0.63;5:0.63-0.76;7:0.76	6,65					
28880	n/y/y/n	2:0.35-0.5;3:0.5-0.6;7:0.6	6,65					2,4
28881	y/y/y/n	2:0.35-0.5;3:0.5-0.6;7:0.6	6.65					2,4
28883	n/v/n/n	<b>2:0.35-0.5:3:0.5-0.6:7:0.6</b>	6.65				ASDEX	Upgrade
28884	y/y/y/n	3:0.5-0.6;7:0.6	6,65					2,4
28885	y/y/y/n	2:0.35-0.5;3:0.5-0.6;7:0.6	6,65					2,4
30383	y/y/y/n	7: 0.26-0.75	6,65	Hmode		FILD FHA FIPM 09	I	2,6
30945	n/y/n/n	2:0.28-0.376;6:0.382-0.697	6.65	dis@4s			I	2,2
30946	y/y/n/y	2:0.28-0.445;6:0.451-0.928	6,65	Lmode	no heating!	later TAE???	I	2,2
30947	y/n/n/y	2:0.28-0.478:6:0.482-0.928	6.65	dis@4s	H mode	EGAM @1s 100kHz		2.2
30948	n/y/y/n	2:0.28-0.491;3:0.497-0.789	6,65	dis@1.2s	Q6@0.789		1	2,2
30949	y/y/n/n	2:0.35-0.5;3:0.38-0.79;6:0.79;7:1.0;8:1.2	6,65	dis@1.5		late EGAMs	I	2,2
30950	y/y/y/n	<u>3:0.28-0.295;7:0.312-0.797</u>	6,65	dis@1.5	3:0.8-0.92;6,8@0.9		I	2,2
30951	n/y/n/n	<u>3:0.28-0.295;5:0.312-0.552,8</u>	6,65	dis@1.7	8-0.84;3:-0.99		1	2,2
30952	y/y/y/n	<u>3:0.28-0.295;7:0.312-0.797</u>	6.65	dis@1.18	<u>Q6@0.8</u>		1	2,2
30953	y/y/n/n	<u>3:0.28-0.295;6:0.3 2-0.753</u>	6,65	dis@1.11	Q2@0.76++		1	2,2
31213	y/y/y/n	<u>3:0.28-0.295;7:0.296-1.033</u>	7.13	dis@1.7	O6@1.0		1	2.2
31214	y/y/y/n	<u>3:0.28-0.295;7:0.296-1.033</u>	6,05	dis@1.0			1	2,2
31215	y/y/y/n	<u>3:0.28-0.295;7:0.296-1.033</u>	6,65	dis@1.0			I	2,2
31216	y/y/y/n	3:0.28-0.295;7:0.296-3.045+blips	6,65	Lmode		g=2 and ga>4!	1	2.2
31233	y/y/y/n	<u>3:0.28-0.501;7:0.506-3.227</u>	7,13	Hmode	Q6@1.0			2,2
31234	y/n/y/n	3:0.28-0.310;7:0.318-0.813	7.13	dis@ 0.8			1	2.2
32326	y/n/y/y	7: 0.28 +blips	7.13	EGAMS, TAEs				2.2
32327	y/n/y/n	7: 0.28 +blips: 82kV	7.13	transition			I	2.2
32328	n/n/n/n	7: 0.28 +blips +0.5 ECRH	7.13	only turbulence				2.2
32329	n/n/n/n	7: 0.28 + blips+0.5 ECRH	7.13	only Alfvenic turb				2.2
32384	y/n/y/n	7:0.28 +blips 93kV	7.13	too high density			1	2.2
32386	y/n/n/n	7: 0.28 +blips: 65kV	7.13				I	2.2
32387	y/n/y/y	7+6: 0.28 +blips: 65kV	7.13				1	2.2
32388	y/y/y/y	7:0.28 +blips + higher density 93kV	7.13				1	2.2
33872	y/y/y/y	7:0.28 +blips + higher density 93kV	7.13		diff breakdown	no Te inversion		2.2
33873	y/y/y/y	7:0.28 +blips + higher density 93kV	7.13		diff breakdown	no Te inversion	<b> </b>	2.5
33874	y/y/y/y	7:0.28 +blips + higher density 93kV	7.13	dis@1.0	std brkdwn	no Te inversion	<b> </b>	2.0
33875	y/y/y/y	7:0.28 +blips + higher density 93kV	7.13	dis@1.0s	std brkdwn	no Te inversion	<b> </b>	2.2
34184	y/y/y/y	7:0.28 +blips + higher density 93kV	7.13	shape scan t>0.8		Te inversion		2.2
34185	y/y/y/y	7:0.28 +blips + higher density 93kV	7.13	shape scan_t>0.8		Te inversion	<b> </b>	2.2
34186	y/y/y/y	7:0.28 +blips + higher density 93kV	7.13	std		Te inversion	<b> </b>	2.2
34187	y/y/y/y	7:0.28 +blips + higher density 93kV	1 6.65	std		Te inversion	1	2.2

#### comparison to reference discharge (#32388/#34185)





MILLO meeting, 10.-22. September 2017



#### 32326(NB) vs 32328 (NB+ECRH)



